

# Quantum Mechanics and Atomic Physics

## Lecture 15:

### Angular momentum and Central Potentials

<http://www.physics.rutgers.edu/ugrad/361>

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# Last time

- S.E. in 3D and in spherical coordinates
- For a mass  $\mu$  moving in a central potential  $V(r)$

$$-\frac{\hbar^2}{2\mu} \left\{ \frac{1}{r^2} \frac{\partial}{\partial r} \left( r^2 \frac{\partial}{\partial r} \right) + \frac{1}{r^2} \frac{1}{\sin\theta} \frac{\partial}{\partial \theta} \left( \sin\theta \frac{\partial}{\partial \theta} \right) + \frac{1}{r^2} \frac{1}{\sin^2\theta} \left( \frac{\partial^2}{\partial \varphi^2} \right) \right\} \psi(r, \theta, \varphi) + V(r) \psi(r, \theta, \varphi) = E \psi(r, \theta, \varphi)$$

- Solutions separated into angular and radial parts

$$\psi(r, \theta, \varphi) = R(r) \underbrace{Y(\theta, \varphi)}_{\Theta(\theta) \Phi(\varphi)} = R(r) \Theta(\theta) \Phi(\varphi)$$

**Rewrite this in terms of angular momentum of the system**

$$-\frac{\hbar^2}{2\mu} \left\{ \frac{1}{R} \frac{\partial}{\partial r} \left( r^2 \frac{\partial R}{\partial r} \right) \right\} + (V - E) r^2 = -\frac{1}{2\mu} L_{op}^2 Y$$

$$L_{op}^2 = -\hbar^2 \left[ \frac{1}{\sin\theta} \frac{\partial}{\partial\theta} \left( \sin\theta \frac{\partial}{\partial\theta} \right) + \frac{1}{\sin^2\theta} \left( \frac{\partial^2}{\partial\phi^2} \right) \right]$$

# Angular Solutions

Solutions to  $\Phi(\varphi)$ :

$$\Phi(\varphi) = \sqrt{\frac{1}{2\pi}} e^{im_l\varphi}$$

Magnetic quantum number

$$m_l = 0, \pm 1, \pm 2, \dots$$

Solutions to  $\Theta(\theta)$ :

Associated Legendre functions

$$P_{l,m_l}(\cos\theta)$$

$$l = 0, 1, 2, 3, \dots$$

Orbital angular quantum number

Solutions to  $Y(\theta, \varphi)$ : Spherical harmonic functions

$$Y_{l,m_l}(\theta, \varphi)$$

$$l = 0, 1, 2, \dots$$

$$|m_l| \leq l$$

$$Y_{l,m_l}(\theta, \varphi) = \Theta_{l,m_l}(\theta) \cdot \Phi_{m_l}(\varphi)$$

# Space Quantization

- The magnetic quantum number  $m_\ell$  expresses the quantization of direction of **L**

$$L_z = m_\ell \hbar \quad L_z = L \cos \theta$$
$$\Rightarrow \cos \theta = \frac{L_z}{L} = \frac{m_\ell \hbar}{\sqrt{\ell(\ell+1)} \hbar} = \frac{m_\ell}{\sqrt{\ell(\ell+1)}}$$

Space quantization.

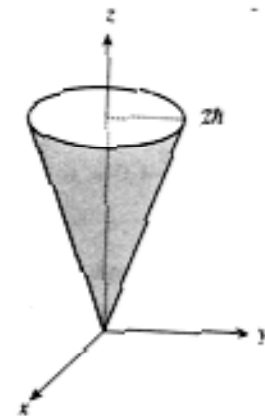


FIGURE 6.6 For  $\ell = 2$ ,  $m = 2$ ,  $L$  lies anywhere on the cone.

- So **L** can assume only certain angles, given above, with respect to the z-axis.
- **This is called space quantization.**

# Example

- For a particle with  $\ell=2$ , what are the possible angles that  $\mathbf{L}$  can make with the z-axis?

$$\ell=2, \quad L=\sqrt{6}\hbar \quad L_z=m\hbar$$

$$m_\ell = -2, -1, 0, 1, 2$$

$$\cos\theta = \frac{m_\ell}{\sqrt{\ell(\ell+1)}} = \frac{m_\ell}{\sqrt{6}}$$

- $m_\ell=2 \Rightarrow \cos\theta = \frac{2}{\sqrt{6}} \Rightarrow \theta = 35.3^\circ$
- $m_\ell=1 \Rightarrow \cos\theta = \frac{1}{\sqrt{6}} \Rightarrow \theta = 65.9^\circ$
- $m_\ell=0 \Rightarrow \cos\theta = 0 \Rightarrow \theta = 90^\circ$
- $m_\ell=-1 \Rightarrow \cos\theta = -\frac{1}{\sqrt{6}} \Rightarrow \theta = 114^\circ \text{ or } -66^\circ$
- $m_\ell=-2 \Rightarrow \cos\theta = -\frac{2}{\sqrt{6}} \Rightarrow \theta = 145^\circ \text{ or } -35^\circ$

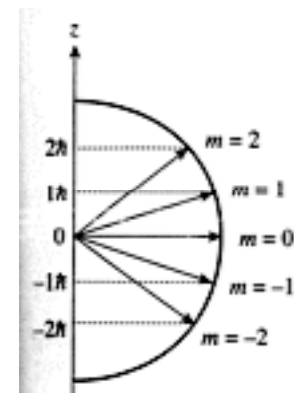


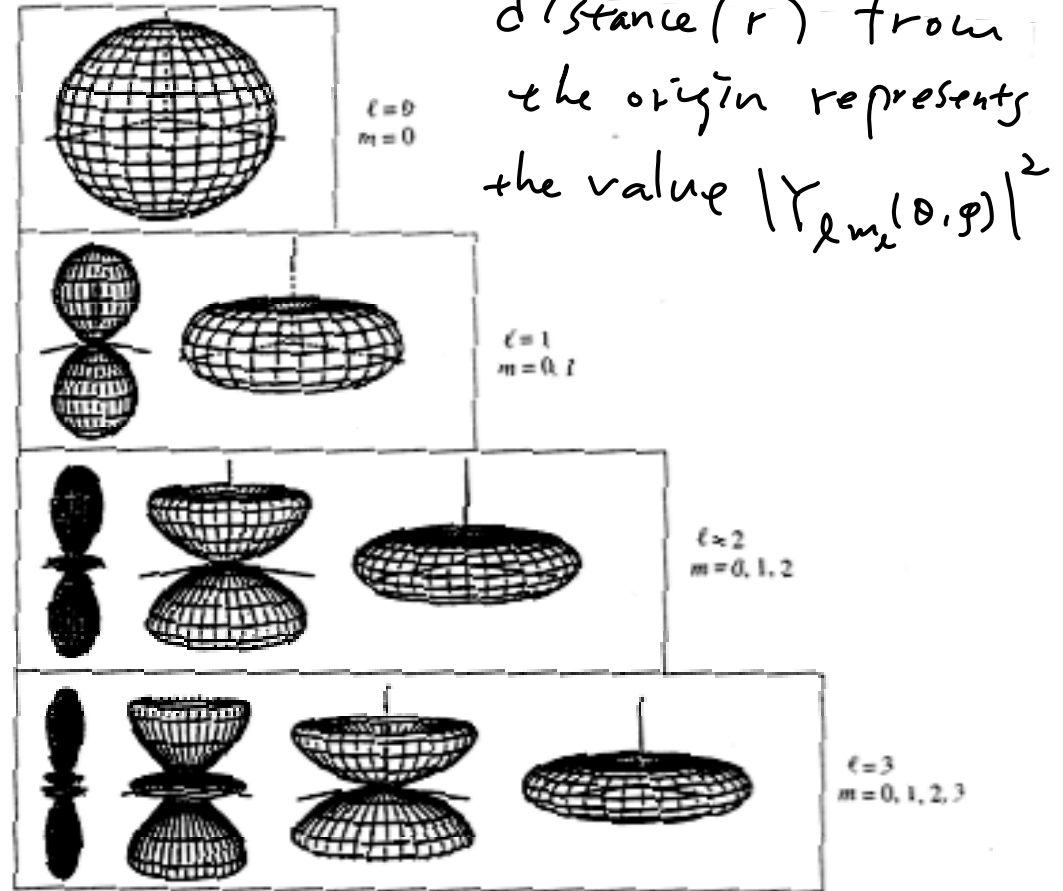
FIGURE 6.5 Possible orientations of  $\mathbf{L}$  for  $\ell = 2$ ,  $|\mathbf{L}| = \sqrt{6}\hbar$ .

# Plot of Spherical harmonics

- “3D” plots of:

$$r = \left| Y_{\ell m}(\theta, \varphi) \right|^2$$

- $\theta$  is measured from the +z axis
- Independent of  $\phi$ 
  - So rotationally symmetric around z-axis
- These manifest themselves as the probability distributions



**FIGURE 6.7** Polar diagrams of the absolute squares of spherical harmonics up to  $\ell = 3$ . Reproduced from Siegmund Brandt and Hans Dieter Dohmer, *The Picture Book of Quantum Mechanics*, John Wiley & Sons, Inc., New York, 1985, with the kind permission of the publisher.

# Summary: Quantization of $L$ , $L_z$ and space

$$L_{op}^2 Y_{\ell, m_\ell}(\theta, \varphi) = [\hbar^2 \cdot \ell(\ell+1)] Y_{\ell, m_\ell}(\theta, \varphi)$$

$$\Rightarrow L = \sqrt{\ell(\ell+1)} \hbar$$

$$(L_z)_{op} \Phi(\varphi) = m_\ell \hbar \Phi(\varphi)$$

$$\Rightarrow L_z = m_\ell \hbar$$

$$\cos\theta = \frac{L_z}{L} = \frac{m_\ell}{\sqrt{\ell(\ell+1)}}$$

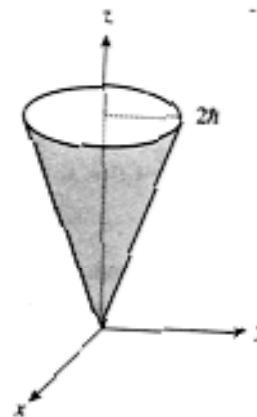


FIGURE 6.6 For  $\ell = 2$ ,  $m = 1$ ,  $\ell$  lies anywhere on the cone.



# Radial Solutions

- To find radial solutions:

$$-\frac{\hbar^2}{2\mu} \left\{ \frac{1}{R} \frac{\partial}{\partial r} \left( r^2 \frac{\partial R}{\partial r} \right) \right\} + \{V(r) - E\} r^2 = -\frac{1}{2\mu} \overbrace{\left[ \hbar^2 \ell(\ell+1) \right]}^{L_{op}^2 Y} Y$$

$$\Rightarrow -\frac{\hbar^2}{2\mu} \left\{ \frac{1}{R} \frac{\partial}{\partial r} \left( r^2 \frac{\partial R}{\partial r} \right) - \ell(\ell+1) \right\} + V(r) r^2 = E r^2$$

- This is called the *radial equation*

# Solving the radial equation

- Note:

$$\frac{1}{R} \frac{\partial}{\partial r} \left( r^2 \frac{\partial R}{\partial r} \right) = \frac{r}{R} \frac{\partial^2}{\partial r^2} (r R)$$

$$\Rightarrow -\frac{\hbar^2}{2\mu} \left\{ \frac{r}{R} \frac{\partial^2}{\partial r^2} (r R) - \ell(\ell+1) \right\} + V(r) r^2 = E r^2$$

- Let's introduce the auxiliary function:

$$u(r) = r R(r)$$

- And plug it in above

$$\Rightarrow -\frac{\hbar^2}{2\mu} \left( \frac{u}{R^2} \frac{\partial^2}{\partial r^2} (u) - \ell(\ell+1) \right) + V(r) \frac{u^2}{R^2} - E \frac{u^2}{R^2} = 0$$

Multiply by  $\frac{R^2}{u}$

$$-\frac{\hbar^2}{2\mu} \left[ \frac{\partial^2 u}{\partial r^2} - \ell(\ell+1) \underbrace{\frac{R^2}{u}}_{\frac{u}{r^2}} \right] + u(V-E) = 0$$

Multiply by  $-\frac{2\mu}{\hbar^2}$  and  $\frac{\partial^2}{\partial r^2} \Rightarrow \frac{d^2}{dr^2}$

$$\Rightarrow \frac{d^2}{dr^2} u + \frac{2\mu}{\hbar^2} \left[ E - V(r) - \frac{\ell(\ell+1)}{r^2} \cdot \frac{\hbar^2}{2\mu} \right] u = 0$$

# The Total Wavefunction

- The total wavefunction

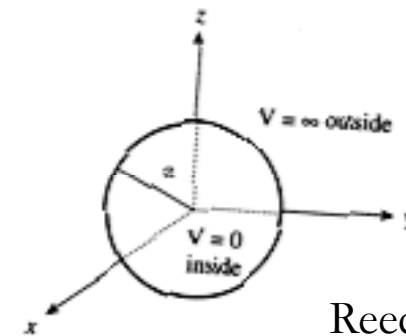
$$\psi_{n\ell m_\ell}(r, \theta, \varphi) = R_{n\ell}(r) Y_{\ell m_\ell}(\theta, \varphi) = \frac{U_{n\ell}(r)}{r} Y_{\ell m_\ell}(\theta, \varphi)$$

- We will soon define  $n$  to be the quantum number that dictates the quantization of energy
- First let's solve this for two simple cases
  - Infinite and finite spherical wells
  - Spherical analogs of particle in a box
  - Interest in nuclear physics: nuclei modeled as spherical potential wells 10's of MeV deep and  $10^{-14}$  m in radius
- Then we will obtain a detailed solution to the Coulomb potential for hydrogen (next time)

# The Infinite Spherical Well

- A particle of mass  $\mu$  is trapped in a spherical region of space with radius  $a$  and an impenetrable barrier

$$V(r) = \begin{cases} \infty & r \geq a \\ 0 & r < a \end{cases}$$



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FIGURE 7.1 Infinite Spherical Well.

- We will only deal with the simplest case of zero angular momentum

$$l = 0$$

- (Otherwise:  $l \neq 0 \Rightarrow$  Bessel & Neumann functions. )

$$\text{So } U(r) \Rightarrow U_0(r)$$

$$\frac{d^2 U_0(r)}{dr^2} + k^2 U_0(r) = 0 \quad \text{inside well.}$$

$$k^2 = \left[ E - V(r) - \frac{l(l+1)}{r^2} \frac{\hbar^2}{2\mu} \right] \frac{2\mu}{\hbar^2}$$

$$= \frac{2\mu E}{\hbar^2}$$

since  $V(r) = 0$  inside.  
and  $l = 0$

- Looks just like for 1D infinite potential well!
- So solutions are:

$$U_0(r) = A \sin kr + B \cos kr$$

# Boundary Conditions

- Recall:  $R = U/r$
- Boundary conditions should prevent divergence for  $R$  as  $r \rightarrow 0$ : Thus  $U = R * r = 0$  as  $r \rightarrow 0$ :

$$\Rightarrow U(r=0) = 0 \Rightarrow B = 0$$

- Boundary conditions also require that the barrier is impenetrable:

$$\Rightarrow U(r=a) = 0 \Rightarrow 0 = A \sin ka \Rightarrow ka = n\pi$$

- This gives us quantization of energy!

$$\Rightarrow E_n(l=0) = \frac{n^2 \pi^2 \hbar^2}{2\mu a^2}$$

# Total Wavefunction

$$\begin{aligned}\Psi_{n00}(r, \theta, \varphi) &= R_n(r) Y_{0,0}(\theta, \varphi) \\ &= \frac{A}{\sqrt{4\pi} r} \sin\left(\frac{n\pi r}{a}\right)\end{aligned}$$

$$\text{since: } Y_{0,0} = \frac{1}{\sqrt{4\pi}} \quad \text{and}$$

$$R_n(r) = \frac{u_n(r)}{r} = \frac{A}{r} \sin\left(\frac{n\pi r}{a}\right)$$

- To obtain A, we apply normalization ...



# Normalization

$$\int_0^a \int_0^\pi \int_0^{2\pi} \psi^* \psi \underbrace{r^2 \sin\theta \, d\varphi \, d\theta \, dr}_{d\text{Volume}} = 1$$

$$\int_0^a \int_0^\pi \int_0^{2\pi} \frac{A^2}{4\pi r^2} \sin^2\left(\frac{n\pi r}{a}\right) \cdot r^2 \sin\theta \, d\varphi \, d\theta \, dr = 1$$

$$\int_0^{2\pi} d\varphi = 2\pi, \quad \int_0^\pi \sin\theta \, d\theta = 2 \quad \int_0^a \sin^2\left(\frac{n\pi r}{a}\right) dr = \frac{a}{2}$$

$$\Rightarrow \frac{A^2}{4\pi} \cdot 2\pi \cdot 2 \cdot \frac{a}{2} = 1 \Rightarrow A = \sqrt{\frac{2}{a}}$$

So,

$$\Psi_{n,0,0} = \frac{1}{\sqrt{2\pi a}} \sin\left(\frac{n\pi r}{a}\right)$$

# Normalization: simpler approach

$$\psi_{n00}(r, \theta, \varphi) = R_n(r) Y_{00}(\theta, \varphi)$$

Because  $Y_{00}$  is already normalized, only  $R_n(r)$  needs  
( $\frac{1}{\sqrt{4\pi}}$ )

to be normalized. In other words

$$1 = \int_0^a \int_0^\pi \int_0^{2\pi} |\psi|^2 r^2 \sin\theta d\varphi d\theta dr$$

$$= \int_0^a |R_n|^2 r^2 dr \underbrace{\left( \int_0^\pi \int_0^{2\pi} |Y_{00}|^2 \sin\theta d\theta d\varphi \right)}_{1}$$

$$= \int_0^a \left( \frac{A}{r} \sin\left(\frac{n\pi}{a} r\right) \right)^2 r^2 dr$$

$$= A^2 \int_0^a \sin^2\left(\frac{n\pi}{a} r\right) dr = A^2 \cdot \frac{a}{2} \Rightarrow A = \sqrt{\frac{2}{a}}$$

# Example: what is $\langle r \rangle$ for a particle in an infinite spherical well?

$$\langle r \rangle = \int_0^a \int_0^\pi \int_0^{2\pi} [\Psi_{n00}^* r \Psi_{n00}] r^2 \sin\theta d\varphi d\theta dr$$

$$= \frac{1}{2\pi a} \left[ \int_0^a r \sin^2\left(\frac{n\pi r}{a}\right) dr \right] \underbrace{\left[ \int_0^\pi \sin\theta d\theta \right]}_2 \underbrace{\left[ \int_0^{2\pi} d\varphi \right]}_{2\pi}$$

$$= \frac{2}{a} \int_0^a r \sin^2\left(\frac{n\pi r}{a}\right) dr$$

$$= \frac{2}{a} \left[ \frac{r^2}{4} - \frac{r \sin\left(\frac{n\pi r}{a}\right)}{4\left(\frac{n\pi}{a}\right)} - \frac{\cos\left(\frac{n\pi r}{a}\right)}{8\left(\frac{n\pi}{a}\right)^2} \right]_0^a$$

$$= \frac{2}{a} \cdot \frac{a^2}{4} = \frac{a}{2}$$

$$\langle r \rangle = \frac{a}{2}$$

$\langle r \rangle$  is independent of  $n$ !

Increasing the energy of the particle changes the probability distribution but not its average position.

This is not the case for non-zero angular momentum!

# The Finite Spherical Well

- Analog of 1-D finite potential well.
- Could describe a particle trapped inside a nucleus

$$V(r) = \begin{cases} V_0 & r \geq a \\ 0 & r < a \end{cases}$$

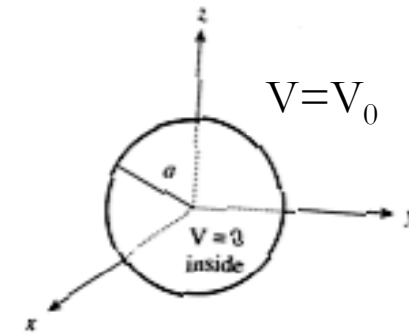


FIGURE 7.1 Infinite Spherical Well.

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- Let's find the bound state solutions,  $E < V_0$
- Again we consider only  $l=0$  case.

# Inside and outside the well

## ■ Inside the well

$$r < a \quad E > V(r) \Rightarrow E > 0$$

$$\frac{d^2 U_0^{\text{in}}}{dr^2} + k_1^2 U_0^{\text{in}} = 0$$

$$k_1^2 = \frac{2\mu E}{\hbar^2}$$

Just like infinite spherical well

$$R_0^{\text{inside}}(r) = \frac{A \sin k_1 r}{r} \quad r < a$$

## ■ Outside the well

$$r > a \quad E < V_0$$

$$\frac{d^2 U_0^{\text{out}}}{dr^2} = k_2^2 U_0^{\text{out}}$$

$$k_2^2 = \frac{2\mu}{\hbar^2} (V_0 - E)$$

$$U_0^{\text{out}}(r) = C e^{k_2 r} + D e^{-k_2 r}$$

$$\Rightarrow R_0^{\text{out}}(r) = \frac{C e^{k_2 r}}{r} + \frac{D e^{-k_2 r}}{r}$$

# Boundary conditions outside the well

- Boundary condition at  $r \rightarrow \infty$ ,  $U(r) \rightarrow 0$   
 $\Rightarrow C = 0$

$$R_{\text{out}}(r) = \frac{D e^{-k_2 r}}{r}$$

- Continuity @  $r=a$  of  $R$  and  $\frac{dR}{dr}$

$$A \sin k_1 a = D e^{-k_2 a}$$

$$-\frac{A \sin k_1 a}{a^2} + \frac{A k_1 \cos k_1 a}{a} = -\frac{D k_2 e^{-k_2 a}}{a} - \frac{D e^{-k_2 a}}{a^2}$$

$$\Rightarrow A [k_1 \cos k_1 a - \sin k_1 a] = -D e^{-k_2 a} (a k_2 + 1)$$

- Divide the two equations:

$$\begin{aligned} k_1 a \cot k_1 a - 1 &= -(a k_2 + 1) \\ k_1 a \cot k_1 a &= -a k_2 \end{aligned}$$

- Define:

$$\xi = k_1 a, \quad \eta = k_2 a \quad \xi \cot \xi = -\eta$$

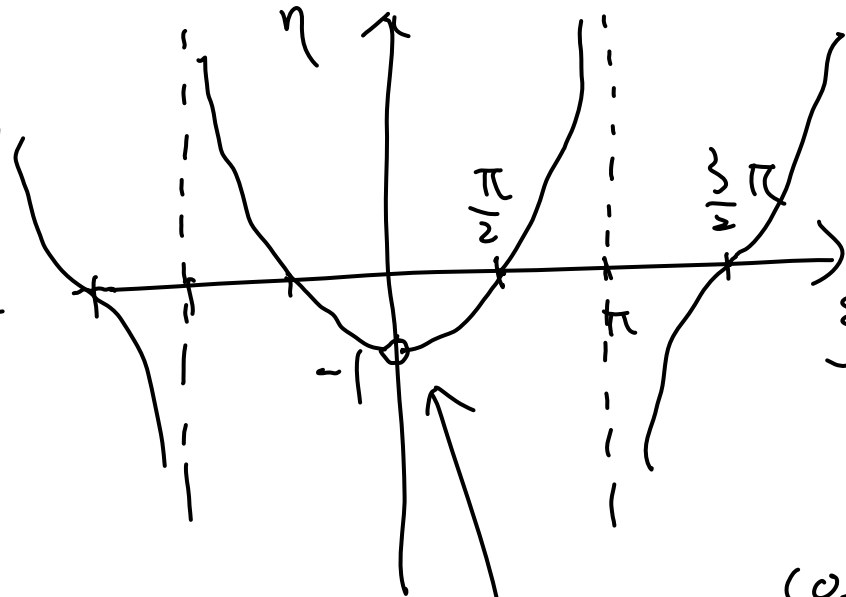
This is a transcendental equation!

$$\begin{aligned} \xi^2 + \eta^2 &= a^2 (k_1^2 + k_2^2) = a^2 \left[ \frac{2\mu E}{\hbar^2} + \frac{2\mu (V_0 - E)}{\hbar^2} \right] \\ &= a^2 \cdot \frac{2\mu V_0}{\hbar^2} = \text{const} \sim \rho^2 \end{aligned}$$

$\rho^2$  is the strength parameter

# Sketch functions

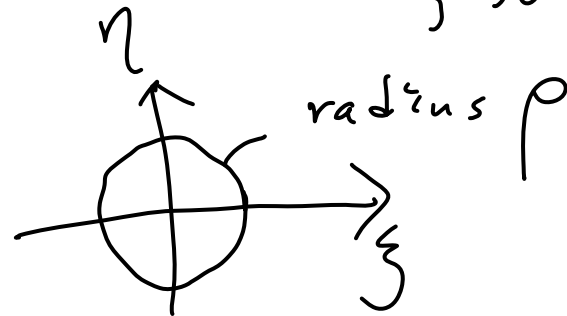
$$\eta = - \underbrace{\left( \underbrace{\xi}_{\text{odd}} \cot \underbrace{\xi}_{\text{odd}} \right)}_{\text{even}}$$



$$\cot \frac{\pi}{2} = \frac{\cos \frac{\pi}{2}}{\sin \frac{\pi}{2}} = \frac{0}{1} = 0$$

$$\lim_{\xi \rightarrow 0} \xi \frac{\cos \xi}{\sin \xi} = 1$$

$$\xi^2 + \eta^2 = \rho^2 \Rightarrow$$





# Energy Eigenvalues

- This is an equation that describes a circle of radius  $\rho$  in the  $(\xi, \eta)$  plane.

$$\xi^2 + \eta^2 = \rho^2$$

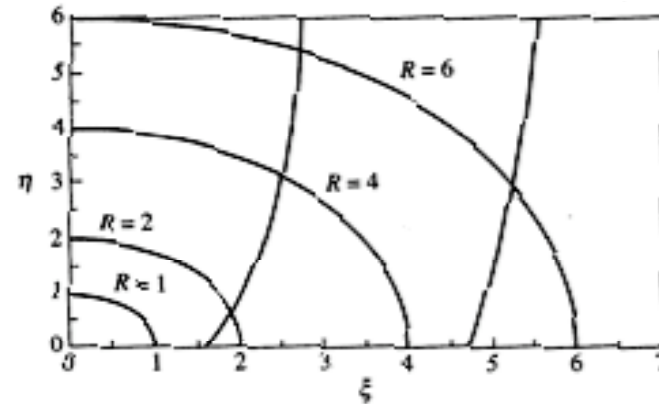


FIGURE 7.2 Illustration of the Solution for Energy Eigenvalues for a Finite Spherical Well.

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- To find allowed bound state energies need to satisfy this equation and transcendental equation simultaneously
- Points of intersection of circles and the cotangent curves correspond to the quantized energy levels

# Energy Eigenvalues

- Zeros of cotangent occur at:

$$\cot z = 0$$

$$(2n-1)\frac{\pi}{2}, \quad n=1, 2, 3, \dots$$

- For a spherical well to possess  $n$  bound states it must have:

$$\rho \geq (2n-1)\frac{\pi}{2}$$

- Or ...
- $$\rho^2 \geq (2n-1)^2 \frac{\pi^2}{4} \Rightarrow V_0 a^2 \geq \frac{(2n-1)^2 \pi^2 \hbar^2}{8\mu}$$

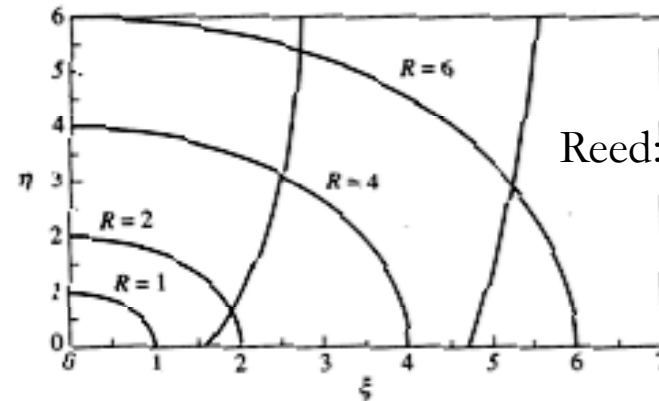


FIGURE 7.2 Illustration of the Solution for Energy Eigenvalues for a Finite Spherical Well.

Note  $\rho=1$  has no bound states!

$$\begin{aligned} \xi^2 + \eta^2 &= a^2(k_1^2 + k_2^2) = a^2 \left( \frac{2\mu E}{\hbar^2} + \frac{2\mu(V_0 - E)}{\hbar^2} \right) \\ &= a^2 \cdot \frac{2\mu V_0}{\hbar^2} = \text{const} = \rho^2 \end{aligned}$$

# Example

- How many energy states are available to an alpha-particle trapped in a finite spherical well of depth 50 MeV and radius  $10^{-14}$  m? Assume zero angular momentum.

$$(2n-1)^2 \leq \frac{8\mu a^2 V_0}{\hbar^2 \pi^2}$$

$$(2n-1)^2 \leq \frac{8(6.646 \times 10^{-27} \text{ kg})(10^{-14} \text{ m})^2 (\overbrace{50 \text{ MeV} \times 1.6 \times 10^{-19} \frac{\text{J}}{\text{eV}}})}{(1.055 \times 10^{-34} \text{ J}\cdot\text{s})^2 \pi^2}$$

$$(2n-1)^2 \leq 387$$

$$n \leq 10.3$$

- What is the energy of the lowest energy bound state for this system?

$$\zeta^2 + \eta^2 = \frac{2\mu a_0^2 V_0}{\hbar^2} = 955$$

$$\zeta \cot \zeta = -\eta$$

$$\zeta \cot \zeta + \sqrt{955 - \zeta^2} = 0$$

Minimize  $\zeta = 3.04305$

$$E = \frac{\hbar^2 \zeta^2}{2\mu} = \frac{\zeta^2}{a_0^2} \frac{\hbar^2}{2\mu} = \frac{(3.04305)^2 (1.055 \times 10^{-31} \text{ J})^2}{(10^{-14} \text{ m})^2 \cdot 2 \cdot (6.646 \times 10^{-27} \text{ kg})}$$

$$= 7.75 \times 10^{-14} \text{ J} = \underline{0.48 \text{ MeV}}$$

# Summary/Announcements

- Next time: The Coulomb Potential of the Hydrogen atom