

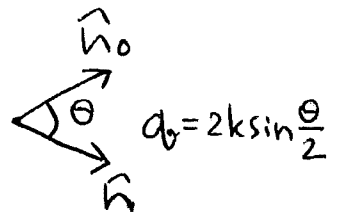
HW #6 solutions

(10.9) a) Use the 1st Born approx'n
(ϵ_r close to 1):

$$\frac{\vec{E}^* \cdot \vec{A}_{sc}^{(1)}}{D_0} = \frac{k^2}{4\pi\epsilon_0} \int d^3x e^{i\vec{q} \cdot \vec{x}} \left[\vec{E}^* \cdot \vec{E}_0 \frac{\delta\epsilon}{\epsilon_0} + (\hat{n}_0 \times \vec{E}_0) (\hat{n} \times \vec{E}^*) \frac{\delta\mu}{\mu_0} \right] \quad (10.31)$$

Here, $\vec{q} = k(\hat{n}_0 - \hat{n})$, so that

$$q^2 = 2k^2(1 - \cos\theta) = 4k^2 \sin^2 \frac{\theta}{2}$$



Further, $\delta\mu = 0$ and

$$\frac{\delta\epsilon}{\epsilon_0} = \begin{cases} \epsilon_r - 1 & r \leq a \\ 0 & r > a \end{cases}$$

Then

$$\frac{\vec{E}^* \cdot \vec{A}_{sc}^{(1)}}{D_0} = \frac{k^2}{4\pi\epsilon_0} (\epsilon_r - 1) (\vec{E}^* \cdot \vec{E}_0) \times \underbrace{\int_{r < a} d^3x e^{i\vec{q} \cdot \vec{x}}}_I$$

$$\begin{aligned} I &= 2\pi \int_0^a dr r^2 \int_{-1}^1 dx e^{iqr} x = \\ &= \frac{4\pi}{q} \int_0^a dr r \sin(qr) = \frac{4\pi}{q^3} [\sin(qa) - qa \cos(qa)] \end{aligned}$$

Consequently,

$$\frac{\vec{E}^* \cdot \vec{A}_{sc}^{(1)}}{D_0} = \frac{(ka)^2}{q} (\epsilon_r - 1) (\vec{E}^* \cdot \vec{E}_0) \times$$

$$\times \underbrace{\frac{\sin(qa) - qa \cos(qa)}{(qa)^2}}_{j_1(qa) - \text{Bessel f'n with } l=1}$$

$$\text{Then } \frac{d\sigma}{d\Omega} = \left| \frac{\vec{E}^* \cdot \vec{A}_{sc}^{(1)}}{D_0} \right|^2 = k^4 a^6 |\epsilon_r - 1|^2 \times$$

$$\times |\vec{E}^* \cdot \vec{E}_0|^2 \left(\frac{j_1(qa)}{qa} \right)^2.$$

The unpolarized cross-section is

$$\frac{d\sigma^{unp}}{d\Omega} = k^4 a^6 |\epsilon_r - 1|^2 \left(\frac{j_1(qa)}{qa} \right)^2 \frac{1 + \cos^2 \theta}{2}$$

Note that if $ka \ll 1 \Rightarrow qa \ll 1$,

and $j_1(qa) \rightarrow \frac{qa}{3}$ as $qa \rightarrow 0$.

$$\text{So } \frac{d\sigma^{unp}}{d\Omega} \xrightarrow{qa \ll 1} \frac{k^4 a^6}{9} |\epsilon_r - 1|^2 \frac{1 + \cos^2 \theta}{2}.$$

If $ka \gg 1 \Rightarrow qa = 2ka \sin \frac{\theta}{2} \gg 1$,

and $j_1(qa) \xrightarrow{qa \gg 1} -\frac{\cos(qa)}{qa}$

true for all angles away from $\theta = 0$

So, for θ away from 0

$$\frac{d\sigma_{\text{unp}}}{d\Omega} \underset{qa \gg 1}{\sim} \frac{1}{(qa)^4} = \frac{1}{(2ka \sin \frac{\theta}{2})^4}$$

falls off rapidly

In σ_{unp} , $\int d\Omega$ will be highly peaked around $\theta=0$:

$$\sigma_{\text{unp}} \sim \int d\Omega \underbrace{\frac{1+\cos^2\theta}{2}}_{\approx 1} \left(\frac{j_1(qa)}{qa} \right)^2 \approx$$

$$\approx 2\pi \int_0^\pi \sin\theta d\theta \left(\frac{j_1(qa)}{qa} \right)^2 \approx 2\pi \int_0^\pi d\theta \theta \left(\frac{j_1(ka\theta)}{ka\theta} \right)^2 \quad (\approx)$$

$qa \approx ka\theta$ if θ is small
 extend to $+\infty$

$$\approx \frac{2\pi}{(ka)^2} \int_0^\infty j_1^2(x) \frac{dx}{x} = \frac{\pi}{2(ka)^2}$$

Finally,

$$\sigma_{\text{unp}} \approx \frac{\pi a^2}{2} (ka)^2 |E_r^{-1}|^2, \text{ as desired}$$

Note that $ka \ll 1 \Rightarrow \sigma_{\text{unp}} \sim k^4$

$ka \gg 1 \Rightarrow \sigma_{\text{unp}} \sim k^2$

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11.1 a) $y' = y$ (motion in x direction)
 $z' = z$

Homogeneity of space-time \Rightarrow linear

transform:
$$\begin{cases} x' = f_1 x + f_2 t \\ t' = g_1 x + g_2 t \end{cases}$$

f_1, f_2, g_1, g_2 do not depend on x & t but are otherwise arbitrary functions of v .

Inspired by Galilean relativity, we take v out as follows:

$$\begin{cases} x' = f_1 x - v \overbrace{f_2}^{\tilde{f}_2} t & \Leftarrow \text{expect} \\ t' = g_2 t - v \underbrace{\tilde{g}_1}_{g_1} x & \begin{matrix} x' = x - vt \\ \text{as } \frac{v}{c} \rightarrow 0 \end{matrix} \end{cases}$$

But due to homogeneity the points along x cannot spread out as time increases $\Rightarrow f_1 = \tilde{f}_2$.

Moreover, due to isotropy expect

f_1, g_2, \tilde{g}_1 to be functions of v^2 (no dependence on sign of v)

Finally,
$$\begin{cases} x' = f_1(v^2)x - f_1(v^2)v t, & y' = y, \\ t' = g_2(v^2)t - \tilde{g}_1(v^2)v x, & z' = z \end{cases}$$

The inverse can be obtained by

$v \rightarrow -v$:

$$\begin{cases} x = f_1(v^2) x' + f_1(v^2) v t', & y = y', \\ t = g_2(v^2) t' + \tilde{g}_1(v^2) v x', & z = z'. \end{cases}$$

b) Transform $K \rightarrow K'$ & then $K' \rightarrow K$,
should recover original coordinates:

$$\begin{aligned} x &= f_1 x' + f_1 v t' = f_1 (f_1 x - f_1 v t) + \\ &\quad + f_1 v (g_2 t - \tilde{g}_1 v x) = \\ &= \underbrace{(f_1^2 - f_1 \tilde{g}_1 v^2)}_{\substack{\Downarrow \\ f_1(f_1 - v^2 \tilde{g}_1) = 1}} x + \underbrace{(f_1 v g_2 - f_1^2 v)}_{\substack{= 0 \\ \Downarrow \\ f_1 = g_2}} t \end{aligned}$$

Likewise,

$$\begin{aligned} t &= g_2 t' + \tilde{g}_1 v x' = g_2 (g_2 t - \tilde{g}_1 v x) + \\ &\quad + \tilde{g}_1 v (f_1 x - f_1 v t) = \\ &= \underbrace{(\tilde{g}_1 v f_1 - \tilde{g}_1 v g_2)}_{\substack{= 0 \\ \Downarrow \\ f_1 = g_2}} x + \underbrace{(g_2^2 - \tilde{g}_1 f_1 v^2)}_{\substack{= 1 \\ \Downarrow \\ g_2^2 = 1 + \tilde{g}_1 f_1 v^2}} t \end{aligned}$$

$f_1^2 - \tilde{g}_1 f_1 v^2 = 1$, same as above

So now
$$\begin{cases} \tilde{g}_1 = \frac{f_1^2 - 1}{f_1 v^2}, \\ g_2 = f_1 \end{cases}$$

The transform only depends on $f_1(v^2)$.

c)
$$u = \frac{dx}{dt} = \frac{f(dx' + v dt')}{f dt' + v dx' \frac{f^2 - 1}{f v^2}} \quad (\equiv)$$

↗
rename $f_1 \rightarrow f$ for simplicity

$$u' = \frac{dx'}{dt'}$$

$$(\equiv) \frac{u' + v}{1 + v u' \underbrace{\frac{f^{-1} f}{f v^2}}_{\tilde{g}_1 / f}}$$

Universal limiting speed C :

$$C = \frac{C + v}{1 + v C \frac{f^{-1} f}{f v^2}} \Rightarrow f = \frac{1}{\sqrt{1 - \frac{v^2}{C^2}}} = \gamma$$

Finally,
$$\begin{cases} x' = \gamma(x - vt) \\ t' = \gamma\left(t - \frac{vx}{C^2}\right) \end{cases}, \quad \begin{cases} y' = y \\ z' = z \end{cases}$$

$$\frac{f^2 - 1}{f v} = \frac{\gamma^2 - 1}{\gamma v} = \frac{\gamma v}{C^2} \left(\frac{\gamma - 1/\gamma}{\gamma \left(\frac{v}{C}\right)^2} \right) = \frac{C}{\gamma v} \frac{\gamma^2 - 1}{\gamma^2 \left(\frac{v}{C}\right)^2} = \frac{\gamma v}{C^2}$$

Likewise,

$$\begin{cases} x = \gamma(x' + vt'), & y = y', \\ t = \gamma(t' + \frac{vx'}{c^2}), & z = z'. \end{cases}$$

$$v \rightarrow -v$$

$c = c$ (speed of light)
by experiments

11.3 Consider two Lorentz boosts along x -axis:

$$A_1 = \begin{pmatrix} \gamma_1 & -\beta_1 \gamma_1 & 0 & 0 \\ -\beta_1 \gamma_1 & \gamma_1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix} \quad \beta_1 = \frac{v_1}{c}$$

$$\gamma_1 = \frac{1}{\sqrt{1-\beta_1^2}}$$

and same for A_2 (with β_2, γ_2).

Then

$$A_2 A_1 = \begin{pmatrix} \gamma_1 \gamma_2 (1 + \beta_1 \beta_2) & -\gamma_1 \gamma_2 (\beta_1 + \beta_2) & 0 & 0 \\ -\gamma_1 \gamma_2 (\beta_1 + \beta_2) & \gamma_1 \gamma_2 (1 + \beta_1 \beta_2) & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$$

Note that

$$\begin{aligned} \gamma_1 \gamma_2 (1 + \beta_1 \beta_2) &= \frac{1}{\sqrt{\frac{(1-\beta_1^2)(1-\beta_2^2)}{(1+\beta_1 \beta_2)^2}}} = \\ &= \frac{1}{\sqrt{1 - \left(\frac{\beta_1 + \beta_2}{1 + \beta_1 \beta_2}\right)^2}} = \frac{1}{\sqrt{1 - \frac{1}{c^2} \left(\frac{v_1 + v_2}{1 + \frac{v_1 v_2}{c^2}}\right)^2}} = \\ &= \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} \equiv \gamma \quad \beta = \frac{v}{c} \end{aligned}$$

Likewise,

$$\begin{aligned} -\gamma_1 \gamma_2 (\beta_1 + \beta_2) &= -\frac{\beta_1 + \beta_2}{1 + \beta_1 \beta_2} \overbrace{\gamma_1 \gamma_2 (1 + \beta_1 \beta_2)}^{\gamma} = \\ &= -\frac{v}{c} \gamma = -\beta \gamma \end{aligned}$$

Finally,

$$A_2 A_1 = \begin{pmatrix} \gamma & -\beta \gamma & 0 & 0 \\ -\beta \gamma & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}.$$

This is a Lorentz boost with

$$v = \frac{v_1 + v_2}{1 + \frac{v_1 v_2}{c^2}}.$$

Note that $v \neq v_1 + v_2$ at relativistic speeds