

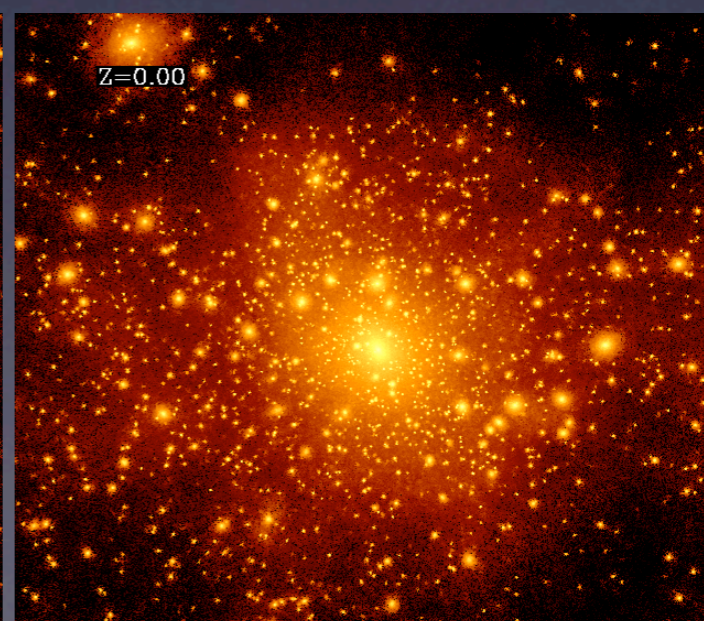
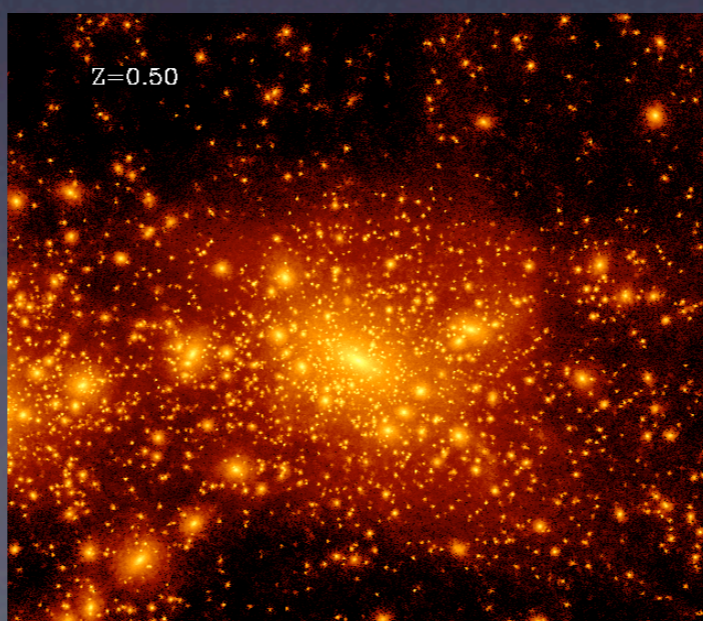
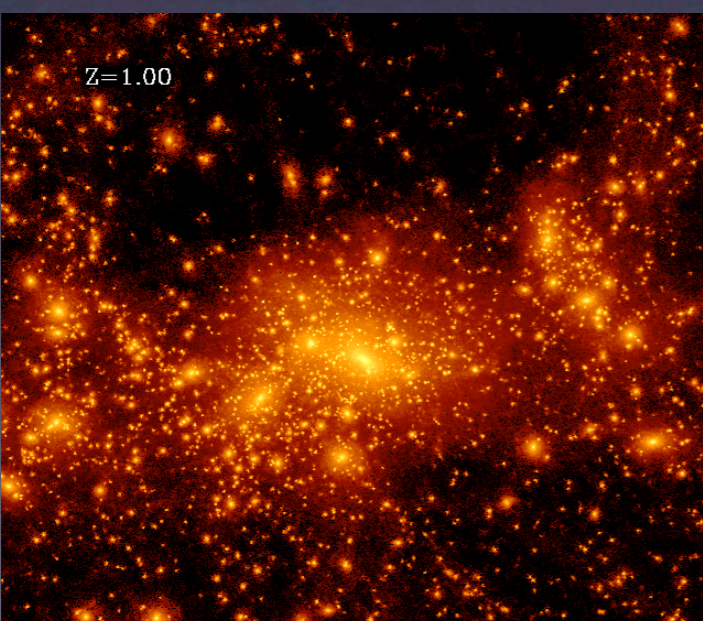
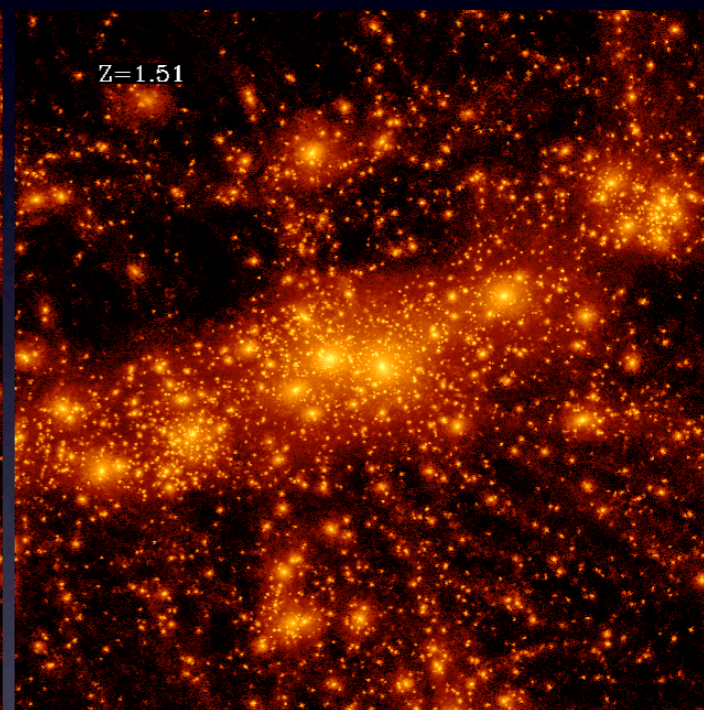
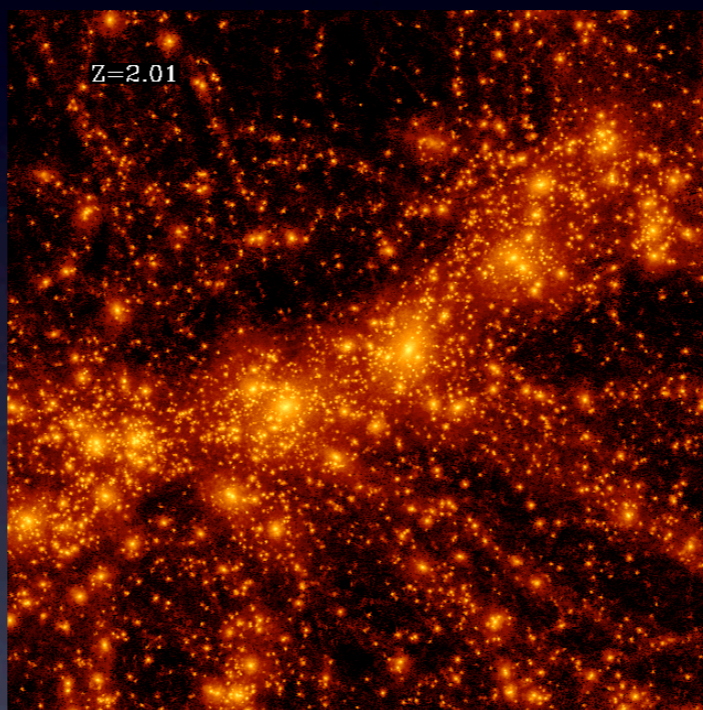
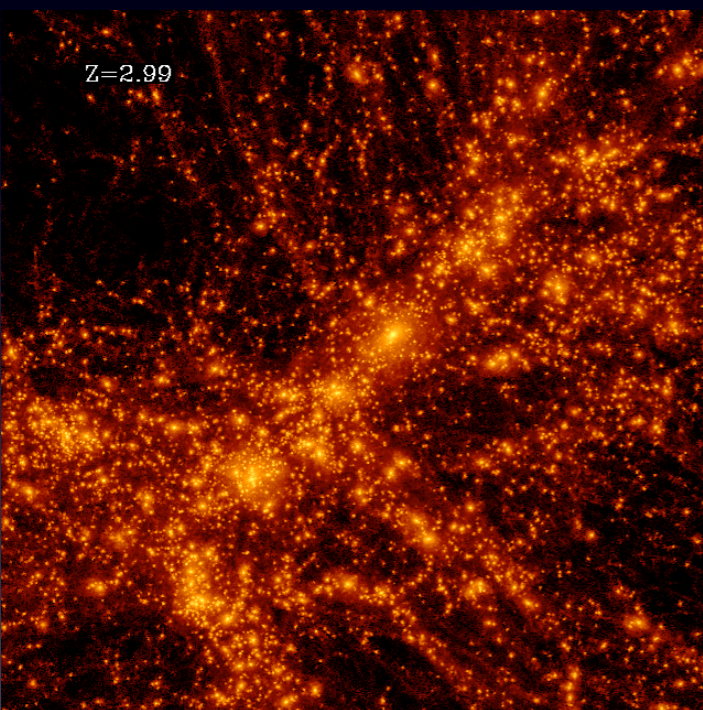
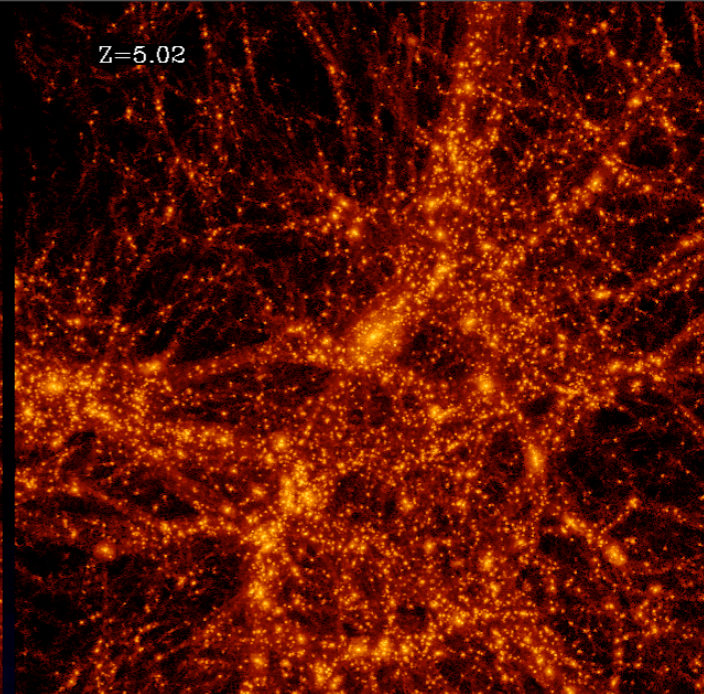
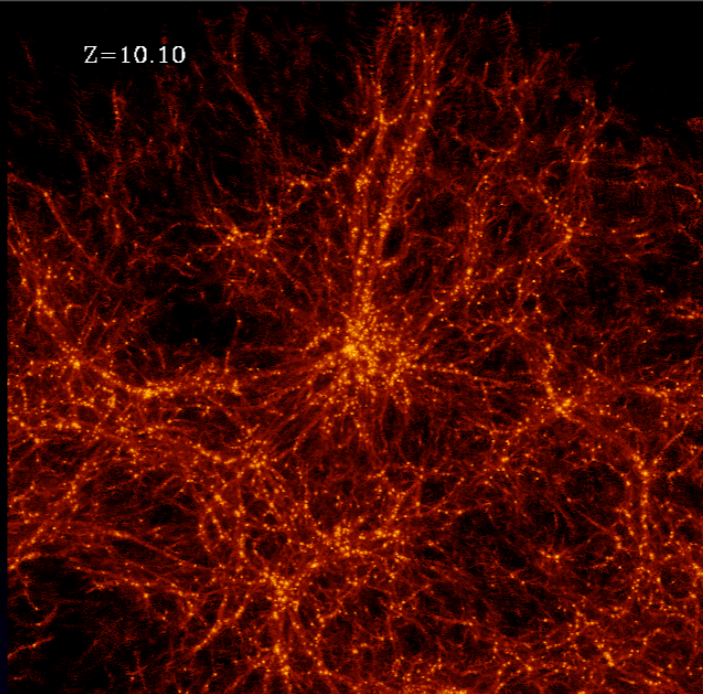
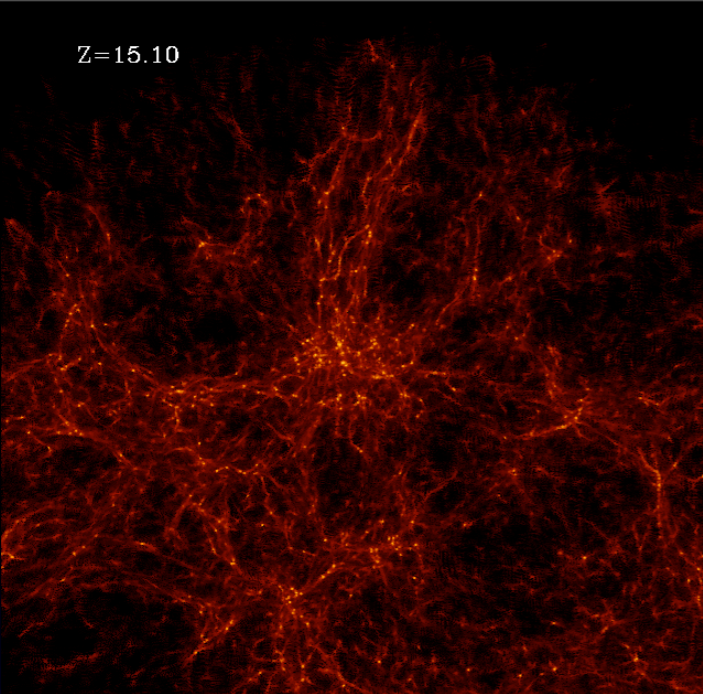
# Dark Matter Substructure and Dwarf Galactic Satellites

By Andrey Kravtsov  
Presented By Curtis McCully

# Outline



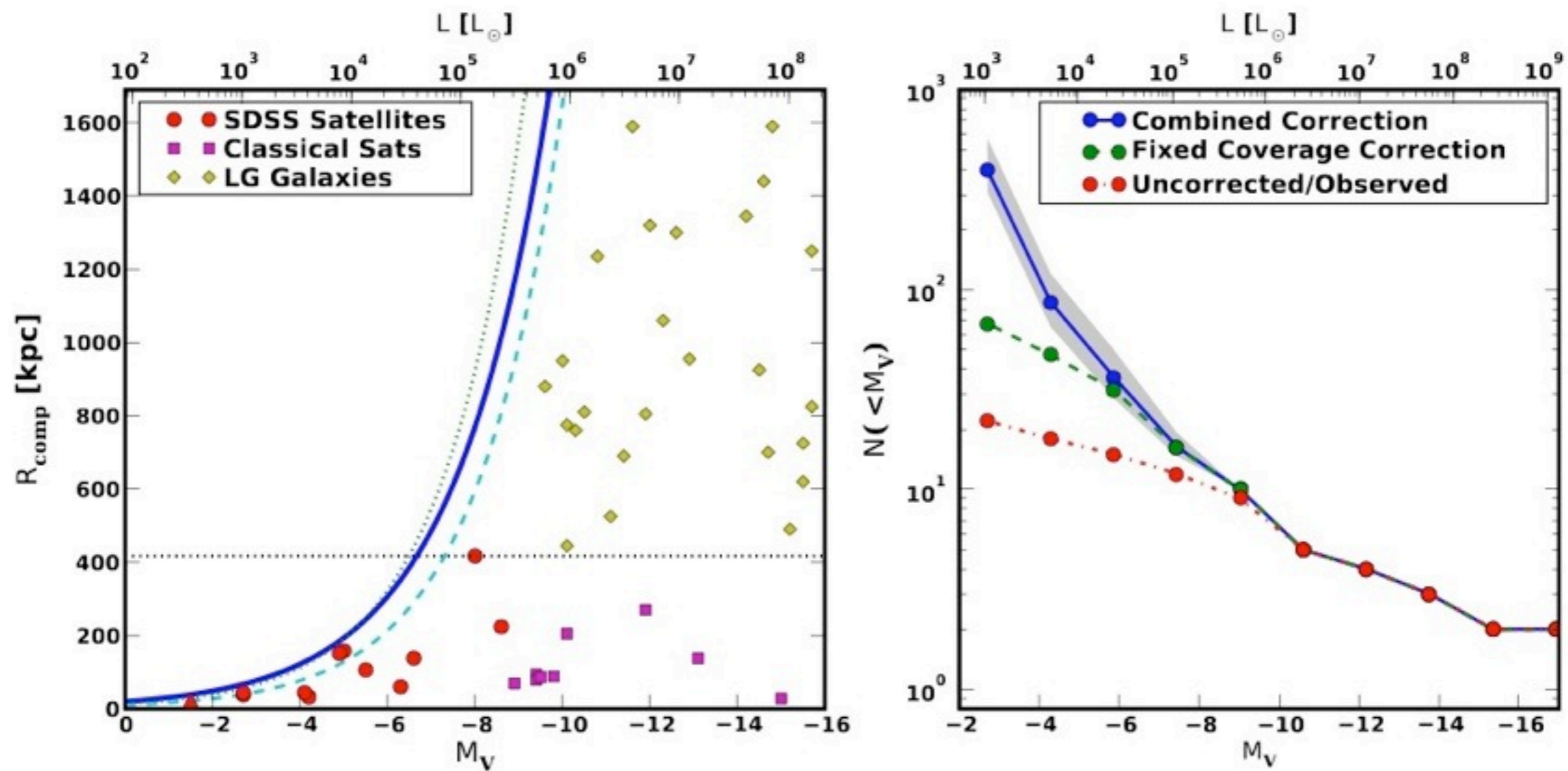
- Introduction
  - CDM Simulations vs Observations
- Quantifying the Substructure Problem
  - Dark Satellites
  - Luminous Satellites
- Possible Solutions
  - New Cosmological Model
  - Galaxy Formation Suppression



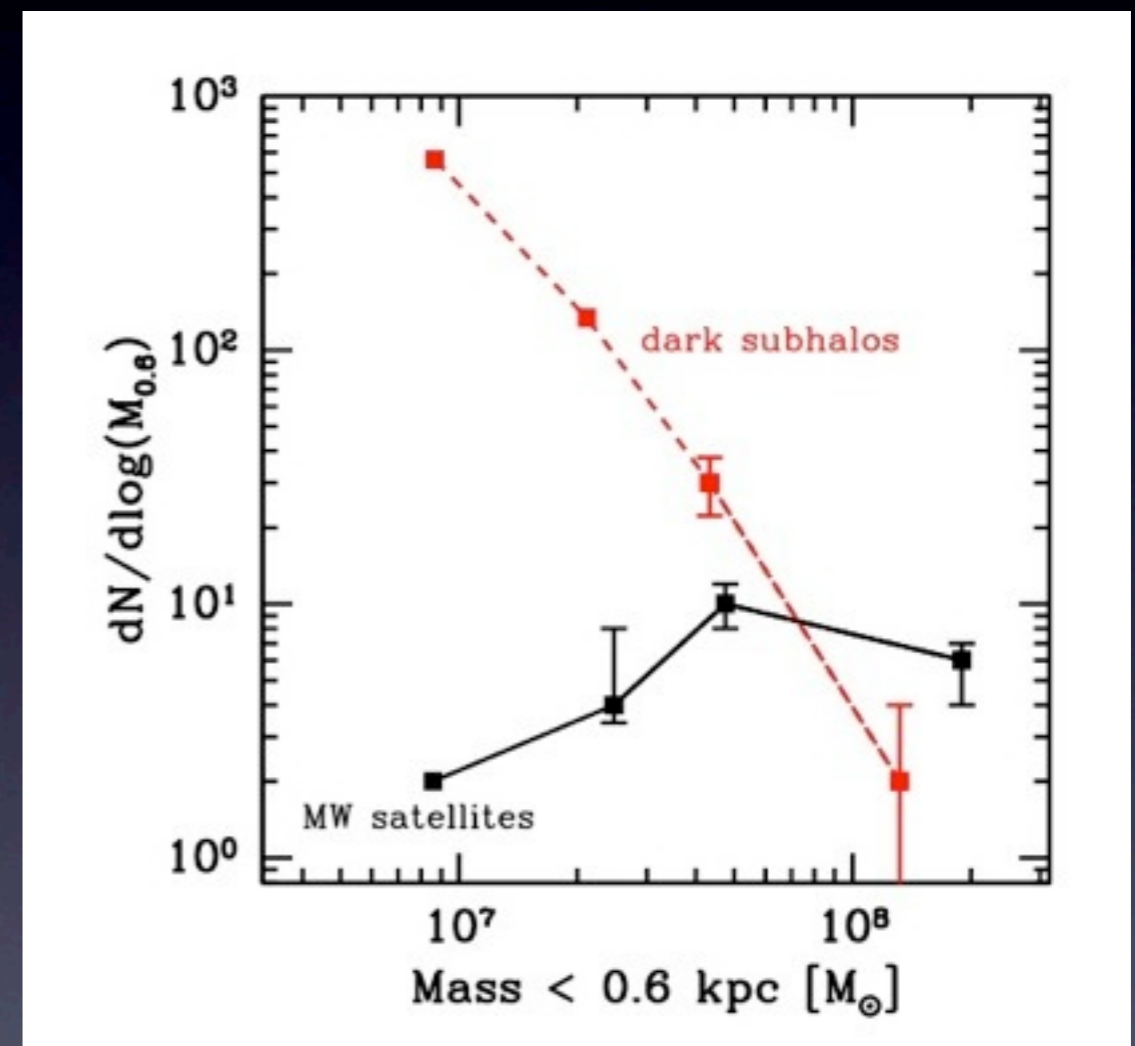
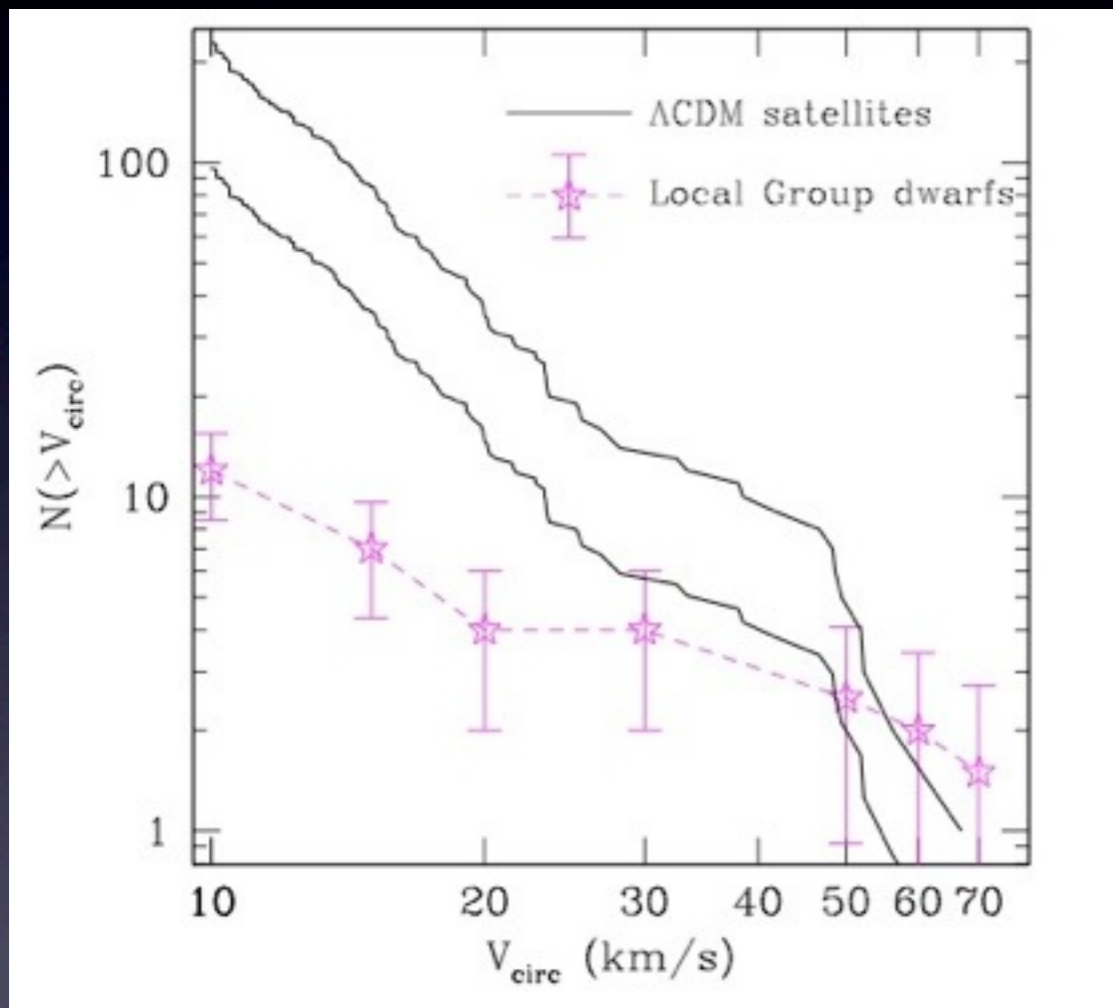
# Via Lactea Movies can be found at the following url

- <http://www.ucolick.org/%7Ediemand/vl/movies.html>

# Luminous Satellite Galaxies



# Circular Velocity and $m_{0.6}$ Slopes





# Defining the “Substructure Problem”

“The substructure problem can be stated as the discrepancy in the slopes of the circular velocity and  $m_{0.6}$  mass functions inferred for observed satellites of the Milky Way and the slopes of these functions predicted for dark matter subhalos in the MW-sized host halos formed in the concordance  $\Lambda$ CDM cosmology”

# Different Cosmology



- Suppressing density fluctuations
- Warm Dark Matter

$$\begin{aligned}d &= 2R = 2 \left( \frac{3M}{4\pi\Omega_{m0}\rho_{\text{crit}0}} \right)^{1/3} & (2) \\ &= 360.4 \text{ kpc} \left( \frac{M}{10^9 M_{\odot}} \frac{0.3}{\Omega_m} \right)^{1/3} \left( \frac{H_0}{70} \right)^{-2/3},\end{aligned}$$

where  $\Omega_m$  is the present-day total matter density in units of the present-day critical density,  $\rho_{\text{crit}0} \equiv 3H_0^2/8\pi G$  and  $H_0$  is the current Hubble constant in units of km/s/Mpc.



# Galaxy Formation Subhalo Suppression



- Gas photoevaporated after reionization
- Supernovae Feedback
- Satellites could be in larger halos

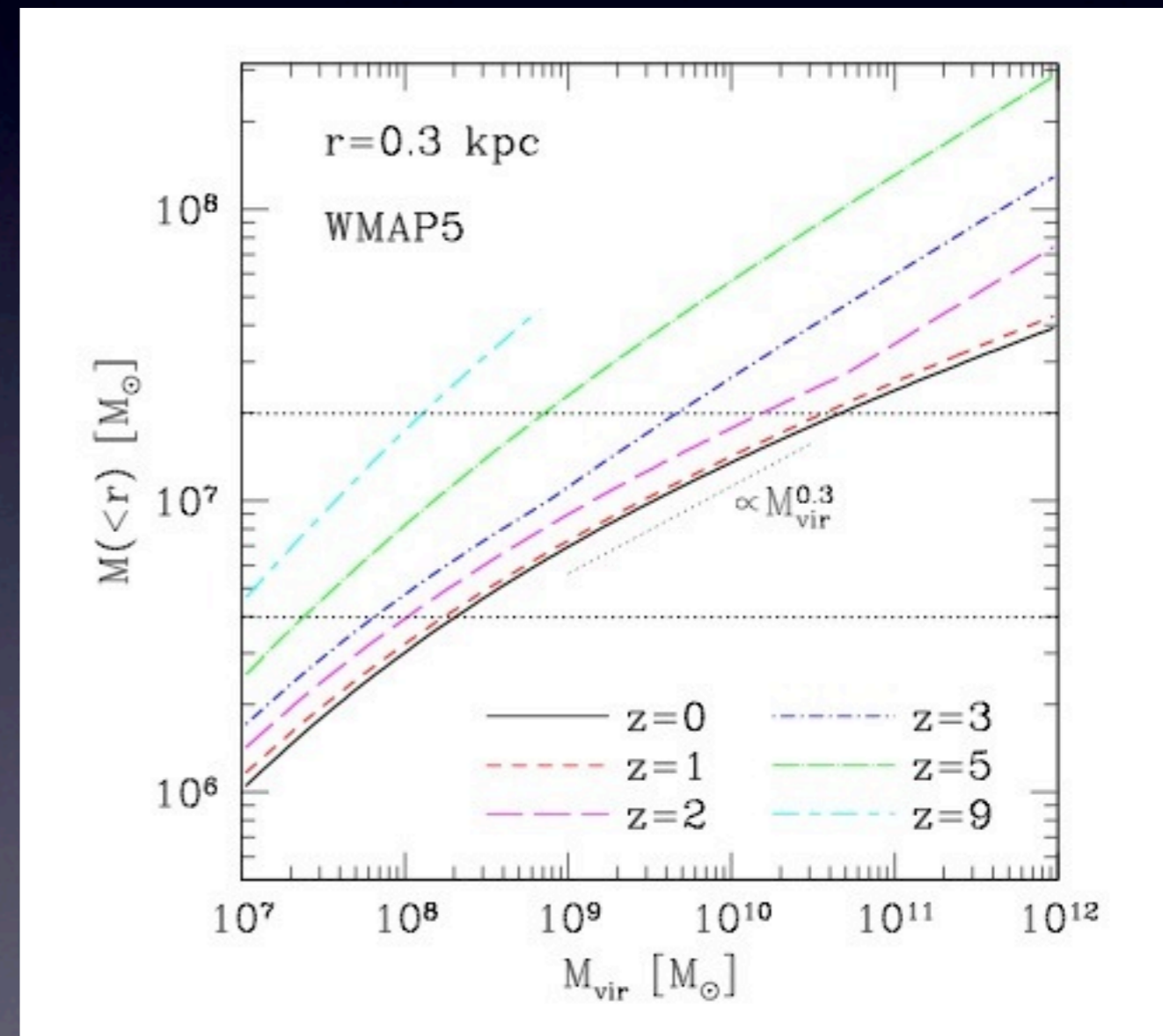
# Conclusions



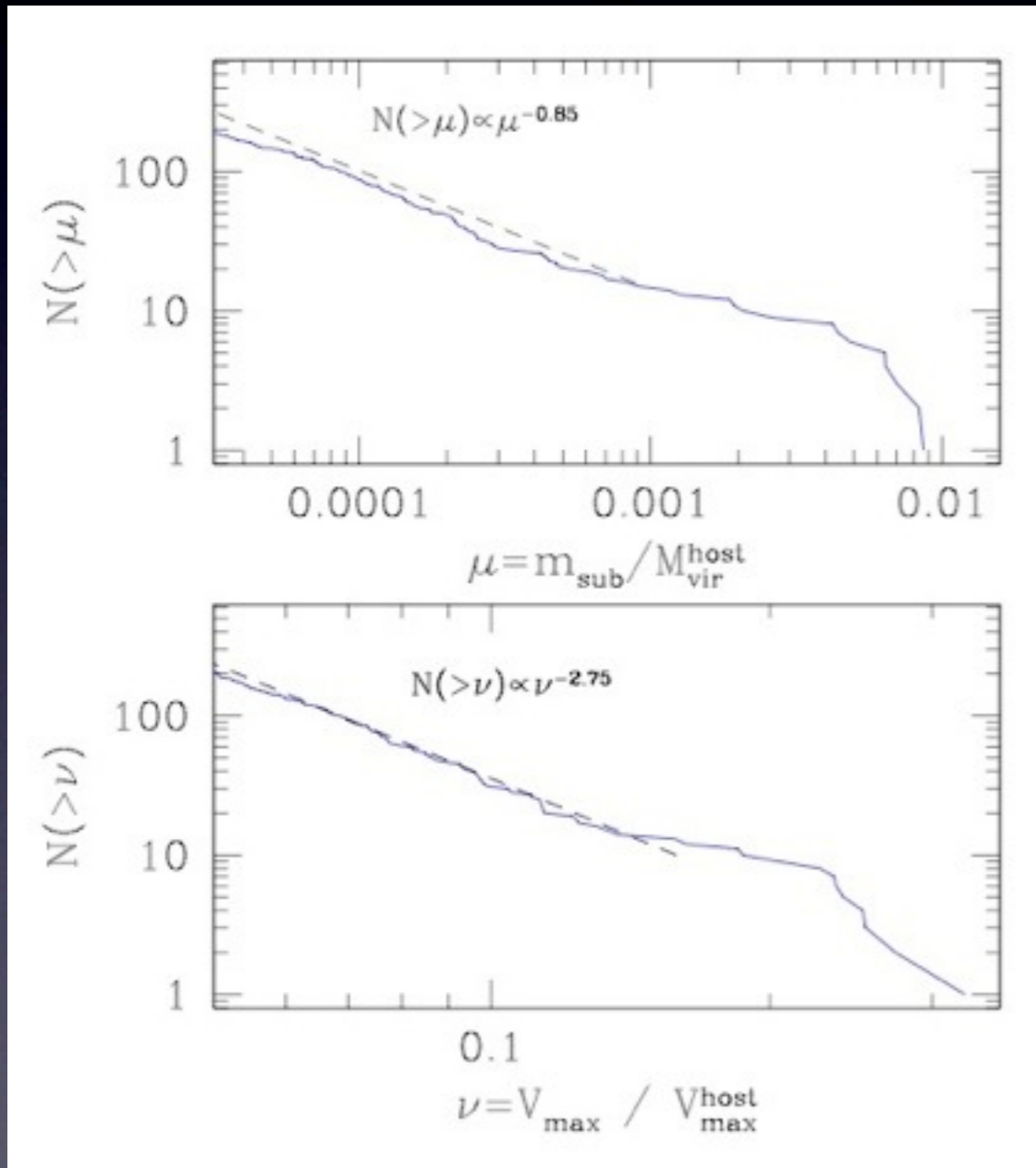
- Models and Observation do not agree
- Satellites could be mostly dark
- If satellites are dark, need low star forming efficiency
- Mass-dependent suppression mechanism

# Luminous Satellite Models

- Threshold Galaxy Formation
- Selective Galaxy Formation



# Subhalo Populations



$$V_{\text{max}} = \text{max} \left( \frac{Gm(<r)}{r} \right)^{1/2}$$

$$m(<r) = 4\pi \int \rho(r)r^2 dr$$

# Luminous Satellites (Cont.)

