

Superconductivity near a quantum critical point in the extreme retardation regime

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We study fermions at quantum criticality with extremely retarded interactions of the form $V(\omega_l) = (g/|\omega_l|)^\gamma$, where ω_l is the transferred Matsubara frequency. This system undergoes a normal-superconductor phase transition at a critical temperature $T = T_c$. The order parameter is the frequency-dependent gap function $\Delta(\omega_n)$ as in the Eliashberg theory. In general, the interaction is extremely retarded for $\gamma \gg 1$, except at low temperatures $\gamma > 3$ is sufficient. We evaluate the normal state specific heat T_c , the jump in the specific heat $\Delta(\omega_n)$ near T_c , and the Landau free energy. Our answers are asymptotically exact in the limit $\gamma \rightarrow \infty$. At low temperatures, we prove that the global minimum of the free energy is nondegenerate and determine the order parameter, the free energy, and the specific heat. These answers are exact for $T \rightarrow 0$ and $\gamma > 3$. We also uncover and investigate an instability of the γ model: Negative specific heat at $T \rightarrow 0$ and just above T_c .

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I. INTRODUCTION

Critical fluctuations of the order parameter field mediate strong electron-electron interactions near metallic quantum critical points. Coupling to this bosonic order parameter field induces attraction between electrons, leading to the enhancement of superconductivity near such points in many quantum materials, such as the cuprates, iron pnictides, and certain heavy-fermion materials [1–4]. On the other hand, same interactions destroy Fermi liquid quasiparticles thus undermining superconductivity, which builds on quasiparticle coherence. As a result, the enhancement does not occur in other quantum critical materials, such as heavy-fermion metals $\text{CeCu}_{6-x}\text{Au}_x$ and YbRh_2Si_2 [5].

It has been argued [6] that many quantum critical system are described by the γ model—effective electron-electron interactions of the form $V(\omega_l) = (g/|\omega_l|)^\gamma$, where $\omega_l = 2\pi Tl$ is the transferred Matsubara frequency. Examples include a 2D nematic ($\gamma = 1/3$) and magnetic ($\gamma = 1/2$) quantum critical points, a spin-liquid model for the cuprates ($\gamma = 0.7$), 2D pairing mediated by an undamped propagating boson ($\gamma = 1$), and the strong coupling limit of phonon mediated superconductivity ($\gamma = 2$). Interactions of this form obtain by integrating out the bosonic fields and averaging over the Fermi surface, taking into account that the order parameter field is massless at criticality.

Here we study the thermodynamic properties of the γ model in the extreme retardation regime, when $V(\omega_l)$ is effectively local in the Matsubara frequency domain. This corresponds to $\gamma > 3$ at low temperatures and to $\gamma \gg 1$ at arbitrary T . Even though in all the above examples $0 \leq \gamma \leq 2$, there is no fundamental reason why larger γ are impossible. Moreover, properties of the system at finite temperature are continuous in γ . For example, the relative difference between

the exact superconducting transition temperature T_c for $\gamma = 2$ and the large γ asymptote for T_c is only 6% (see below).

The attractive feature of the γ model is that there is a single intrinsic dimensionless parameter in the model—the exponent γ itself. The coupling constant g sets the energy units. The momentum-dependence of the fermion-boson coupling and the spectrum of mediating bosons become irrelevant at criticality. The Fermi energy is the largest energy in the problem, i.e., we are in the limit $\varepsilon_F/g \rightarrow \infty$. This suggests that the γ model is the minimal description of the strong coupling fixed point (quantum critical point) in certain materials.

The interaction as we wrote it diverges at $\omega_l = 0$. Moving slightly away from criticality regularizes this divergence. The particular form of the regularization is unimportant, because it does not affect our results. A convenient regularization is $V(\omega_l) = \frac{g^\gamma}{|\omega_l|^\gamma + \Omega^{2\gamma}}$, where Ω has the meaning of the critical boson mass. Now it is straightforward to analyze the interaction in the time domain. In the $\gamma \rightarrow \infty$ limit, $V(0) = \frac{g^\gamma}{\Omega^{2\gamma}} \equiv \lambda$ and $V(\omega_l) = 0$ for $\omega_l \neq 0$. Its Fourier transform is $\tilde{V}(\tau' - \tau) = \lambda T$. This interaction is *maximally retarded*—independent of the time separation between the interacting particles both in real and imaginary time domains. By contrast, instantaneous interactions, e.g., the pairing interaction in the Bardeen-Cooper-Schrieffer (BCS) theory, are proportional to $\delta(\tau' - \tau)$. When γ is large but finite, the interaction is *extremely retarded*.

The γ model is a generalization of the strong coupling limit of the Eliashberg theory [7–9], which corresponds to $\gamma = 2$. It presents two phases: Normal and superconducting. The order parameter is the frequency dependent gap function $\Delta_n \equiv \Delta(\omega_n)$ as in the Eliashberg theory. Here $\omega_n = \pi T(2n + 1)$ is the fermionic Matsubara frequency. We focus on the thermodynamic properties near the superconducting transition temperature T_c and at low T . Our results near T_c are asymp-

totically exact in the limit $\gamma \rightarrow \infty$ and the low temperature answers are similarly exact for $T \rightarrow 0$ at any $\gamma > 3$.

First, we determine the T_c itself,

$$T_c(\gamma) = \frac{ga^{\frac{1}{\gamma}}}{2\pi}, \quad a \approx 1.1843, \quad (1)$$

and then the order parameter near T_c

$$\Delta_n = \eta\psi a^{-\frac{1}{\gamma}} |\omega_n| J_{n+\frac{1}{2}+a^{-1}}(a^{-1}), \quad (2)$$

where $\eta \approx 19.20$, $J_\alpha(x)$ is the Bessel function of the first kind, and the complex number ψ is the Landau order parameter. Note that Eq. (1) appeared in an earlier paper [10].

The free energy expanded to the fourth power of $|\psi|$ near the transition (the Landau free energy) is

$$f_L = Rv_0g^2 \left[\frac{2\pi\gamma}{g} (T - T_c) |\psi|^2 + \frac{|\psi|^4}{2} \right], \quad (3)$$

where $R \approx 0.62$ and v_0 is the density of states at the Fermi level. This allows us to evaluate $|\psi|$ near T_c ,

$$|\psi| = \left[\frac{2\pi\gamma}{g} \right]^{1/2} (T_c - T)^{1/2}, \quad (4)$$

the jump in the specific heat at T_c ,

$$\delta c = c_s - c_n = 2\pi Rv_0g^2 a^{\frac{1}{\gamma}}, \quad (5)$$

and the thermodynamic critical field

$$H_c = 4\pi^{3/2}\gamma\sqrt{Rv_0}(T_c - T). \quad (6)$$

As expected $|\psi|$ and H_c have the usual mean-field critical exponents, but note also their γ dependence.

We prove that the global minimum of the free energy is unique for $T \rightarrow 0$ up to an overall phase $e^{i\phi}$ of the gap function $\Delta(\omega_n)$ and determine the leading small T asymptotic behavior of $\Delta(\omega_n)$. Specifically,

$$\frac{\Delta(\omega_n)}{\omega_*} = Y\left(\frac{\omega_n}{\omega_*}\right), \quad Y(x) = x \tan \theta_0(x), \quad (7)$$

where

$$\omega_* = g \frac{[\zeta(\gamma - 2)]^{\frac{1}{3}}}{2^{\frac{1}{3}}} \left(\frac{g}{2\pi T} \right)^{\frac{\gamma}{3}-1} \quad (8)$$

and $\theta_0(x)$ is the solution of the universal gap equation $\theta'' = x \sin \theta$. Equation (7) says that plots of the gap functions $\Delta(\omega_n)$ vs ω_n collapse onto the same curve (Fig. 4 below) for all $\gamma > 3$, g , and $T \rightarrow 0$, when both $\Delta(\omega_n)$ and ω_n are measured in units of ω_* . Further, we evaluate $2\Delta(0)/T_c$ and find that it is finite for all $\gamma > 3$ and $T > 0$, but diverges as $|\gamma - 3|^{-1/3}$ for $\gamma \rightarrow 3^+$ and as $T^{1-\gamma/3}$ for $T \rightarrow 0$.

We also determine the normal state specific heat c_n at all temperatures and the specific heat c_s in the superconducting state just below T_c and for $T \rightarrow 0$. The analysis of the specific heat reveals a pathology in the γ model: c_n is negative just above T_c and c_s is negative at $T \rightarrow 0$. This indicates that the γ model is thermodynamically unstable [11] at these temperatures. The resolution of this problem depends on the microscopic Hamiltonian underlying the γ model, which is not given. We discuss several scenarios. In particular, there is a scenario that removes this pathology without affecting any of the above answers.

The content of this paper is as follows. In Sec. II, we derive the free energy functional of the γ model and map it to a classical Heisenberg spin chain using the approach we proposed in an earlier paper. In Sec. III, we obtain the equation for the stationary points of the effective action for arbitrary γ and in Sec. IV we specialize the free energy to the extreme retardation regime. In Sec. V, we first devise a numerical approach for determining T_c for arbitrary γ and then obtain the solution of the linearized gap equation for $\gamma \rightarrow \infty$. We derive the large γ asymptote of T_c in Sec. VI. Section VII contains our results for the Landau free energy, jump in the specific heat at T_c , Landau order parameter, and thermodynamic critical field. We work out the specific heat in the normal state and in the superconducting state just below T_c in Sec. VIII. In Sec. IX, we solve the gap equation at low temperatures, and evaluate the specific heat in the superconducting state at $T \rightarrow 0$. In the final section, we summarize and pose an important open problem. The Appendix provides a proof of the uniqueness of the global minimum of the free energy at small T .

II. MAPPING TO THE SPIN CHAIN

In previous paper [9] we mapped the effective action for electrons interacting via phonons to a classical spin chain. In the strong coupling limit where phonon frequencies go to zero, this is the γ model with $\gamma = 2$. The same approach works for arbitrary γ . We define the γ model by its Euclidean action

$$S = T \sum_{\mathbf{p}\sigma} \psi_{\mathbf{p}\sigma}^* G_{\mathbf{p}}^{-1} \psi_{\mathbf{p}\sigma} - \frac{T^3 \delta}{2} \sum_{\mathbf{p}_i \sigma'} \frac{g^\gamma}{|\omega_l|^\gamma + \Omega^\gamma} \psi_{\mathbf{p}_1 \sigma}^* \psi_{\mathbf{p}_2 \sigma} \psi_{\mathbf{p}_3 \sigma'}^* \psi_{\mathbf{p}_4 \sigma'}, \quad (9)$$

where $G_{\mathbf{p}}^{-1} = -i\omega_n + \xi_{\mathbf{p}}$, ω_n are fermionic Matsubara frequencies, $\xi_{\mathbf{p}}$ are single-particle levels counted from the chemical potential, $\psi_{\mathbf{p}\sigma}$ is the fermionic field, and \mathbf{p}_i are frequency-momentum 4-vectors constrained by $\mathbf{p}_1 - \mathbf{p}_3 = \mathbf{p}_2 - \mathbf{p}_4 = (\omega_l, \mathbf{q})$. The single-particle level spacing δ is a combination of the density of states at the Fermi energy v_0 and the system volume V , $\delta = (v_0 V)^{-1}$. We added a constant $\Omega \ll T$ to avoid divergence at zero frequency. In the end we take the limit $\Omega \rightarrow 0$.

This action is exact for the electron-phonon system ($\gamma = 2$) in the limit when phonon frequencies go to zero, the Fermi energy $\varepsilon_F \rightarrow \infty$, and assuming the electron-phonon coupling depends only on the magnitude of the transferred momentum. Its status is unclear in most other examples listed above, when the bosonic field is a collective mode containing fermions themselves, such as, e.g., the nematic or magnetic order parameter. In this paper, we simply postulate Eq. (9) and study its properties regardless of its origin.

The mapping to the spin chain involves several steps [9]. First, we decouple the interaction with three Hubbard-Stratonovich fields Σ_\uparrow , Σ_\downarrow , and Φ and integrate out the fermions [12]. Stationary point equations for these fields are the Eliashberg equations generalized to arbitrary γ from $\gamma = 2$. Fluctuations around the stationary point are negligible as long as the Fermi energy is the largest energy in the problem

[13]. To unveil the spin chain, we expand the action in Ω and then send $\Omega \rightarrow 0$ [14]. A key ingredient of the mapping is the observation that at the stationary points the normal and anomalous Green's functions integrated over $\xi_{\mathbf{p}}$ satisfy the constraint $G_n^2 + |F_n|^2 = 1$. The normal average G_n is real and the anomalous average F_n is complex. Moreover, the mass of fluctuations violating this constraint is proportional to $\Omega^{-\gamma} \rightarrow \infty$ [13]. The constraint is therefore rigidly enforced even away from the stationary point.

It is now natural to introduce classical spins \mathbf{S}_n of unit length as

$$S_n^z = G_n, \quad S_n^x = \text{Re}(F_n), \quad S_n^y = \text{Im}(F_n). \quad (10)$$

In terms of the spins, the action expanded in Ω becomes $S_{\text{eff}} = \nu_0 V H_s$, where H_s is the spin chain Hamiltonian

$$H_s = -2\pi \sum_n \omega_n S_n^z - \pi^2 T g^\gamma \sum_{nm} \frac{\mathbf{S}_n \cdot \mathbf{S}_m - 1}{|\omega_n - \omega_m|^\gamma}. \quad (11)$$

In addition to the expansion in Ω , we regularized the interaction term [9] by replacing $\mathbf{S}_n \cdot \mathbf{S}_m \rightarrow \mathbf{S}_n \cdot \mathbf{S}_m - 1$ and then took the limit $\Omega \rightarrow 0$, which is now harmless. The free energy density of a given field configuration is [15]

$$f = \frac{TS_{\text{eff}}}{V} = \nu_0 T H_s. \quad (12)$$

The Boltzmann weight of a given field configuration is $e^{-fV/T} = e^{-H_s/\delta}$, i.e., the spin chain is at an effective temperature $T_s = \delta$. We work in the thermodynamic limit where $\delta \rightarrow 0$ and therefore the spin chain is at zero temperature.

We introduce the frequency-dependent superconducting gap function $\Delta_n \equiv \Delta(\omega_n)$ through [9,16]

$$S_n^z = \frac{\omega_n}{\sqrt{\omega_n^2 + |\Delta_n|^2}}, \quad S_n^+ = \frac{\Delta_n}{\sqrt{\omega_n^2 + |\Delta_n|^2}}, \quad (13)$$

where $S_n^+ = S_n^x + iS_n^y$.

In the normal state, $\Delta_n = 0$ and as a consequence $\mathbf{S}_n = \text{sgn}(\omega_n)\hat{z}$, where \hat{z} is the unit vector along the z axis. A characteristic feature of this state is a sharp domain wall between ω_{-1} and ω_0 . In the superconducting state, $\Delta_n \neq 0$ and the domain wall softens, see Fig. 1.

III. STATIONARY POINTS OF THE FREE ENERGY

In Ref. [9], we discussed stationary points of the free energy functional f for $\gamma = 2$. These results in fact apply to any $\gamma > 1$. In particular, we know that with a proper choice of the arbitrary overall phase $e^{i\phi}$, the gap function at the minimum of the free energy has the following properties: $\Delta(\omega_n) > 0$ for all ω_n or $\Delta(\omega_n) = 0$ for all ω_n , $\Delta(\omega_n)$ is an even function, and $\Delta(\omega_n) \rightarrow 0$ as $\omega_n \rightarrow \pm\infty$. These properties hold for $T = 0$ as well, in which case ω_n becomes a continuous variable ω and $\Delta(\omega_n) > 0$ implies $\Delta(0) > 0$. We also know that the stationary points of f and of the spin chain H_s are identical, because we are in the strong coupling limit $\lambda = \frac{g^\gamma}{\Omega^\gamma} \rightarrow \infty$ (equivalent to $\Omega \rightarrow 0$).

Given that $\Delta_n \geq 0$, Eq. (10) implies $S_n^y = 0$ and $S_n^x \geq 0$. Let $\theta_n \equiv \theta(\omega_n)$ be the angle \mathbf{S}_n makes with the z axis. Then

$$S_n^z = \cos \theta_n, \quad S_n^x = \sin \theta_n, \quad 0 \leq \theta_n \leq \pi. \quad (14)$$

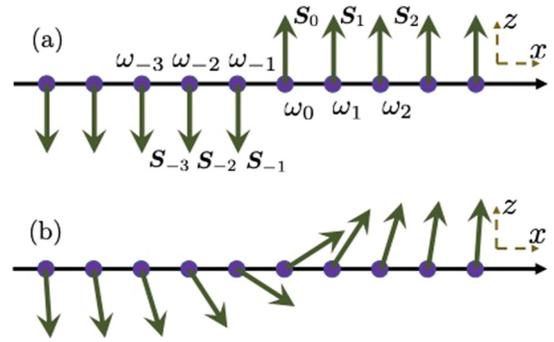


FIG. 1. Transition from (a) the normal state to (b) a superconducting state in terms of classical spins. Positions of the spins \mathbf{S}_n are the fermionic Matsubara frequencies ω_n . Spin-spin interactions are purely ferromagnetic and the spins are subject to a Zeeman magnetic field $2\pi\omega_n$ along the z -axis. In the superconducting state, spins acquire x components, which implies nonzero anomalous Green's function. The normal-superconductor transition is seen as the softening of the sharp domain wall at the origin of the spin chain.

The angle θ_n cannot exceed π due to the condition $S_n^x \geq 0$. This also prevents θ_n from winding. The remaining properties of the gap function

$$\Delta_n = \omega_n \tan \theta_n \quad (15)$$

translate into

$$\theta(-\omega_n) = \pi - \theta(\omega_n), \quad \lim_{\omega_n \rightarrow \pm\infty} \theta_n = \frac{\pi}{2} \mp \frac{\pi}{2}. \quad (16)$$

The spin chain Hamiltonian reads in terms of θ_n

$$H_s = -2\pi \sum_n \omega_n \cos \theta_n - \pi^2 T g^\gamma \sum_{nm} \frac{\cos(\theta_n - \theta_m) - 1}{|\omega_n - \omega_m|^\gamma}.$$

Differentiating this with respect to θ_n , we obtain the equation for the stationary points of the free energy of the γ model

$$\omega_n \sin \theta_n = g^\gamma \pi T \sum_{m \neq n} \frac{\sin(\theta_m - \theta_n)}{|\omega_m - \omega_n|^\gamma}. \quad (17)$$

This is the superconducting *gap equation*. It is a version of the Eliashberg gap equation as it follows from the generalized Eliashberg equations mentioned in the previous section. It takes the form of the standard gap equation for the strong coupling limit of the Eliashberg theory [7,8] when we rewrite it in terms of Δ_n and set $\gamma = 2$.

The *normal state*, $\Delta_n = 0$, is always a stationary point. The corresponding spin configuration $\mathbf{S}_n = \text{sgn}(\omega_n)\hat{z}$ is shown in Fig. 1(a). In this state $\theta_n = 0$ for $\omega_n > 0$ and $\theta_n = \pi$ for $\omega_n < 0$, which indeed satisfies the stationary point equation (17). It is also not difficult to show that since S_n^z is the normal Green's function integrated over $\xi_{\mathbf{p}}$, in the normal state all states below the chemical potential are occupied and above—empty. At $T = 0$, the Matsubara frequency $\omega_n \rightarrow \omega$ takes values on the entire real axis. The angle $\theta(\omega)$ is then discontinuous at $\omega = 0$ in the normal state. In contrast, in the ground state (free energy minimum at $T = 0$) we have

$$\theta(0) = \frac{\pi}{2}. \quad (18)$$

This follows from Eqs. (13) and (14) together with the condition $\Delta(0) > 0$.

The symmetry (16) enables us to express the free energy density (12) in terms of $\theta_{n \geq 0}$ only. It is also convenient to introduce along the way the nondimensionalized free energy density \bar{f} as follows:

$$\bar{f} \equiv \frac{f}{v_0 g^2} = -2\tau \sum_{n=0}^{\infty} \bar{\omega}_n \cos \theta_n - \frac{\tau^{2-\gamma}}{2} \sum_{n,m=0}^{\infty} \times \left[\frac{\cos(\theta_n - \theta_m) - 1}{(n-m)^\gamma} - \frac{\cos(\theta_n + \theta_m) + 1}{(n+m+1)^\gamma} \right], \quad (19)$$

where the dimensionless temperature τ and fermionic Matsubara frequency $\bar{\omega}_n$ are

$$\tau = \frac{2\pi T}{g}, \quad \bar{\omega}_n = \frac{\omega_n}{g} = \tau \left(n + \frac{1}{2} \right) \quad (20)$$

IV. EXTREME RETARDATION REGIME

In this paper we are primarily interested in the extreme retardation regime of the γ model. In this regime, the range of the interaction is extremely short in the frequency domain and therefore extremely long (retarded) in the time domain. In terms of the spin chain (11), this means that a given spin S_n interacts only with its nearest neighbors or, in the continuum $T \rightarrow 0$ limit, only with spins in its immediate vicinity. There are two overlapping cases when this happens. One is $\gamma \gg 1$ at arbitrary T . The other is low T for any $\gamma > 3$. Let us consider these two cases.

A. Large γ , arbitrary T

At large γ , the $|n-m| \geq 2$ interaction terms in Eq. (11) are suppressed by a factor of $2^{-\gamma}$ or smaller compared to the $|n-m|=1$ terms. Discarding exponentially small terms, we obtain a spin chain with nearest-neighbor interactions,

$$H_s = -2\pi \sum_n \omega_n S_n^z - \frac{\pi g^\gamma}{(2\pi T)^{\gamma-1}} \sum_n (S_n \cdot S_{n+1} - 1). \quad (21)$$

The corresponding nondimensionalized free energy density for the coplanar spin distribution (14) is

$$\bar{f} \equiv \frac{f}{v_0 g^2} = -\tau \sum_n \bar{\omega}_n \cos \theta_n - \frac{\tau^{2-\gamma}}{2} \sum_n [\cos(\theta_{n+1} - \theta_n) - 1], \quad (22)$$

and the expression (19) in terms of $\theta_{n \geq 0}$ becomes

$$\bar{f} = -2\tau \sum_{n \geq 0} \bar{\omega}_n \cos \theta_n + \tau^{2-\gamma} \frac{\cos(2\theta_0) + 1}{2} - \tau^{2-\gamma} \sum_{n \geq 0} [\cos(\theta_{n+1} - \theta_n) - 1]. \quad (23)$$

We use this expression for the free energy in subsequent sections to determine the critical temperature T_c , the jump in the specific heat at the transition, etc.

B. Any $\gamma > 3$, low T

We will see below that at $T \ll g$ the γ model is local for any $\gamma > 3$. The gap equation becomes a nonlinear ordinary differential equation (ODE), and the interaction part of the free energy corresponds to the continuum limit of the classical ferromagnetic Heisenberg spin chain with nearest-neighbor interactions. Again, the interaction is as local in the frequency domain as it can be.

Consider the solution $\theta_n = \theta(\omega_n)$ of the gap equation (17) under the conditions (16). One can always extend $\theta(\omega_n)$ to an infinitely differentiable function $\theta(\omega)$ for all real ω that coincides with θ_n when $\omega = \omega_n$, and such that conditions (16) hold with ω in place of ω_n . This also extends $\Delta(\omega_n)$ to $\Delta(\omega)$ with the help of Eq. (15) and fixes $\theta(0) = \frac{\pi}{2}$ [17] by continuity of $\theta(\omega)$. There are uncountably many such functions but we will determine a distinguished $\theta(\omega)$ that is real analytic on the entire real ω axis [18] and captures the small T asymptotic behavior of $\theta(\omega_n)$. We will see in Sec. IX that $\theta(\omega)$ varies on a scale $\omega_* \propto T^{1-\frac{\gamma}{3}}$. The difference $|\omega_m - \omega_n|$ in Eq. (17) is of the order of ω_* when $|m-n| \sim (T/g)^{-\frac{\gamma}{3}} \gg 1$. Terms with $|m-n| \gg 1$ are negligible due to the rapid convergence of the sum over m . Therefore, $\theta(\omega_m)$ is close to $\theta(\omega_n)$ for m whose contribution to Eq. (17) is significant and

$$\begin{aligned} \sin(\theta_m - \theta_n) &\approx \theta_m - \theta_n \\ &\approx \frac{d\theta_n}{d\omega_n} (\omega_m - \omega_n) + \frac{1}{2} \frac{d^2\theta_n}{d\omega_n^2} (\omega_m - \omega_n)^2. \end{aligned} \quad (24)$$

Substituting Eq. (24) into the gap equation (17) and dropping the discrete index n [$\omega_n \rightarrow \omega$ and $\theta_n \rightarrow \theta(\omega)$], we find the nonlinear ODE

$$\bar{\omega} \sin \theta = \frac{\tau^{3-\gamma}}{2} \frac{d^2\theta}{d\bar{\omega}^2} \sum_{k=1}^{\infty} \frac{1}{k^{\gamma-2}}, \quad (25)$$

where $\bar{\omega} = \omega/g$. We see that the approximations made require $\gamma > 3$ and low T so that the summation over k converges and the replacement (24) is accurate. More precisely, in Sec. IX we show that the local approximation (25) for the gap equation and the corresponding approximation for the free energy are accurate when $\tau^{\frac{r}{3}} \ll 1$, where $r = \gamma - 3$ for $3 < \gamma \leq 5$ and $r = 2$ for $\gamma \geq 5$. We use Eq. (25) in Sec. IX to determine the scaling laws at low T .

The nondimensionalized free energy density in this approximation becomes

$$\bar{f} = -\tau \sum_n \bar{\omega}_n \cos \theta_n + \zeta(\gamma - 2) \frac{\tau^{4-\gamma}}{4} \sum_n \left(\frac{d\theta_n}{d\bar{\omega}} \right)^2, \quad (26)$$

where $\zeta(x)$ is the Riemann zeta function. The variation of this expression with respect to θ_n gives Eq. (25). It is also instructive to compare formulas (26) and (22) for the free energy. The low temperature limit of Eq. (22) and the large γ limit of Eq. (26) should coincide. This is indeed the case, because $1 - \cos(\theta_{n+1} - \theta_n) \approx \frac{1}{2} \tau^2 (d\theta_n/d\bar{\omega}_n)^2$ at low T and $\zeta(\gamma - 2) \rightarrow 1$ as $\gamma \rightarrow \infty$.

In terms of spins the free energy (26) takes a particularly simple form,

$$\bar{f} = \kappa^{-2} \int_{-\infty}^{\infty} dx \left\{ -xS^z + \frac{1}{2} \left(\frac{dS}{dx} \right)^2 \right\}, \quad (27)$$

where

$$\kappa^3 = \frac{2\tau^{\gamma-3}}{\zeta(\gamma-2)}, \quad x = \kappa\bar{\omega}. \quad (28)$$

This is the continuum limit of the classical ferromagnetic Heisenberg spin chain with nearest-neighbor interactions and a Zeeman field. Note that the free energy diverges as $T^{2-\frac{2\gamma}{3}}$ in the $T \rightarrow 0$ limit. This divergence is cutoff by the mass Ω of the critical boson. In other words, the γ model requires an infrared cutoff for $\gamma > 3$. Physically, we are working in the regime $T \gg \Omega > 0$ and by low T we mean $g \gg T \gg \Omega$.

V. SUPERCONDUCTING TRANSITION

Here we first investigate the superconducting transition in the γ model for *arbitrary* γ and then turn to $\gamma \rightarrow \infty$ limit. As mentioned in Sec. III, the normal state is a stationary point of the free energy functional at any temperature. It is the global minimum above a certain (critical) temperature T_c . Below T_c , which is nonzero for all γ , it is a saddle point implying the emergence of a new (superconducting) global minimum at $T = T_c$.

Interestingly, there is a single unstable direction (normal mode) at any $T < T_c$ for $\gamma \gtrsim 2.7$ and for $\gamma \lesssim 0.2$. In other words, there is an orthonormal basis in the configuration space of the free energy, such that it increases along all axes except one, along which it decreases. Therefore, a single mode is responsible for the Cooper instability at all T in the γ model with these values of γ . The same is true in the BCS theory [19,20], which corresponds to $\gamma = 0$. It is an amplitude (Higgs) mode, because the phase of Δ_n is zero for it.

Coherent dynamics of the BCS condensate in response to sudden perturbations can be understood as undamped, underdamped, and overdamped oscillations of the Higgs mode [20–22]. At the first glance, this suggests that the γ model for $\gamma \gtrsim 2.7$ and $\gamma \lesssim 0.2$ has the same three nonequilibrium phases as the BCS condensate. In Phase I the order parameter decays to zero at long times, in Phase II it goes to a nonzero constant, and in Phase III it oscillates persistently and periodically [23]. However, we have to keep in mind that the normal modes we obtain for the γ model are the modes of the free energy functional and not of the Hamiltonian dynamics as in the BCS case. Because of this it is far from certain that the far from equilibrium dynamics of the two systems are indeed similar.

A. Arbitrary γ

As usual, to determine the type of the stationary point, we need to expand the free energy to quadratic order around it. It is convenient to work with the free energy (19) formulated on the positive Matsubara axis. Expanding Eq. (19) around the

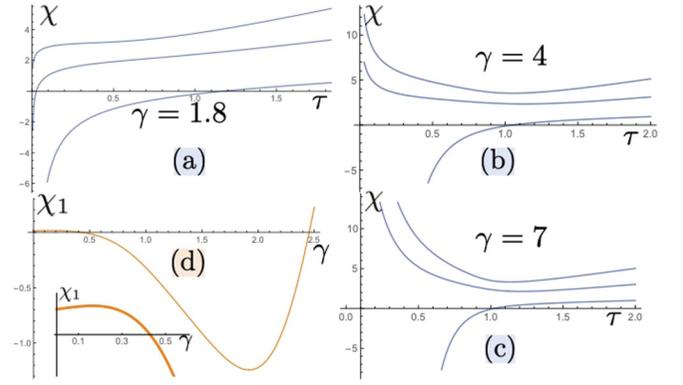


FIG. 2. Three lowest eigenvalues of the second derivative matrix X vs reduced temperature τ for three different values of γ [(a), (b), (c)] and the second lowest eigenvalue χ_1 as a function of γ at $\tau = 0.01$ (d). The inset in (d) magnifies the small γ region of the graph to show that $\chi_1(\gamma)$ becomes positive again at small γ . Eigenvalues of X determine the stability of the normal state—it is stable when all of them are positive and unstable otherwise. For any γ there is a critical temperature τ_c below which the normal state is unstable. When $0.2 \lesssim \gamma \lesssim 2.7$ two eigenvalues are negative at $\tau \rightarrow 0$. Otherwise, there is a single negative eigenvalue at all $\tau < \tau_c$.

normal state, we obtain a quadratic form

$$\delta\bar{f} = \tau \sum_{n,m=0}^{\infty} \theta_n X_{nm} \theta_m = \tau \boldsymbol{\theta}^T X \boldsymbol{\theta}, \quad (29)$$

where $\delta\bar{f} = \bar{f} - \bar{f}_n$, \bar{f}_n is the nondimensionalized normal state free energy, $\boldsymbol{\theta}$ is a column vector with components θ_n , and the second derivative matrix X (the Hessian) reads

$$\begin{aligned} \tau^{\gamma-1} X_{nn} &= \tau^{\gamma} \left(n + \frac{1}{2} \right) + \sum_{k=1}^n \frac{1}{k^{\gamma}} - \frac{1}{2(2n+1)^{\gamma}}, \\ \tau^{\gamma-1} X_{n \neq m} &= -\frac{1}{2(n+m+1)^{\gamma}} - \frac{1}{2|n-m|^{\gamma}}. \end{aligned} \quad (30)$$

The normal state is a minimum when all eigenvalues of X are positive. It becomes a saddle point and a transition to the superconducting state takes place when one of the eigenvalues of X vanishes. In a mechanical interpretation, where $\delta\bar{f}$ is the potential energy of a system of 1D point particles with coordinates θ_n and equal mass, eigenvalues and eigenvectors of X have the meaning of squares of the frequencies and normal modes of small oscillations around the normal state, respectively. The normal state is unstable when one of the frequencies becomes imaginary.

To analyze the eigenvalues of X numerically, we truncate it to a finite $L \times L$ matrix. The size L should be such that the Zeeman field overwhelms the interaction near the ends of the chain as it does at large Matsubara frequencies in an infinite chain. The first line in Eq. (30) shows that this requires $L \gg \tau^{-\gamma}$ for $\gamma > 1$ and $(L\tau)^{\gamma} \gg 1$ for $\gamma < 1$. For comparison, the critical dimensionless temperature is $\tau_c \approx 1$ for $\gamma \geq 2$ (see below). It increases with decreasing γ and diverges, $\tau_c \sim \gamma^{-\frac{1}{\gamma}}$ as $\gamma \rightarrow 0^+$ [10].

Figure 2 shows three lowest eigenvalues $\chi_0 < \chi_1 < \chi_2$ of X as functions of the reduced temperature τ for several γ and

the second eigenvalue χ_2 as a function of γ at a fixed temperature $\tau = 0.01 < \tau_c$. We observe that: (a) χ_0 is negative below a certain value of τ and positive above it for all γ , i.e., there is a superconducting transition for any γ ; (b) χ_2 is positive at all γ and temperatures; and (c) χ_1 changes sign twice as a function of γ at low temperatures.

Point (c) comes as a surprise—the second eigenvalue χ_1 also becomes negative at very low temperatures for $0.2 \lesssim \gamma \lesssim 2.7$. Outside of this interval of γ only one eigenvalue is negative for all $T < T_c$. It is interesting to see whether this means that the dynamics of the Cooper instability for $\gamma \gtrsim 2.7$ and $\gamma \lesssim 0.2$ are as in the BCS model, where there is a single negative eigenvalue as well. Suppose we prepare the system in the normal state and then suddenly turn on the interaction (interaction quench). The BCS order parameter first grows exponentially with the growth exponent set by the negative eigenvalue [19,24]. Nonlinear effects stop the growth at some point after which the order parameter oscillates persistently and periodically. The fact that oscillations occur with a single basic frequency is a consequence of having only one unstable direction.

B. Large γ

The problem of finding the large γ asymptotic behavior of T_c as well as of the eigenvalues and eigenvectors of X is exactly solvable. First, it is not difficult to see that at $\gamma = \infty$ the critical temperature $\tau_c = 1$ ($T_c = \frac{g}{2\pi}$). Indeed, the interaction term in Eq. (22) is proportional to $\tau^{2-\gamma}$. At $\gamma = \infty$ the interaction vanishes for $\tau > 1$. Then, the system simply minimizes the Zeeman term as shown in Fig. 1(a), i.e., we are in the normal state, where $\theta_n = \frac{\pi}{2} - \frac{\pi}{2} \text{sgn}(\omega_n)$. For $\tau < 1$ the interaction diverges. The sharp domain wall at the origin (the jump from $\theta_{-1} = \pi$ to $\theta_0 = 0$) now costs infinite energy as the interaction is ferromagnetic and favors parallel spin alignment. It is more advantageous to spread the jump in θ_n over a large energy interval, i.e., the domain wall softens [25]. In fact, in this limit all spins at finite ω_n are along the x -axis and $\vec{f} = 0$ for $\tau < 1$ [26].

Now let us evaluate the leading order asymptotic behavior of T_c and other observables for large γ . Near the transition the system is close to the normal state. We therefore expand the free energy (23) around the normal state to the second order,

$$\delta \bar{f} = \tau \sum_{n=0}^{\infty} \bar{\omega}_n \theta_n^2 + \frac{\tau^{2-\gamma}}{2} \sum_{n=0}^{\infty} (\theta_{n+1} - \theta_n)^2 - \tau^{2-\gamma} \theta_0^2. \quad (31)$$

The notations here the same as in Eq. (29).

Finding eigenvalues of the Hessian X is equivalent to finding the stationary points of $\bar{f}_\chi = \bar{f} - \chi \tau \sum_n \theta_n^2$. Setting the derivative of \bar{f}_χ with respect to θ_n to zero, we obtain one equation for $n \geq 1$,

$$\theta_{n+1} = 2[\tau^\gamma (n + \frac{1}{2}) - \tau^{\gamma-1} \chi + 1] \theta_n - \theta_{n-1}, \quad (32)$$

and another one for $n = 0$,

$$\theta_1 = [\tau^\gamma - 1 - 2\tau^{\gamma-1} \chi] \theta_0. \quad (33)$$

Similar to a second-order linear differential equation, the recurrence relation (32) has two linearly independent solutions.

TABLE I. Values of the superconducting transition temperature $T_c(\gamma)$ for various γ versus the large γ asymptote (39). The agreement is reasonable already for $\gamma = 2$. The relative error—the last column of the table—is roughly $\frac{2-\gamma}{2\gamma}$ consistent with $O(2^{-\gamma})$ term in Eq. (39).

γ	$\frac{T_c(\gamma)}{g}$	$\frac{T_c(\gamma \rightarrow \infty)}{g}$	$\frac{T_c(\gamma) - T_c(\gamma \rightarrow \infty)}{T_c(\gamma \rightarrow \infty)}$
2	0.183	0.173	6%
3	0.171	0.168	2%
4	0.1669	0.1660	0.5%
8	0.16258	0.16256	0.01%
20	0.160506897	0.160506899	$-10^{-6}\%$

Equation (33) provides one boundary condition. The second boundary condition is $\theta_n \rightarrow 0$ for $n \rightarrow \infty$, see Eq. (16).

Observe that Eq. (32) is the recurrence relation for Bessel functions,

$$Z_{\alpha+1}(x) = \frac{2\alpha}{x} Z_\alpha - Z_{\alpha-1}, \quad (34)$$

with $x = \tau^{-\gamma}$ and $\alpha = n + \frac{1}{2} - \tau^{-1} \chi + \tau^{-\gamma}$. The Bessel function $Z_\alpha(x)$ that goes to zero as $\alpha \rightarrow \infty$ is $J_\alpha(x)$ —the Bessel function of the first kind. Therefore, the normal mode with eigenvalue χ is

$$\theta_{n\chi} = J_{n+\frac{1}{2}+\tau^{-\gamma}-\chi\tau^{-1}}(\tau^{-\gamma}), \quad (35)$$

up to a normalization constant. The boundary condition (33) now determines the eigenvalues χ at temperature τ ,

$$\frac{J_{\frac{3}{2}+\tau^{-\gamma}-\chi\tau^{-1}}(\tau^{-\gamma})}{J_{\frac{1}{2}+\tau^{-\gamma}-\chi\tau^{-1}}(\tau^{-\gamma})} = \tau^\gamma - 1 - 2\tau^{\gamma-1} \chi. \quad (36)$$

We use this formula to determine the critical temperature and derive the Landau free energy near the critical point.

VI. CRITICAL TEMPERATURE

As discussed above, the superconducting transition temperature $\tau_c = \frac{2\pi T_c}{g}$ corresponds to one of the eigenvalues χ of the matrix X crossing zero. Setting $\chi = 0$ in Eq. (36), we obtain an equation for $a = \tau_c^\gamma$,

$$\frac{J_{\frac{3}{2}+a^{-1}}(a^{-1})}{J_{\frac{1}{2}+a^{-1}}(a^{-1})} = a - 1. \quad (37)$$

Numerically, we find that this equation has a unique solution

$$a \approx 1.1843, \quad (38)$$

and therefore

$$T_c(\gamma \rightarrow \infty) = [a + O(2^{-\gamma})]^\frac{1}{\gamma} \frac{g}{2\pi}. \quad (39)$$

The correction $O(2^{-\gamma})$ to a is due to interactions with non-nearest-neighbor spins, which we neglected. This formula gives the leading term in the asymptotic expansion of $T_c(\gamma)$ around $\gamma = \infty$. Note that $T_c(\infty) = \frac{g}{2\pi}$ in agreement with our reasoning at the beginning of Sec. VB.

In Table I, we compare the large γ asymptote (39) with numerically exact values of T_c for several $2 \leq \gamma \leq 20$. The agreement is reasonable already at $\gamma = 2$. To determine T_c numerically, we diagonalize the Hessian X in Eq. (30) as a

function of the reduced temperature $\tau = \frac{2\pi T}{g}$ and compute the value of τ at which the lowest eigenvalue of X crosses zero.

VII. THERMODYNAMICS NEAR THE TRANSITION

Near T_c we need to keep only the unstable mode $\theta_{n\chi_0}$ whose eigenvalue χ_0 changes sign at the transition. Then, $\theta_n = \epsilon\theta_{n\chi_0}$, where the amplitude of the unstable mode ϵ is our order parameter. The free energy \bar{f} expanded near the transition to order ϵ^4 is the Landau free energy from which various thermodynamic properties, such as the jump in the specific heat, thermodynamic critical field H_c etc., follow.

At the transition, the unstable mode [Eq. (35) with $\chi = 0$ and $\tau = \tau_c$] is

$$\theta_{n0} = J_{n+\frac{1}{2}+a^{-1}}(a^{-1}), \quad (40)$$

where we used $\tau_c^\gamma = a$. This function is positive for all $n \geq 0$. It decays with n monotonically and very quickly, approximately as n^{-n} . Its values at the first ($n=0$) and second ($n=1$) Matsubara frequencies contribute 96.66% and 3.28% to $\sum_{n=0}^{\infty} \theta_{n0}^2$. The superconductivity is therefore confined to few small Matsubara frequencies, mostly to $\pm\pi T$, consistent with the short ranginess of the interaction, cf. [10]. In terms of spins, the domain wall that develops below T_c (see Fig. 1) measures only a couple of sites long.

To obtain the Landau free energy, we expand Eq. (22) to quartic order in θ_n and substitute $\theta_n = \epsilon\theta_{n\chi_0}$. The quadratic part simplifies since $\theta_{n\chi_0}$ is the eigenstate of matrix X with eigenvalue χ_0 and we find

$$\delta\bar{f} = \chi_0\epsilon^2\tau_cb_2 + \epsilon^4\tau_c^2b_4. \quad (41)$$

As $\chi_0 \sim \epsilon^2$ at the minimum with respect to ϵ , it is sufficient to evaluate the coefficients at $\chi_0\epsilon^2$ and ϵ^4 at $\tau = \tau_c$ —corrections to them in $\delta\tau = \tau - \tau_c$ contribute terms of order ϵ^6 to the free energy. Then, $b_2 = \sum_n \theta_{n0}^2 \approx 6.074 \times 10^{-2}$ and similarly we determine $b_4 \approx 8.445 \times 10^{-4}$.

We also need χ_0 as a function of $\delta\tau$. Expanding Eq. (36) in χ and $\delta\tau$ around $\chi = 0$ and $\tau = \tau_c$, we find $\chi_0 = h\gamma\delta\tau$, where $h \approx 0.5339$. It is convenient to redefine the order parameter as $\tau_c\epsilon^2 = \eta^2|\psi|^2$, where $\eta^2 = \frac{hb_2}{2b_4}$. We have

$$\delta\bar{f} = R \left[\gamma\delta\tau|\psi|^2 + \frac{|\psi|^4}{2} \right], \quad R \approx 0.6226. \quad (42)$$

We made $\psi = |\psi|e^{i\phi}$ complex to restore the arbitrary overall phase $e^{i\phi}$ of Δ_n , which we set to one until now. Since $\theta_{n \geq 0}$ is small, Eqs. (13) and (14) imply $\Delta_n = \omega_n\theta_n$ for $\omega_n > 0$ and therefore

$$\Delta(\omega_n) = \eta|\psi|a^{-\frac{1}{\gamma}}\omega_n J_{n+\frac{1}{2}+a^{-1}}(a^{-1})e^{i\phi}, \quad \omega_n > 0,$$

$$\Delta(-\omega_n) = \Delta(\omega_n), \quad \eta \approx 19.20. \quad (43)$$

Note that $\Delta(\omega_n)e^{-i\phi} > 0$ in agreement with the theorem mentioned earlier that $\Delta(\omega_n)$ must be non-negative at the minimum of the free energy up to an overall phase.

Let us also restore the units in the Landau free energy (42),

$$f_L \equiv f - f_n = Rv_0g^2 \left[\frac{2\pi\gamma}{g}(T - T_c)|\psi|^2 + \frac{|\psi|^4}{2} \right], \quad (44)$$

where f_n is the normal state free energy density.

Minimizing the condensation energy (44) with respect to $|\psi|^2$, we find

$$|\psi| = \left[\frac{2\pi\gamma}{g} \right]^{1/2} (T_c - T)^{1/2}, \quad (45a)$$

$$f - f_n = -2\pi^2 Rv_0\gamma^2 (T - T_c)^2. \quad (45b)$$

From here the jump in the specific heat $c = -T \frac{\partial^2 f}{\partial T^2}$, i.e., the difference between superconducting and normal state specific heats at $T = T_c$ is

$$\delta c = c_s - c_n = 2\pi Rv_0g\gamma^2 a^{\frac{1}{\gamma}}, \quad (46)$$

where we used Eq. (39). For $\gamma = 2$ this formula gives $\delta c \approx 24v_0g$, which is not too far from the exact answer (rounded to two significant digits) $\delta c = 17v_0g$ [7] considering that $\gamma = 2$ is well outside of $\gamma > 3$ range where our large γ theory is supposed to be accurate.

The thermodynamic critical field H_c is the magnetic field above which the energy cost $\frac{H^2}{8\pi}$ of expelling the magnetic field (Meissner effect) exceeds the energy gain (45b) due to superconductivity [27]. We find

$$H_c = 4\pi^{3/2}\gamma\sqrt{Rv_0}(T_c - T). \quad (47)$$

Both $|\psi|$ and H_c have the usual mean-field scaling with $(T_c - T)$ for a scalar theory.

VIII. HEAT CAPACITY ABOVE AND BELOW T_c

Here we evaluate the normal state specific heat c_n for $\gamma > 2$ and the specific heat c_s in the superconducting state just below T_c . It turns out that c_n is negative in a temperature interval (T_c, T_n) above T_c for any $\gamma \geq 2$. We will discuss the significance of this later in this section.

We saw above that in the normal state $S_n^z = \text{sgn}(\omega_n)$. The free energy density in terms of the spin chain is $f = v_0TH_s$. The Zeeman term in Eq. (21) contributes

$$f_n^Z = -2\pi v_0T \sum_{n=-\infty}^{\infty} |\omega_n| = -8\pi^2 v_0T^2 \sum_{n=0}^{\infty} \left(n + \frac{1}{2} \right) \quad (48)$$

to the normal state free energy. Looking back at the derivation of the spin chain [9], we see that up to a T -independent constant, f_n^Z must be the $\epsilon_F \rightarrow \infty$ limit of the free energy of the noninteracting Fermi gas [11],

$$f_n^{(0)} = u_0\epsilon_F - \frac{1}{3}\pi^2 v_0T^2, \quad (49)$$

where u_0 is a constant that depends only on the number of spatial dimensions, e.g., $u_0 = 3/5$ in 3D.

Summations in Eq. (48) diverge. The reason is that we took the limit $\epsilon_F \rightarrow \infty$ when deriving the spin chain and the Fermi gas free energy (49) diverges in this limit. However, this affects only the T -independent part, which is of no interest to us here. The standard way to deal with this divergence is to apply the Poisson summation formula to Eq. (48) discarding the T -independent part. We obtain

$$f_n^Z = -\frac{1}{3}\pi^2 v_0T^2. \quad (50)$$

It is instructive to also derive this answer using the zeta function regularization technique [28]. Recall the definition of the

Hurwitz zeta function

$$\zeta(s, p) = \sum_{n=0}^{\infty} \frac{1}{(n+p)^s}. \quad (51)$$

We interpret the second summation in Eq. (50) as $\zeta(-1, \frac{1}{2})$ and since $\zeta(-1, \frac{1}{2}) = \frac{1}{24}$, we obtain Eq. (50) with this approach too.

Now note that $S_n \cdot S_m - 1 = \text{sgn}(\omega_n \omega_m) - 1$ vanishes when ω_n and ω_m have the same sign and is equal to -2 otherwise. This observation allows us to rewrite the interaction part of the free energy as

$$\begin{aligned} f_{\text{int}} &= v_0 g^\gamma (2\pi T)^{2-\gamma} \sum_{l=1}^{\infty} \frac{l}{l^\gamma} \\ &= 4\pi^2 v_0 \zeta(\gamma-1) \left(\frac{g}{2\pi}\right)^\gamma T^{2-\gamma}. \end{aligned} \quad (52)$$

Here we reduced the summation over n and m to a single sum over $l = n + m + 1$ using $\omega_n - (-\omega_m) \propto (n + m + 1)$ and the fact that there are l ways to choose n and m for a given l .

Adding Eq. (52) to Eq. (50), we obtain the normal state free energy up to a T -independent constant

$$f_n = -\frac{1}{3}\pi^2 v_0 T^2 + 4\pi^2 v_0 \zeta(\gamma-1) \left(\frac{g}{2\pi}\right)^\gamma T^{2-\gamma}, \quad (53)$$

and

$$c_n(T) = \frac{2\pi^2 v_0}{3} T \left[1 - \left(\frac{T_n}{T}\right)^\gamma \right], \quad (54)$$

where

$$T_n = [3(\gamma-1)(\gamma-2)\zeta(\gamma-1)]^{\frac{1}{\gamma}} \frac{g}{2\pi} > T_c. \quad (55)$$

The inequality $T_n(\gamma) > T_c(\gamma)$ holds for all $\gamma > 2$ because the difference $T_n(\gamma) - T_c(\gamma \rightarrow \infty)$ grows with γ and exceeds $0.25T_c$ already at $\gamma \rightarrow 2^+$ [29], while $T_c(\gamma \rightarrow \infty)$ given by Eq. (39) underestimates T_c by 6% or less as seen from Table I.

Therefore, the heat capacity is negative for $T_c < T < T_n$. Note that $T_n \rightarrow T_c$ as $\gamma \rightarrow \infty$. Nevertheless, there is a sliver of temperature where the heat capacity is negative. The interpretation of this depends on the origin of the γ model. Any subsystem that does not interact with other subsystems and has negative heat capacity is thermodynamically unstable [11]. One scenario is that the quasiparticles are ill defined, they do not form a Fermi liquid, and therefore the stationary point we started with (the solution of the gap equation) is not the global minimum of the total free energy. Another scenario is that the effective fermion-fermion interaction $V(\omega_l)$ changes with temperature, so that $V(\omega_l) = \frac{g^\gamma}{|\omega_l|^\gamma}$ below T_c and something else above T_c . In other words, the γ model kicks in only below T_c . This scenario is in principle possible when the bosons that mediate the interaction are collective excitations of the fermions themselves as is the case for many $\gamma < 1$. Then, the superconducting transition modifies the bosonic propagator and therefore $V(\omega_l)$. However, if the interaction is mediated by phonons or other true bosons, the γ model is unphysical for $\gamma \geq 2$ at least in a certain range of temperatures, see also Ref. [30] where we addressed this issue for phonon mediated electron-electron interactions ($\gamma = 2$).

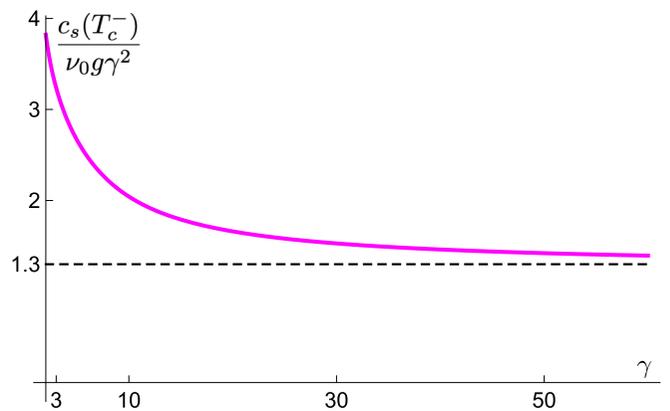


FIG. 3. Specific heat c_s in the superconducting state just below T_c as a function of γ , normalized by $v_0 g \gamma^2$. We see that c_s is positive at $T = T_c^-$ for all $\gamma > 3$. Since $c_s(T)$ is continuous, it will remain positive for a finite temperature range below T_c . Note also that $c_s(T_c^-) \approx 1.3 v_0 g \gamma^2$ at large γ .

Due to the jump at T_c , the specific heat becomes positive in the superconducting state for a range of temperatures $T_- < T \leq T_c$. We encountered this situation before in the $\gamma = 2$ case, where the superconducting state was free of the pathologies of the normal state—the opening of the gap stabilized the system. To evaluate the specific heat c_s just below T_c , i.e., at $T = T_c^-$, we use Eqs. (46) and (54) and

$$c_s(T_c^-) = c_n(T_c^-) + \delta c. \quad (56)$$

We plot $c_s(T_c^-)$ normalized by $v_0 g \gamma^2$ as a function of γ in Fig. 3. We see that it is positive at $T = T_c^-$ for all $\gamma > 3$. By continuity it must also remain positive in a certain finite temperature range $(T_-, T_c^-]$. Note also the large γ asymptote $c_s(T_c^-) \approx 1.3 v_0 g \gamma^2$. We will see in Sec. IX B that $T_- > 0$, i.e., c_s becomes negative at low temperatures. This is unlike the $\gamma = 2$ case where the specific heat is always positive in the superconducting state and vanishes when $T \rightarrow 0$ [30] as it should.

IX. LOW-TEMPERATURE PROPERTIES—UNIVERSAL GAP FUNCTION AND SPECIFIC HEAT

We saw that the γ model is a superconductor in thermal equilibrium at $T < T_c$ —the anomalous averages are nonzero. Having addressed its properties near T_c and in the normal state, we now turn to the superconducting state at $T \ll T_c$.

For all $\gamma > 3$, the problem of determining the gap function $\Delta(\omega)$ and thermodynamic properties in the low-temperature regime reduces to a single parameterless second-order ODE up to corrections of relative order $(T/g)^{r\gamma/3}$, where $r > 0$ is a function of γ . In this section, we first solve this ODE to evaluate the gap function on the Matsubara axis, the free energy and the specific heat, and then discuss the corrections. Similar to the normal state, the absolute specific heat is negative, but the difference between the superconducting and normal state specific heats is positive.

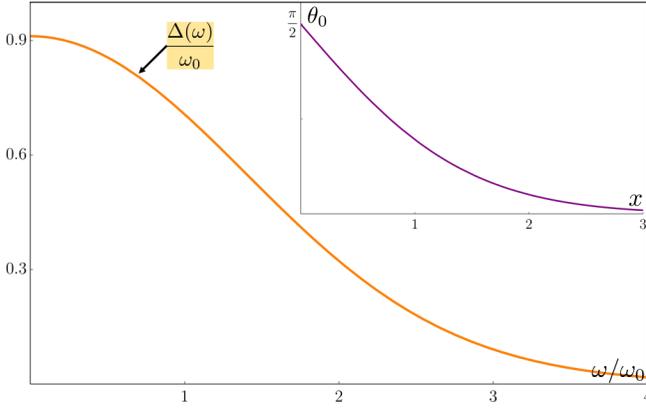


FIG. 4. Universal gap function $\Delta(\omega)$ and the corresponding angle $\theta_0(x)$ spins make with the z axis (inset). Gap functions $\Delta(\omega_n)$ for all $\gamma > 3$ and all values of the coupling g collapse onto this plot in the limit $T \rightarrow 0$ as long as both Δ and ω are measured in units of the energy constant ω_* given by Eq. (58).

A. Gap function

We showed in Sec. IV B that at low temperatures and $\gamma > 3$ the gap equation reduces to Eq. (25). Rescaling the variable $\bar{\omega}$ as in Eq. (28), we obtain a *universal* low-temperature gap equation

$$\frac{d^2\theta}{dx^2} = x \sin \theta(x), \quad (57)$$

where

$$x = \frac{\omega}{\omega_*}, \quad \omega_* = g \frac{[\zeta(\gamma - 2)]^{\frac{1}{3}}}{2^{\frac{1}{3}}} \left(\frac{g}{2\pi T} \right)^{\frac{\gamma}{3}-1} \quad (58)$$

The energy constant ω_* is the coupling g rescaled by a factor that depends on g/T and γ . Eq. (57) is universal in that it is parameterless and therefore independent of γ , g , and T —all dependence on these parameters is in the energy scale ω_* .

Recall from Sec. IV B that $\theta(0) = \frac{\pi}{2}$. Further, the requirement $\Delta_n \geq 0$ discussed in the beginning of Sec. III together with Eqs. (14) and (15) imply $0 \leq \theta_n \leq \frac{\pi}{2}$ for $\omega_n > 0$. Therefore, for $x \geq 0$ we have

$$\theta(0) = \frac{\pi}{2}, \quad \theta(+\infty) = 0, \quad 0 \leq \theta(x) \leq \frac{\pi}{2}. \quad (59)$$

In the Appendix we show that under the conditions (59) there exists a unique solution $\theta_0(x)$ of the nonlinear differential equation (57) and note that $\theta_0(x)$ is real analytic on the entire x -axis by standard theorems of the theory of ODEs. This is a parameterless function, which we plot in the inset to Fig. 4 for $x \geq 0$. Observe also that since $\theta(x) \rightarrow \pi - \theta(-x)$ leaves Eq. (57) invariant and by uniqueness, $\theta_0(x)$ must map into itself under this transformation, i.e., $\theta_0(x) = \pi - \theta_0(-x)$.

Having determined $\theta_0(x)$, we know the gap function for all low temperatures, g , and $\gamma > 3$. Indeed, Eq. (15) implies

$$\frac{\Delta(\omega)}{\omega_*} = Y\left(\frac{\omega}{\omega_*}\right), \quad Y(x) = x \tan \theta_0(x). \quad (60)$$

[Recall that we extended $\Delta(\omega_n)$ to $\Delta(\omega)$ defined on the entire real ω axis in Sec. IV B.] This equation provides the leading small T asymptotic behavior of $\Delta(\omega)$ for any g and γ . Graphically, plots of the gap function $\Delta(\omega_n)$ vs Matsubara frequency

ω_n for any γ and g tend to the same universal curve shown in Fig. 4 as $T \rightarrow 0$, when both $\Delta(\omega_n)$ and ω_n are measured in units of ω_* .

It is straightforward to work out the expansion of $\Delta(\omega)$ at small ω and its large ω asymptote. Both depend on a single constant that needs to be determined numerically. At large ω/ω_* , the angle θ_0 is small and $Y(x) \approx x\theta_0(x)$. Equation (57) becomes the Airy equation. Its solution that goes to zero at infinity is the Airy function of the first kind, $\theta_0(x) \propto \text{Ai}(x)$. Therefore at large $\omega/\omega_* \equiv x$

$$\frac{\Delta(\omega)}{\omega_*} \approx 4.58 \text{Ai}(x) \approx 1.29x^{\frac{3}{4}} \exp\left(-\frac{2x^{\frac{3}{2}}}{3}\right), \quad (61)$$

where we determined the constant of proportionality 4.58 numerically and used the known asymptotic expansion for $\text{Ai}(x)$ [31].

Similarly, at small $x = \omega/\omega_*$,

$$\frac{\Delta(\omega)}{\omega_*} \approx 0.91 - 0.10x^2 + 0.05x^4 + O(x^6). \quad (62)$$

In particular, recalling the definition (58) of ω_* , we find

$$\Delta(0) = 0.72g[\zeta(\gamma - 2)]^{\frac{1}{3}} \left(\frac{g}{2\pi T} \right)^{\frac{\gamma}{3}-1}. \quad (63)$$

Since this result is for $\gamma > 3$, $\Delta(0)$ diverges as $T^{1-\frac{\gamma}{3}}$ for $T \rightarrow 0$ [32]. For fixed T and $\gamma \rightarrow 3^+$ we have

$$\frac{2\Delta(0)}{T_c} \approx 8.47|\gamma - 3|^{-\frac{1}{3}}, \quad (64)$$

where we took the value of T_c for $\gamma = 3$ from Table I. Reference [32] found the same answer, but for $\gamma \rightarrow 3^-$ and with a prefactor 4π instead of 8.47. Note that unlike in the BCS theory, $2\Delta(0)$ here is not the gap in the spectrum [33].

B. Low T free energy and specific heat

To evaluate the free energy \bar{f} , we use Eq. (27). As discussed in Sec. VIII, the free energy contains a diverging T -independent constant because we took the limit $\varepsilon_F \rightarrow \infty$. This divergence is present for superconducting states as well, since in these states $S_n^z \rightarrow \text{sgn}(\omega_n)$ when $|\omega_n| \rightarrow \infty$, same as in the normal state. To isolate it, we subtract and add the noninteracting part of the normal state free energy f_n^Z given by Eq. (48), i.e., $f = f - f_n^Z + f_n^Z$. The difference $f - f_n^Z$ is finite and we have already evaluated the T dependence of f_n^Z in Sec. VIII.

Taking the difference $\bar{f} - \bar{f}_n^Z$, where $\bar{f}_n^Z = f_n^Z/(v_0g^2)$, corresponds to the replacement $S^z \rightarrow S^z - \text{sgn}(x)$ in the Zeeman term in Eq. (27), since $S^z(x) = \text{sgn}(x)$ in the normal state. Using also $S^z = \cos \theta$, $dS/dx = d\theta/dx \equiv \theta'$, and the symmetry property (16) of $\theta(x)$, we obtain the nondimensionalized free energy in the superconducting state at $T \ll T_c$ as

$$\begin{aligned} \bar{f}_s = & \frac{[\zeta(\gamma - 2)\tau^{3-\gamma}]^{\frac{1}{3}}}{2^{\frac{2}{3}}} \int_0^\infty dx \{2x(1 - \cos \theta_0) + (\theta'_0)^2\} \\ & - \frac{\tau^2}{12}, \quad \text{where } \tau = \frac{2\pi T}{g}, \quad \bar{f} = \frac{f}{v_0g^2}. \quad (65) \end{aligned}$$

Substituting the numerical solution for $\theta_0(x)$ discussed in the previous subsection into the integral, we find

$$\bar{f}_s = 1.11[\zeta(\gamma - 2)]^{\frac{2}{3}}\tau^{2-\frac{2\gamma}{3}} - \frac{\tau^2}{12}. \quad (66)$$

The specific heat $c_s = -T \frac{\partial^2 f}{\partial T^2}$ in the superconducting state is

$$c_s = \frac{2\pi^2 v_0}{3} T \left[1 - \left(\frac{T_s}{T} \right)^{\frac{2\gamma}{3}} \right], \quad (67)$$

where

$$T_s = [1.48(\gamma - 3)(2\gamma - 3)]^{\frac{3}{2\gamma}} [\zeta(\gamma - 2)]^{\frac{1}{\gamma}} \frac{g}{2\pi}. \quad (68)$$

It is evident that c_s becomes negative as $T \rightarrow 0$ for any $\gamma > 3$.

This indicates that the γ model is pathological for $T \rightarrow 0$ just as it is above T_c , see the discussion at the end of Sec. VIII. Again, there cannot be an unambiguous explanation of this pathology without the knowledge of the system Hamiltonian. To fix it, we need to modify the effective fermionic action (9), i.e., the γ model itself. One scenario is as follows. Consider the difference between c_s and the normal state specific heat (54) c_n ,

$$c_s - c_n = \frac{2\pi^2 v_0}{3} T \left[\left(\frac{T_n}{T} \right)^\gamma - \left(\frac{T_s}{T} \right)^{\frac{2\gamma}{3}} \right]. \quad (69)$$

This is positive for $T < T_n^3/T_s^2$. A straightforward numerical analysis shows that $T_n^3/T_s^2 > T_c$. Therefore, this difference is always positive when $T < T_c$. Suppose the modification of the fermionic action is such that it introduces a new order parameter independent of S_n . Further, suppose this adds terms to H_s that depend on the new order parameter and not on S_n [see Eqs. (11) and (12)]. Then, it does not affect the superconducting transition and the gap equation, but adds T -dependent terms to the free energy and to the specific heat. The change in the specific heat must be the same for c_s and c_n . Provided the new c_n is positive, so is the new c_s , because Eq. (69) remains valid. This scenario is favorable for the results of this section in the sense that it fixes the pathology while leaving them intact with the exception of the answers for \bar{f}_s and c_s . However, note that $\bar{f}_s - \bar{f}_n$ and $c_s - c_n$ do not change.

C. Accuracy of the local approximation

In this section, we solved the gap equation on the Matsubara axis and evaluated the specific heat at low temperatures for $\gamma > 3$, by replacing the gap equation (17), which we copy here for convenience,

$$\omega_n \sin \theta_n = g^\gamma \pi T \sum_{m \neq n} \frac{\sin(\theta_m - \theta_n)}{|\omega_m - \omega_n|^\gamma}, \quad (70)$$

with the differential equation (57). Now let us investigate the accuracy of this local approximation.

To derive Eq. (57), we expanded $\sin(\theta_m - \theta_n)$ in $(\omega_m - \omega_n)$ to second order, see Eq. (24). The contribution from odd powers of $(\omega_m - \omega_n)$ to the right hand side of Eq. (70) cancels after summation over m . Therefore, the error comes from terms of order 4 and higher. The Lagrange error bound $M(\omega_m - \omega_n)^4/4!$ then gives an upper bound on the difference

between $\sin(\theta_m - \theta_n)$ and our approximation to it. Here M is the maximum value of $d^4 \theta_n / d\omega_n^4$.

Let us split the summation in Eq. (70) into two parts: $|m - n| < L$ and $|m - n| \geq L$, where L is to be determined. For $|m - n| \leq L$ we use the Taylor series expansion of $\sin(\theta_m - \theta_n)$ to third order plus the Lagrange error bound. For $|m - n| > L$, we replace $\sin(\theta_m - \theta_n) \rightarrow 1$, which provides an upper bound for the error from neglecting these terms. Pulling out an overall factor of $2(2\pi T)^{-\gamma}$, we obtain

$$\begin{aligned} & \frac{(2\pi T)^2}{2} \frac{d^2 \theta_n}{d\omega_n^2} \zeta(\gamma - 2) - \frac{(2\pi T)^2}{2} \frac{d^2 \theta_n}{d\omega_n^2} \sum_{k=L}^{\infty} k^{2-\gamma} \\ & + \frac{(2\pi T)^4}{4!} M \sum_{k=1}^L k^{4-\gamma} + \sum_{k=L}^{\infty} k^{-\gamma}. \end{aligned} \quad (71)$$

The first term leads to Eq. (57) after the change of variables (58). The remaining three terms are the error. To estimate the second derivative of $\theta(\omega_n)$ and M , we use the numerical solution of Eq. (57) plotted in Fig. 4. For example, $d^2 \theta / d\omega^2 = \omega_*^{-2} d^2 \theta / dx^2$ with $0 \leq d^2 \theta / dx^2 < 0.6$ and $M \approx 1.4\omega_*^{-4}$. We replace the summations over k in Eq. (71) with integrals and then minimize the error with respect to L . As L turns out to be large, corrections due to the replacement of the summations with integrals are negligible. In this way, we obtain that the relative error [magnitude of the ratio of the sum of the last three terms in Eq. (71) to the first term] is

$$\text{RE} \sim \left(\frac{2\pi T}{g} \right)^{\frac{\gamma}{3}} r = \begin{cases} \gamma - 3, & \text{for } 3 < \gamma \leq 5, \\ 2, & \text{for } \gamma > 5. \end{cases} \quad (72)$$

X. CONCLUSIONS

In this paper, we studied the thermodynamics of a system of fermions near a quantum critical point with extremely retarded interactions of the form $V(\omega_l) = (g/|\omega_l|)^\gamma$, where ω_l is a bosonic Matsubara frequency. The case $\gamma = 2$ of this γ model corresponds to the strong coupling limit of the Eliashberg theory, which is intermediately retarded. Extreme retardation at generic T kicks in for $\gamma \gg 1$ and for $\gamma > 3$ at $T \rightarrow 0$. Note that for $\gamma > 2$, the γ model is a model without a Hamiltonian and is defined through a nonlocal effective Euclidian action only.

The γ model shows two phases: Normal and superconducting. Similarly to the Eliashberg theory, the order parameter is the frequency dependent gap function $\Delta(\omega_n)$, where ω_n is the fermionic Matsubara frequency. We determined the superconducting transition temperature T_c and the order parameter $\Delta(\omega_n)$ near T_c . The amplitude ψ of $\Delta(\omega_n)$ serves as the Landau order parameter. We expanded the free energy functional to order $|\psi|^4$ near the transition to obtain the Landau free energy, from which we derived the jump in the specific heat and the thermodynamic critical field. These answers are asymptotically exact in the limit $\gamma \rightarrow \infty$. We also evaluated the normal state specific heat c_n for arbitrary T and $\gamma > 2$.

Next, we turned our attention to the properties of the γ model at low temperatures. We proved that the global minimum of the free energy is unique (nondegenerate). We derived the universal gap equation, which is a parameterless second-order ODE, and determined the scaling of $\Delta(\omega)$ with T , g , and

γ for all $g, \gamma > 3$, and $T \rightarrow 0$. Building on this, we obtained explicit expressions for the free energy and specific heat in the superconducting state for this range of parameters. We also evaluated $2\Delta(0)/T_c$ and found that it is finite for $T > 0$ and $\gamma > 3$, but diverges as $|\gamma - 3|^{-1/3}$ for $\gamma \rightarrow 3^+$ and as $T^{1-\gamma/3}$ for $T \rightarrow 0$. These results are exact for any $\gamma > 3$ at $T \rightarrow 0$.

Note that “exact” and “asymptotically exact” here and elsewhere in this paper mean exact for the γ model defined by the effective action (9) in the thermodynamic and $\varepsilon_F \rightarrow \infty$ limits. In this regime, fluctuational corrections to the spin chain are negligible and it is at zero effective temperature, i.e., its ground state [which is determined by the gap equation (17)] captures all thermodynamical properties that we evaluated.

We found that the γ model is thermodynamically unstable for $\gamma \geq 2$. Its specific heat is negative above T_c for $\gamma \geq 2$ and also at $T \rightarrow 0$ for $\gamma > 3$. We saw in an earlier paper [30] that when this model is understood as an effective description of the phonon mediated electron-electron interaction ($\gamma = 2$), this instability implies the emergence of a new order above T_c . The new phase breaks the lattice translational symmetry and invalidates the γ model at least in a certain temperature range. For other γ , a microscopic Hamiltonian is similarly necessary to resolve this issue. An interesting open problem is therefore to construct classes of physical many-body Hamiltonians that correspond to the γ model with arbitrary $\gamma > 2$.

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APPENDIX: EXISTENCE AND UNIQUENESS OF SOLUTIONS TO THE UNIVERSAL GAP EQUATION

In this Appendix we show that under the conditions (59) the universal gap equation (57) has a unique solution. Uniqueness is easier to show than existence, so we start with uniqueness, then discuss existence of a solution. We also offer a section about the properties of solutions of Eq. (57) when conditions (59) are dropped.

1. Uniqueness of a solution satisfying conditions (59)

Suppose that both $\theta_0(x)$ and $\tilde{\theta}_0(x)$ are twice continuously differentiable solutions of the ODE (57) that satisfy the conditions (59). Then $\Theta(x) := \theta_0(x) - \tilde{\theta}_0(x)$ vanishes when $x \rightarrow 0$ from the right, and when $x \rightarrow \infty$. Moreover, $\Theta(x)$ is also twice continuously differentiable, and its second derivative is given by

$$\frac{d^2}{dx^2} \Theta(x) = x(\sin[\theta_0(x)] - \sin[\tilde{\theta}_0(x)]), \quad x \in (0, \infty). \quad (A1)$$

We multiply Eq. (A1) by $\Theta(x)$, integrate from $x = 0$ to $x = \infty$, integrate by parts on the left-hand side, and obtain

$$\begin{aligned} 0 &\geq - \int_0^\infty |\Theta'(x)|^2 dx \\ &= \int_0^\infty x |\Theta(x)|^2 \frac{\sin[\theta_0(x)] - \sin[\tilde{\theta}_0(x)]}{\theta_0(x) - \tilde{\theta}_0(x)} dx \geq 0, \quad (A2) \end{aligned}$$

where the prime denotes derivative with respect to x . The second inequality in (A2) follows from the fact that $\sin \theta$ is an increasing function for $\theta \in [-\pi/2, \pi/2]$. Since both inequalities in (A2) are strict if $\theta_0(x_0) \neq \tilde{\theta}_0(x_0)$ for some $x_0 > 0$ (and therefore in some open neighborhood of x_0), the only option compatible with (A2) is: $\theta_0(x) = \tilde{\theta}_0(x)$ for all $x \geq 0$.

Thus uniqueness holds.

2. Existence of a solution satisfying conditions (59)

The uniqueness proof only shows that there cannot exist more than one twice continuously differentiable solution to Eq. (57) that satisfies the conditions (59), while leaving open whether a solution exists at all. In this subsection we show that such a solution does exist, indeed.

We now consider the initial value problem for Eq. (57), with initial data

$$\theta(0) = \frac{\pi}{2}; \quad \theta'(0) = \alpha, \quad (A3)$$

and we treat α as a real parameter that we exhibit explicitly in the solution to the initial value problem, written as $\theta(x; \alpha)$. Our goal is to show that there exists at least one particular value α_0 (possibly not unique, in this section) such that the pertinent solution $\theta(x; \alpha_0)$ of the initial value problem satisfies the remaining conditions in (59); i.e., $\theta(x; \alpha_0)$ vanishes as $x \rightarrow \infty$, and $\theta(x; \alpha_0)$ takes values only in $[0, \pi/2]$ for $x \geq 0$. In this case we may identify $\theta(x; \alpha_0)$ with a sought-after solution $\theta_0(x)$ of Eq. (57) under the conditions (59).

We begin by noting that the Picard-Lindelöf theorem [34] guarantees that for each α the initial value problem for Eq. (57) with initial data (A3) has a unique twice continuously differentiable solution $\theta(x; \alpha)$, which exists for all $x \in \mathbb{R}$. The regularity follows from the facts that the function $x \sin \theta$ is continuous in both x and θ . The uniqueness and global character of such a solution follow from the additional feature that the derivative of $\sin \theta$ is uniformly bounded in absolute value. In fact, $\theta(x; \alpha)$ is analytic in x ; this follows from the fact that the function $x \sin \theta$ is analytic in both x and θ ; see Ref. [35]. Thus it remains to show that there is a solution $\theta_0(x)$ that converges to 0 as $x \rightarrow \infty$, and that this solution does not take values outside of the interval $[0, \pi/2]$ when $x \geq 0$.

Since we demand that the solution $\theta_0(x) \leq \pi/2$, with $\theta_0(0) = \pi/2$, it now follows from Eq. (57) that a necessary condition for the existence of such a solution is that the initial slope $\alpha < 0$. Indeed, for $\theta \in (0, \pi)$ and $x > 0$, the right-hand side of Eq. (57) is > 0 , so that the unique solution $\theta(x; \alpha)$ of Eq. (57) that satisfies the initial data (A3) is convex as long as $\theta(x; \alpha) \in (0, \pi)$. But this means that if there is any $x_0 \geq 0$ for which $\theta(x_0; \alpha) \in (0, \pi)$ and $\theta'(x_0; \alpha) = 0$, then x_0 is a local minimum point of $\theta(x; \alpha)$, and this solution will inevitably increase to values $> \pi/2$.

On the other hand, since we also demand that the solution $\theta_0(x) \geq 0$, its negative initial slope α cannot be too large in magnitude, for it is straightforward to show that there is some $\alpha_0 < 0$ such that for $\alpha < \alpha_0$ there is an $x_0 > 0$ such that the pertinent solution $\theta(x; \alpha)$ satisfies $\theta(x_0; \alpha) = 0$ and $\theta'(x_0; \alpha) < 0$. (We have recycled the symbol x_0 with a new meaning.) In this case it follows right away that $\theta(x_0 + \epsilon; \alpha) < 0$ for some $\epsilon > 0$. To see that there is such an $\alpha_0 < 0$,

recall that $\sin \theta \leq 1$, so that we obtain the estimate

$$\theta''(x; \alpha) \leq x, \quad (\text{A4})$$

and integrating this estimate twice for the stipulated initial data we find that

$$\theta(x; \alpha) \leq \frac{\pi}{2} + \alpha x + \frac{1}{6}x^3. \quad (\text{A5})$$

The cubic polynomial at the right-hand side may have no, or one, or two positive roots, depending on α . If there is at least one positive root, let x_* denote either the unique positive root or the smaller of the two positive roots. Such a root x_* exists if and only if $\alpha + \frac{1}{2}x_*^2 \leq 0$. Setting $\alpha_* = -\frac{1}{2}x_*^2$, our cubic problem becomes $\frac{\pi}{2} - \frac{1}{2}x_*^3 + \frac{1}{6}x_*^3 = 0$, viz., $x_* = (\frac{3\pi}{2})^{\frac{1}{3}}$, which returns

$$\alpha_* = -\frac{1}{2} \left(\frac{3\pi}{2} \right)^{\frac{2}{3}}. \quad (\text{A6})$$

And so, when $\alpha \leq \alpha_*$ the cubic upper bound to the solution $\theta_\alpha(x)$ of our initial value problem, i.e., Eq. (57) with initial conditions (A3), vanishes at x_* with a slope ≤ 0 . Hence, $\theta(x; \alpha)$ itself must have a zero at some $x_0 < x_*$ when $\alpha \leq \alpha_*$. Moreover, $\theta'(x_0; \alpha) < 0$. For suppose $\theta'(x_0; \alpha) = 0$; then both $\theta(x; \alpha)$ and its first x derivative would vanish at x_0 , and by the uniqueness of the solution of the second-order initial value problem formulated with these data at x_0 , the function $\theta(x; \alpha)$ would have to vanish identically, which is a contradiction to the fact that $\theta(0; \alpha) = \frac{\pi}{2}$. So $\theta'(x_0; \alpha) < 0$, and since $\theta(x_0; \alpha) = 0$, it follows that $\theta(x + \epsilon; \alpha) < 0$, which violates the required lower bound 0 for θ_0 . It follows that there is some α_0 with $\alpha_* < \alpha_0 < 0$ such that a further necessary condition for the existence of the desired solution $\theta_0(x)$ is that the initial slope $\alpha \geq \alpha_0$.

Now consider what happens to $\theta(x; \alpha)$ if we start with $\alpha = \alpha_*$ and continuously increase α from there. As just discussed, the solution $\theta(x; \alpha_*)$ to the initial value problem (57), (A3) has a smallest positive zero at $x_0(\alpha_*) < x_*$, and $\theta'(x_0(\alpha_*); \alpha_*) < 0$. Moreover, $\theta'(x; \alpha_*) < 0$ for all $x \in [0, x_0(\alpha_*)]$. Indeed, if there was some $x_* < x_0$ with $\theta'(x_*; \alpha_*) = 0$, then $\theta(x_*; \alpha_*) \in (0, \pi/2)$, and as discussed above, x_* would be a local minimum point of $\theta(x; \alpha_*)$. This solution would increase for $x > x_*$ to values $> \pi/2$ —in contradiction to the cubic upper bound we derived. Thus $\theta(x; \alpha_*)$ decreases monotonically from the value $\frac{\pi}{2}$ at $x = 0$ to the value 0 at $x = x_0(\alpha_*)$. Now increase α continuously above α_* . It is easy to see [just formally integrate Eq. (57) twice, using Eq. (A3)] that for all $0 < x \leq x_0(\alpha)$ one has $\frac{\partial}{\partial \alpha} \theta'(x; \alpha) > 0$ and $\frac{\partial}{\partial \alpha} \theta(x; \alpha) > 0$. Thus the zero $x_0(\alpha)$ moves continuously to the right as α increases. Moreover, as long as $x_0(\alpha) < \infty$, the function $\theta(x; \alpha)$ reaches its first zero at $x_0(\alpha)$ with a nonzero negative slope, $\theta'(x_0(\alpha); \alpha) < 0$. This follows from the already made observation that $\theta(x; \alpha)$ must be identically zero if it vanishes at a finite location $x_0(\alpha)$ with vanishing slope, in contradiction to the initial data $\theta(0; \alpha) = \frac{\pi}{2}$. Thus we can increase α until a value α_0 is reached at which $x_0(\alpha_0) = \infty$, with $\lim_{\alpha \rightarrow \alpha_0} \theta'(x_0(\alpha); \alpha) = 0$ (limit from the left). The solution $\theta(x; \alpha_0)$ is a sought-after solution $\theta_0(x)$.

This demonstrates the existence of a solution to Eq. (57) that satisfies (59).

3. The types of solutions for general initial data

We briefly discuss the general initial value problem for Eq. (57) with initial data at $x = 0$. By the 2π periodicity of the sine function, it suffices to restrict the discussion to data

$$\theta(0) = \vartheta \in [-\pi, \pi]; \quad \text{and} \quad \theta'(0) = \alpha \in \mathbb{R}. \quad (\text{A7})$$

For any pair of such initial data (ϑ, α) there exists a unique analytical solution $\theta(x; \vartheta, \alpha)$ of Eq. (57) for all $x \geq 0$; cf. [35]. The purpose of this subsection is to present a mostly qualitative and partly quantitative overview of the behavior of these solutions.

To get a more intuitive grasp of the possible solutions $\theta(x; \vartheta, \alpha)$ it is helpful to note that for $x > 0$ the variable transformation,

$$\theta(x) = \phi(t), \quad \text{with} \quad t = \frac{2}{3}x^{3/2} \quad (\text{A8})$$

changes Eq. (57) into

$$\ddot{\phi}(t) + \frac{1}{3t} \dot{\phi}(t) = \sin \phi(t) \quad (\text{A9})$$

for $t > 0$. Here we have introduced Newton's dot notation to denote derivatives with respect to t . Equation (A9) describes a damped rigid pendulum, or, equivalently, a point mass moving on an upright circle subject to uniform gravity and Newtonian friction, with a friction coefficient inversely proportional to time t [36]. We count the angle ϕ from the “up” position, i.e., $\phi = 0$ corresponds to the unstable inverted pendulum equilibrium and $\phi = \pi$ to the stable equilibrium. Of course, a priori any integer multiple of 2π may be added to either 0 or π to obtain yet another unstable, respectively stable equilibrium solution for Eq. (A9), because Eq. (A9) in itself does not restrict $\phi(t)$ to lie in any particular interval such as $[0, \pi/2]$.

The damped pendulum Eq. (A9) intuitively suggests that the damping will force the evolution of $\phi(t)$ to converge at $t \rightarrow \infty$ to one of the infinitely many copies of the two possible equilibrium states. Most initial data would lead to a (copy of the) stable equilibrium state $\phi_s = \pi$. However, for each ϑ there should be a discrete set of slopes α , which launch a solution that asymptotically approaches (a copy of) the unstable equilibrium state $\phi_u = 0$. The approach to any stable equilibrium is damped oscillatory, while the approach to an unstable equilibrium is monotonic. Copies of the stable pendulum equilibrium have Newtonian energy $= -1$ and the copies of the unstable one $= +1$.

There is one caveat to what we just wrote. Since the friction coefficient vanishes as $t \rightarrow \infty$, it is in principle conceivable that there also exist solutions that asymptotically approach a dynamical solution of the undamped pendulum equation, instead of converging to an equilibrium solution. We can rigorously rule out an asymptotic approach to a so-called librating solution of the undamped pendulum equation (see below), but we have not rigorously ruled out an asymptotic approach to an oscillating undamped pendulum solution; we expect that it does not occur, though.

In addition to these intuitive insights into the possible types of solution, we offer some exact results.

First, one can show the convergence of the Newtonian energy of the pendulum evolution,

$$E[\phi, \dot{\phi}](t) := \frac{1}{2} |\dot{\phi}(t)|^2 + \cos \phi(t). \quad (\text{A10})$$

The energy is monotonically decreasing with t due to friction. Indeed,

$$\frac{d}{dt}E[\phi, \dot{\phi}](t) = \dot{\phi}(t)[\ddot{\phi}(t) - \sin \phi(t)] = -\frac{1}{3t}|\dot{\phi}(t)|^2 \leq 0, \tag{A11}$$

and for any $0 < t < \infty$ the right-hand side in Eq. (A11) is “ <0 ” if $\dot{\phi} \neq 0$. Since E is bounded below, it must converge as $t \rightarrow \infty$. However, it may or may not converge to its minimum value $E_{\min} = -1$.

We remark that convergence of $E[\phi, \dot{\phi}](t)$ to a constant as $t \rightarrow \infty$ does not in itself imply convergence of $\phi(t)$ as $t \rightarrow \infty$. After all, in the absence of friction the energy is conserved, hence trivially converges to a constant. At the same time, $\phi(t)$ keeps oscillating or librating forever in this case, unless $\phi(t)$ is one of the two equilibrium states. With the help of our energy dissipation identity (A11) we can rule out an asymptotic approach to a librating solution, though. Namely, since $-1 \leq \cos \phi \leq 1$, we can extract from Eq. (A11) the two-sided bounds

$$-1 + C_1 t^{-2/3} \leq E[\phi, \dot{\phi}](t) \leq 1 + C_2 t^{-2/3}, \tag{A12}$$

with C_1 and C_2 some positive constants. Since any asymptotically librating solution has energy $E \geq 1 + \epsilon$ for some

$\epsilon > 0$, it follows from the upper bound in Eq. (A12) that the only possibilities are the asymptotic approach to the stable equilibrium ($E = -1$), to the unstable one ($E = 1$), or to an undamped oscillating pendulum solution ($-1 < E < 1$).

Second, whenever $\phi(t)$ does converge to one of the equilibrium points of the damped rigid pendulum, i.e., to a zero of $\sin \phi$, linearization about that equilibrium yields an accurate approximation at large t . Letting $\xi(t)$ denote the deviation of $\phi(t)$ from the equilibrium point, and dropping terms nonlinear in ξ , we find that $\xi(t)$ is (asymptotically) a solution of one of the following two ODEs, viz.,

$$\ddot{\xi}(t) + \frac{1}{3t}\dot{\xi}(t) = \pm \xi(t). \tag{A13}$$

These are Bessel differential equations. Their solutions for positive t map into the Airy function Ai in the original variables. More precisely, for large enough x the function $\theta(x; \vartheta, \alpha)$ approaches the asymptotic form $C_1 + C_2 \text{Ai}(\pm x + C_3)$. The constant C_1 is an integer multiple of π , and the “+” sign is to be chosen when an unstable pendulum equilibrium is approached monotonically, while the “-” sign pertains to the oscillatory approach to a stable pendulum equilibrium. The constants C_1, C_2 , and C_3 depend on the initial data.

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[12] $\Sigma_\sigma \equiv \Sigma_{\sigma p_1 p_2}$ and $\Phi \equiv \Phi_{p_1 p_2}$, where $\sigma = \uparrow$ or \downarrow , couple to $-i\psi_{p_1\sigma}^* \psi_{p_2\sigma}$ and $\psi_{p_1\uparrow}^* \psi_{-p_2\downarrow}^*$, respectively, see Ref. [9] for more detail.

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[14] This derivation of the spin chain holds everywhere in the configuration space (not only at the stationary points) when $\Omega \rightarrow 0$. A different derivation is available for general Ω . See Sec. II and VII in Ref. [9].

[15] To be more precise, taken at its minimum f is the grand potential per volume as we work at fixed chemical potential. Nevertheless, we colloquially refer to it as the free energy or free energy density.

[16] Here we do not consider the spin-flip saddle points [9] where the sign of the right hand sides of Eqs. (13) flips for certain n .

[17] Cf. Eq. (18) and observe that an infinitely differentiable $\theta(\omega)$ exists for the normal state at any *finite* T as well. However, in this case $\theta(\omega)$ changes abruptly from $\frac{\pi}{2}$ to 0 near $\omega = 0$ for small T and dropping higher order derivatives in Eq. (24) is not justified. In this way, the boundary condition $\theta(0) = \frac{\pi}{2}$ eliminates the normal state as a solution of Eq. (25).

[18] Not to be confused with the real frequency axis for the original complex frequency z . In terms of z , Matsubara frequencies ω_n correspond to discrete points $z_n = i\omega_n$ on the imaginary axis and the “real ω axis”—to imaginary z axis, $z = i\omega$.

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