Solution QM -A:

- (a) Atomic configuration of Eu: $1s^2 2s^2 2p^6 3s^2 2p^6 4s^2 3d^{10} 4p^6 5s^2 4d^{10} 5p^6 6s^2 4f^7$.
- (b) The 5*d* level has l = 3 and s = $\frac{1}{2}$. We have, therefore $j = l \pm s$ which gives $j = \frac{5}{2}$ and $j = \frac{3}{2}$, so the orbital spits into two levels $5d^{5/2}$ and $5d^{3/2}$. Each level can accommodate 2j + 1 electrons, so there are 6 electrons in the $5d^{5/2}$ level and 4 electrons in the $5d^{3/2}$ level.
- (c) For Eu³⁺, we consider the total spin and total angular momentum of the electrons, so S = 3, L = 3, so J = |L S| = 0 (less than half-filled). Thus, Eu³⁺ is expected to have a moment of 0 $\mu_{\rm B}$. For Tb³⁺, S = 3, L = 3, so J = |L + S| = 6 (more than half-filled). Thus, Tb³⁺ is expected to have a moment of $12\mu_{\rm B}$. (Using $g\sqrt{J(J+1)}\mu_B$, rather than $gJ\mu_B$, should be O.K. too)

SOLUTION QB

Taking into account

$$\hat{H} = \frac{\hat{\mathbf{p}}^2}{2m} - \mathbf{F}_0 \mathbf{r}_1$$

we find

$$\frac{d\hat{\mathbf{G}}}{dt} = \frac{\partial\hat{\mathbf{G}}}{\partial t} + \frac{i}{\hbar}[\hat{H}, \hat{\mathbf{G}}] = -\mathbf{F}_0 - \frac{i}{\hbar}[\mathbf{F}_0\mathbf{r}, \hat{\mathbf{p}}] = 0,$$

so that the average

$$\langle \hat{\mathbf{G}} \rangle = \langle \hat{\mathbf{p}} \rangle - \mathbf{F}_0 t = \text{const.}$$

This is a natural quantum mechanical generalization of the statement of classical mechanics that for a particle moving in a uniform field the vector $\mathbf{p}_0 = \mathbf{p}(t) - \mathbf{F}_0 t$ is an integral of motion (since the velocity $\mathbf{v}(t) = \mathbf{v}(0) + \mathbf{F}_0 t/m$) and is equal to particle's momentum at t = 0.

Solution QM-C1

2. (a) From the Hamiltonian,

$$H = \frac{P^2}{2m} + \frac{1}{2}m\omega^2 x^2 = \hbar\omega \left(\hat{a}^+ \hat{a} + \frac{1}{2}\right)$$

, where

$$\hat{a} = \frac{1}{\sqrt{2}} \left(\xi + \frac{\partial}{\partial \xi} \right)$$

$$\hat{a}^{+} = \frac{1}{\sqrt{2}} \left(\xi - \frac{\partial}{\partial \xi} \right)$$
 and $\xi^{2} \equiv \frac{m\omega}{\hbar} x^{2}$.

The ground state can be calculated from the condition, $\,\hat{a}\varphi_{0}=0$:

$$\varphi_0(x) = A e^{-\frac{1}{2}\xi^2}$$

For normalization,

$$1 = \int_{-\infty}^{\infty} |\varphi_0|^2 dx = A^2 \left(\frac{\hbar}{m\omega}\right)^{\frac{1}{2}} \int_{-\infty}^{\infty} e^{-\xi^2} d\xi = A^2 \left(\frac{\pi\hbar}{m\omega}\right)^{\frac{1}{2}}$$

Therefore $\varphi_0 = \left(\frac{m\omega}{\pi\hbar}\right)^{\frac{1}{4}} e^{-\frac{m\omega}{2\hbar}x^2}$

(b) When
$$\omega \rightarrow \frac{1}{2}\omega$$
, the new ground state is $\varphi_0^f = \left(\frac{m\omega}{2\pi\hbar}\right)^{\frac{1}{4}} e^{-\frac{m\omega}{4\hbar}x^2}$

Therefore, the probability that the particle is still in the ground state of the new oscillator is

$$\left\langle \varphi_{0}^{f} \middle| \varphi_{0} \right\rangle = \left(\frac{1}{2}\right)^{\frac{1}{4}} \left(\frac{4}{3}\right)^{\frac{1}{2}}$$
$$P = \left| \left\langle \varphi_{0}^{f} \middle| \varphi_{0} \right\rangle \right|^{2} = \frac{4}{3\sqrt{2}}$$

Solution QM-C2

(i) The equation of motion for the two-component wave function for t>0 (we have $\jmath=\frac{1}{2}~\omega)$ is,

$$\frac{1}{2}g\mu_B\sigma_x H\psi = i\hbar\partial_t\psi,\tag{1}$$

or

$$\begin{pmatrix} 0 & -i\omega_0 \\ -i\omega_0 & 0 \end{pmatrix} \begin{pmatrix} a \\ b \end{pmatrix} = \begin{pmatrix} \partial_t a \\ \partial_t b \end{pmatrix}$$
(2)

where $\omega_0 = g\mu_B \sigma_x H/2\hbar$ We find: $a = \cos \omega_0 t$, $b = -i \sin \omega_0 t$. (ii) $\langle J_x \rangle = 0$, $\langle J_y \rangle = -\frac{1}{2} \sin(2\omega_0 t)$, $\langle J_z \rangle = \frac{1}{2} \cos(2\omega_0 t)$. (iii) $T > \omega_0^{-1}$.

SOLUTION QM-D1

Recall that operators \hat{a} and \hat{a}^{\dagger} act in a space of states of the form (here and below summation over n in all equations is from n = 0 to $n = \infty$)

$$|\Psi\rangle = \sum c_n |n\rangle = c_0 |0\rangle + c_1 |1\rangle + \dots,$$

where $|n\rangle$ stands for the *n*-particle (quanta) state. Further, recall that

$$\hat{a}|n\rangle = \sqrt{n}|n-1\rangle, \quad \hat{a}^{\dagger}|n\rangle = \sqrt{n+1}|n+1\rangle.$$

Eigenstates $|\alpha\rangle = \sum c_n |n\rangle$ and eigenvalues α of the bosonic annihilation operator \hat{a} are determined by the eigenvalue equation $\hat{a}|\alpha\rangle = \alpha |\alpha\rangle$. Since

$$\hat{a}|\alpha\rangle = \sum c_n \sqrt{n}|n-1\rangle = \sum c_{n+1} \sqrt{n+1}|n\rangle,$$

the eigenvalue equation takes the form

$$\sum \left(c_{n+1}\sqrt{n+1} - \alpha c_n \right) |n\rangle = 0.$$

Therefore, taking into account the independence of states $|n\rangle$, we obtain

$$c_{n+1} = \frac{\alpha}{\sqrt{n+1}}c_n = \frac{\alpha}{\sqrt{n+1}}\frac{\alpha}{\sqrt{n}}c_{n-1} = \dots = \frac{\alpha^{n+1}}{\sqrt{(n+1)!}}c_0$$

We see that eigenvalues of \hat{a} are arbitrary complex numbers, while the eigenstates can be normalized to one. The normalization condition yields

$$\langle \alpha | \alpha \rangle = \sum_{n} |c_{n}|^{2} = |c_{0}|^{2} \sum_{n} \frac{|\alpha|^{2n}}{n!} = 1$$
, i.e. $|c_{0}|^{2} = e^{-|\alpha|^{2}}$.

Thus, the particle number distribution is given by

$$w_n = |c_n|^2 = \exp(-|\alpha|^2) \frac{|\alpha|^{2n}}{n!},$$

which is a Poisson distribution with $\langle n \rangle = |\alpha|^2$.

The eigenvalue problem for the bosonic creation operator $\hat{a}^{\dagger}|\beta\rangle = \beta|\beta\rangle$ does not have solutions, since the component with the lowest particle number in any state $\beta|\beta\rangle$ is absent in $\hat{a}^{\dagger}|\beta\rangle$.

QM – D2, Solutions

For this spherically symmetric potential, the Schrödinger equation:

$$-\frac{\hbar^2}{2m}\nabla^2\psi + V\psi = E\psi$$

for l = 0 and the specified potential reduces to:

$$-\frac{\hbar^2}{2m}\left(R''(r) + \frac{2}{r}R'(r)\right) - \frac{V_o a}{r}e^{-\frac{r}{a}}R(r) = ER(r)$$

(a) Now, we have the trial wave function and can calculate its derivatives:

$$R(r) = e^{-\frac{\beta}{a}r} \Longrightarrow R' = \left(-\frac{\beta}{a}\right)e^{-\frac{\beta}{a}r} \Longrightarrow R'' = \left(\frac{\beta}{a}\right)^2 e^{-\frac{\beta}{a}r}$$

and, integrating by parts we have:

$$\int_{0}^{\infty} e^{-\frac{2\beta}{a}r} r^2 dr = \frac{2}{\left(\frac{2\beta}{a}\right)^3} \text{ and } \int_{0}^{\infty} e^{-\frac{2\beta}{a}r} r dr = \frac{1}{\left(\frac{2\beta}{a}\right)^2} \text{ and, by extension, } \int_{0}^{\infty} e^{-\frac{(2\beta+1)}{a}r} r dr = \frac{1}{\left(\frac{2\beta+1}{a}\right)^2}$$

To employ the variational principle, we wish to find $E(\beta) = \frac{\langle R|H|R\rangle}{\langle R|R\rangle}$ (Note: angular integrals cancel.) Writing $\langle R|R\rangle E(\beta) = \langle R|H|R\rangle$ we have

$$\left[\int_{0}^{\infty} e^{-\frac{2\beta}{a}r} r^{2} dr\right] E(\beta) = -\frac{\hbar^{2}}{2m} \left[\left(\frac{\beta}{a}\right)^{2} \int_{0}^{\infty} e^{-\frac{2\beta}{a}r} r^{2} dr + 2\left(-\frac{\beta}{a}\right) \int_{0}^{\infty} e^{-\frac{2\beta}{a}r} r dr\right] - V_{o}a \int_{0}^{\infty} e^{-\frac{2\beta+1}{a}r} r dr$$

Which reduces to:

$$\begin{bmatrix} \frac{2}{\left(\frac{2\beta}{a}\right)^3} \end{bmatrix} E(\beta) = -\frac{\hbar^2}{2m} \left[\left(\frac{\beta}{a}\right)^2 \frac{2}{\left(\frac{2\beta}{a}\right)^3} - 2\left(\frac{\beta}{a}\right) \frac{1}{\left(\frac{2\beta}{a}\right)^2} \right] - V_o a \frac{1}{\left(\frac{2\beta+1}{a}\right)^2}$$
$$= -\frac{\hbar^2}{2m} \left[\left(\frac{2}{8}\right) \frac{\left(\frac{\beta}{a}\right)^2}{\left(\frac{\beta}{a}\right)^3} - \left(\frac{2}{4}\right) \frac{\left(\frac{\beta}{a}\right)}{\left(\frac{\beta}{a}\right)^2} \right] - \frac{V_o a^3}{\left(2\beta+1\right)^2}$$
$$E(\beta) = \left[\frac{\hbar^2}{2m} \frac{1}{4} \frac{a}{\beta} - \frac{V_o a^3}{\left(2\beta+1\right)^2} \right] \left[\frac{4\beta^3}{a^3} \right] = \left[\frac{\hbar^2}{2m} \frac{\beta^2}{a^2} - \frac{4V_o \beta^3}{\left(2\beta+1\right)^2} \right]$$

(b) For a bound state, must have $E < 0 \rightarrow$ Find condition for $E(\beta) = 0$. From above expression for $E(\beta)$, and seeking a solution for $\beta > 0$, we see that $E(\beta) = 0$ when:

$$\frac{\hbar^2}{2ma^2} = \frac{4V_o\beta}{(2\beta+1)^2} \quad \text{or, equivalently,} \quad \frac{\hbar^2}{8mV_oa^2} (2\beta+1)^2 = \beta \text{. Rewrite this as} \\ A(2\beta+1)^2 = \beta \text{ where } A = \frac{\hbar^2}{8mV_oa^2} \text{ . This gives the quadratic equation:} \\ 4\beta^2 + 4\beta + 1 = \frac{\beta}{A} \Rightarrow 4\beta^2 + (4-\frac{1}{A})\beta + 1 = 0 \text{. Find the roots of this quadratic:} \\ \beta = \frac{-(4-\frac{1}{A})\pm\sqrt{(4-\frac{1}{A})^2-16}}{8} \quad \text{for real solutions, need } (4-\frac{1}{A})^2 - 16 > 0 \text{ so} \\ \pm (4-\frac{1}{A}) - 4 \ge 0 \Rightarrow -4 + \frac{1}{A} \ge 4 \Rightarrow \frac{1}{A} \ge 8 \Rightarrow A = \frac{\hbar^2}{8mV_oa^2} \le \frac{1}{8} \Rightarrow Vo \ge \frac{\hbar^2}{ma^2}$$

(c) Minimum of energy for this trial wave function is given for β which solves:

$$\frac{\partial E(\beta)}{\partial \beta} = 0 \quad \Rightarrow \quad \left[\frac{\hbar^2 \beta}{ma^2} - \frac{12V_o \beta^2}{(2\beta+1)^2} + \frac{8V_o \beta^3 \cdot 2}{(2\beta+1)^3}\right] = 0 \quad \text{which, for } \beta \neq 0 \text{ requires solving:}$$

$$\frac{\hbar^2 (2\beta+1)^3}{ma^2} - 12V_o \beta (2\beta+1) + 16V_o \beta^2 = 0$$