

Searching for Particle Dark Matter with the GLAST-LAT

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WIMPs

- WIMPs are the most promising candidate for a dark-matter particle
- Eluded direct discovery or identification
- Perhaps we can detect them indirectly?
 - Look at the products of WIMP self-annihilation

WIMP-annihilation Products

$$\frac{dN_\gamma}{dE} = \frac{dN_{\text{cont}}}{dE}(E) + \sum_X b_{\gamma X} n_\gamma \delta(E - m_\chi(1 - m_\chi^2/4m_X^2))$$

- First term from tree-level final states
 - “Cascades” and many-body final states
 - For Majorana WIMPs, light fermion production suppressed
 - Primary fermionic products are $b\bar{b}$, $t\bar{t}$, and $\tau\bar{\tau}$

WIMP-annihilation Products

$$\frac{dN_\gamma}{dE} = \frac{dN_{\text{cont}}}{dE}(E) + \sum_X b_{\gamma X} n_\gamma \delta(E - m_\chi(1 - m_\chi^2/4m_X^2))$$

- Second term from 2-body final states

$$\gamma + X \quad \text{or} \quad 2\gamma$$

- In this paper, assume $b\bar{b}$ and 2γ dominate

GLAST



- Detect γ -rays in 20 MeV to 300 GeV range
- 50 times more sensitive; 10 times greater energy range; and 2 times better energy and angular resolution than EGRET
- Silicon detectors rather than spark chambers (EGRET)

Detection

- “Golden” signal
 - Spectral line at WIMP mass
 - Difficult to explain through other causes
 - Very small astrophysical uncertainties
 - Loop-suppressed
 - Low photon count expected

Detection

$$\frac{d\phi_\gamma}{dE_0} = \frac{\sigma v}{8\pi} \frac{c}{H_0} \frac{\bar{\rho}_0^2}{M_\chi^2} \int dz (1+z)^3 \frac{\Delta^2(z)}{h(z)} \frac{dN_\gamma(E_0(1+z))}{dE} e^{-\tau(z, E_0)}$$

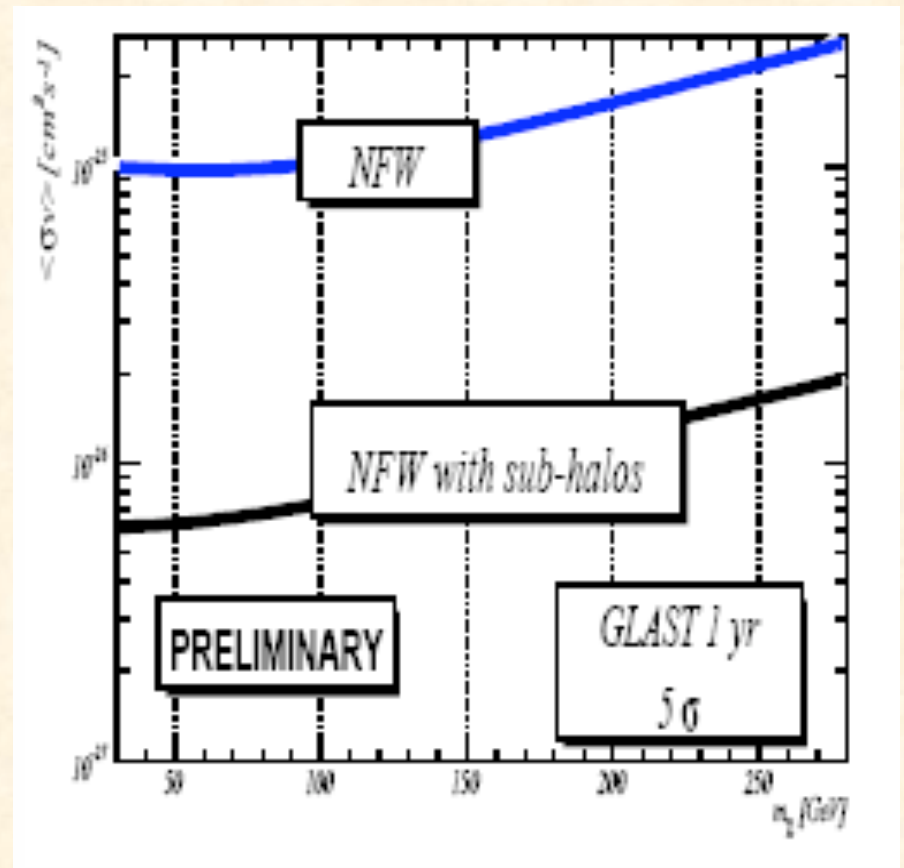
- Particle physicists must find σ and M_χ
- $\frac{dN_\gamma}{dE}$ from before
- This goes as DM density *squared*
 - “Clumpiness” strongly affects signal strength

Halo Substructure

- N-body simulations show small subhalos
- Greater density means more WIMP annihilation products
 - Greater flux of gamma rays
- This can be used to detect satellites within the Milky Way halo

Conclusions

- GLAST can help constrain the cross-section for the WIMP self-annihilation



Conclusions

- GLAST can also give us insight into the existence or absence of subhalos

