

Final solutions

2026

1. Virial expansion

Dilute classical gas with pair potential $u(r_{ij}) = u(|\vec{r}_i - \vec{r}_j|)$, where \vec{r}_i = position of particle i .

(a) Start with the canonical partition function:
(N particles) $Z_N = \frac{1}{N! \lambda_T^{3N}} \int d^N \vec{r} e^{-\beta U(\vec{r}_1, \dots, \vec{r}_N)}$,

where $\lambda_T = \sqrt{\frac{2\pi \hbar^2}{m k_B T}}$ is the thermal de Broglie wavelength

and $U(\vec{r}_1, \dots, \vec{r}_N) = \sum_{i < j} u(r_{ij})$.

Introduce $f_{ij} = e^{-\beta u(r_{ij})} - 1$, then $f_{ij} = f(r_{ij})$

$$e^{-\beta U} = \prod_{i < j} (1 + f_{ij})$$

For low density, we expect:

$$\prod_{i < j} (1 + f_{ij}) \approx 1 + \sum_{i < j} f_{ij} \quad \text{and thus}$$

$$Z_N \approx Z_N^{(0)} + Z_N^{(1)}$$

$$Z_N^{(0)} = \frac{V^N}{N! \lambda_T^{3N}}, \quad \text{whereas}$$

$$Z_N^{(1)} = \frac{1}{N! \lambda_T^{3N}} \sum_{i < j} \underbrace{\int d^N \vec{r} f_{ij}}_{V^{N-2} \int d^3 r_i d^3 r_j f(r_{ij})} =$$

$$\stackrel{\text{transl. inv.}}{\approx} V^{N-1} \int d^3 r f(r).$$

$\sum_{i < j}$ is over $\frac{N(N-1)}{2}$ particle pairs.

$$\text{Thus, } Z_N^{(1)} = Z_N^{(0)} \frac{1}{V^N} \frac{N(N-1)}{2} V^{N-1} \int d^3 r f(r) =$$

$$= Z_N^{(0)} \frac{N(N-1)}{2V} \int d^3 r f(r)$$

$$\text{Next, } F = -k_B T \log Z_N =$$

$$= -k_B T \log \left\{ Z_N^{(0)} \left[1 + \frac{N(N-1)}{2V} \int d^3 r f(r) \right] \right\} \approx$$

expect to be small, $\ll 1$

$$\approx -k_B T \left[\log Z_N^{(0)} + \frac{N(N-1)}{2V} \int d^3 r f(r) \right]$$

$$\text{Finally, } p = k_B T \left(\frac{\partial \log Z_N}{\partial V} \right)_T = - \left(\frac{\partial F}{\partial V} \right)_T.$$

$$p = k_B T \frac{\lambda^{3N} N!}{V^N} \frac{N V^{N-1}}{N! \lambda^{3N}} \stackrel{\leftarrow N \gg 1}{\approx} k_B T \frac{N^2}{2V^2} \int d^3r f(r) =$$

$$= \frac{k_B T N}{V} \left[1 - \frac{N}{2V} \int d^3r f(r) \right]$$

This yields $B_2 = -\frac{1}{2} \int d^3r f(r)$, and

$$\frac{p}{nk_B T} \approx 1 + B_2 n$$

≡≡≡

↑ $n = \frac{N}{V}$

(b) Hard spheres:

$$u(r) = \begin{cases} \infty, & r < 2a \\ 0, & r > 2a \end{cases}$$

Then $e^{-\beta u(r)} = \begin{cases} 0, & r < 2a \\ 1, & r > 2a \end{cases} \Rightarrow f(r) = \begin{cases} -1, & r < 2a \\ 0, & r > 2a \end{cases}$

Thus, $B_2 = -\frac{1}{2} \int_{r < 2a} d^3r (-1) = \frac{1}{2} \frac{4\pi}{3} (2a)^3 = \frac{16\pi a^3}{3}$

≡

$B_2 \sim$ excluded volume of the particle
 Note that $B_2 > 0 \Rightarrow$ pressure is higher than in the ideal gas

(c) If $\underbrace{u(r) < 0}_{\text{attractive interactions}} \Rightarrow e^{-\beta u(r)} > 1 \Rightarrow f(r) > 0$.

Then $B_2 < 0 \Rightarrow$ the pressure is lowered compared to the ideal gas.

2. Blackbody radiation

(a) grand-canonical partition function:

$$Q = \prod_{\vec{k}, \vec{\epsilon}} \frac{1}{1 - e^{-\beta \hbar \omega_{\vec{k}}}}, \quad \text{where } \omega_{\vec{k}} = \underbrace{c |\vec{k}|}_{\substack{\text{speed of} \\ \text{light}}}$$

grand potential:

$$\Omega = -k_B T \log Q = k_B T \sum_{\vec{k}, \vec{\epsilon}} \log(1 - e^{-\beta \hbar \omega_{\vec{k}}}) =$$

$$= 2 k_B T \sum_{\vec{k}} \log(1 - e^{-\beta \hbar \omega_{\vec{k}}})$$

↑
sum over
polarizations

Now, use $\sum_{\vec{k}} \rightarrow \frac{V}{(2\pi)^3} \int d^3k$ and standardize the integral:

$$\Omega = 2 k_B T \frac{V}{(2\pi)^3} 4\pi \int_0^{\infty} dk k^2 \log(1 - e^{-\beta \hbar c k}) \quad \text{⊖}$$

↑
 $x = \beta \hbar c k$

$$\text{⊖} \quad \frac{V k_B T}{\pi^2} \frac{1}{(\beta \hbar c)^3} \underbrace{\int_0^{\infty} dx x^2 \log(1 - e^{-x})}_{= -\frac{\pi^4}{45}}$$

Finally, $\Omega = -\frac{V k_B T}{\pi^2 (\beta \hbar c)^3} \frac{\pi^4}{45} = -\frac{\pi^2}{45} \frac{V}{\hbar^3 c^3} (k_B T)^4$

note that $\Omega = A \beta^{-4}$, $A = -\frac{\pi^2}{45} \frac{V}{\hbar^3 c^3}$

-4-

(b)
$$P = - \left(\frac{\partial \Omega}{\partial V} \right)_T = - \frac{\Omega}{V}, \text{ since}$$

$\Omega \sim V$ from part (a).

Thus,
$$P = \frac{\pi^2}{45} \frac{(k_B T)^4}{\hbar^3 c^3} \sim T^4$$

===== p steeply increases with T

(c) Internal energy

$$U = + \frac{\partial}{\partial \beta} (\beta \Omega)_V = + \frac{\partial}{\partial \beta} (A \beta^{-3}) =$$

$$= -3 \beta^{-4} A = 3 \underbrace{\frac{\pi^2}{45} \frac{(k_B T)^4}{\hbar^3 c^3}}_{\text{"}p\text{"}} V, \text{ or}$$

$$u = \frac{U}{V} = 3p \Rightarrow p = \frac{u}{3}$$

=====

(d) Entropy

$$S = - \left(\frac{\partial \Omega}{\partial T} \right)_V \Rightarrow S' = - A k_B^4 4 T^3 =$$

$$= -3 A \beta^{-4} \times \frac{1}{T} \frac{4}{3} = \frac{4}{3} \frac{u}{T}$$

=====

Indeed, $\Omega = U - TS - \mu N \Rightarrow S' = \frac{U - \Omega}{T}$

$\mu=0$
here

But $u = 3pV$, $\Omega = -pV$:

$$S' = \frac{4pV}{T} = \frac{4}{3} \frac{u}{T}, \text{ as desired.}$$
$$pV = \frac{u}{3}$$

3. Mean-field Ising model long-range coupling

(a) Consider
$$H = -\frac{2J}{N} \sum_{\substack{i \in A \\ j \in B}} S_i S_j - h_A \sum_{i \in A} S_i - h_B \sum_{j \in B} S_j, \quad S_i = \pm 1.$$

a bi-partite lattice with 2 sublattices, with $\frac{N}{2}$ spins each.

Define
$$\begin{cases} m_A = \langle S_i \rangle_{i \in A} \\ m_B = \langle S_j \rangle_{j \in B} \end{cases}$$
 per-spin magnetization

Mean-field: each spin in A feels

$$\sum_{j \in B} S_j \Rightarrow \underbrace{\frac{N}{2} m_B}_{\text{on average}}$$

Likewise, each spin in B feels

$$\sum_{i \in A} S_i \Rightarrow \underbrace{\frac{N}{2} m_A}_{\text{on average}}$$

Then
$$H_{MF} = -\frac{J}{N} \sum_{i \in A} S_i \left(\frac{N}{2} m_B \right) - h_A \sum_{i \in A} S_i -$$

$$-\frac{J}{N} \sum_{j \in B} S_j \left(\frac{N}{2} m_A \right) - h_B \sum_{j \in B} S_j =$$

$$= - \underbrace{\left(\frac{J}{2} m_B + h_A \right)}_{h_A^{eff}} \sum_{i \in A} S_i - \underbrace{\left(\frac{J}{2} m_A + h_B \right)}_{h_B^{eff}} \sum_{j \in B} S_j$$

Thus, we have:

$$\begin{cases} m_A = \tanh \left[\beta \left(\frac{J}{2} m_B + h_A \right) \right], \\ m_B = \tanh \left[\beta \left(\frac{J}{2} m_A + h_B \right) \right]. \end{cases} \quad (*)$$

(b) now, consider $h_A = 0, h_B = 0$:

$$\begin{cases} m_A = \tanh \left[\frac{\beta J}{2} m_B \right], \\ m_B = \tanh \left[\frac{\beta J}{2} m_A \right]. \end{cases}$$

Expand near T_c :

$$\begin{cases} m_A \approx \frac{\beta J}{2} m_B, \\ m_B \approx \frac{\beta J}{2} m_A \end{cases}$$

similar
for m_B

$$\Rightarrow m_A = \left(\frac{\beta J}{2} \right)^2 m_A$$

have to have
this = 1, otherwise
only $m_A = 0$ works

Then $\beta_c = \frac{2}{J} \Rightarrow T_c = \frac{J}{2k_B}$. ferromagnetic
transition

We can also do: $\begin{pmatrix} m_A \\ m_B \end{pmatrix} = \begin{pmatrix} 0 & \frac{\beta J}{2} \\ \frac{\beta J}{2} & 0 \end{pmatrix} \begin{pmatrix} m_A \\ m_B \end{pmatrix}$, or

$$\begin{pmatrix} 1 & -\frac{\beta J}{2} \\ -\frac{\beta J}{2} & 1 \end{pmatrix} \begin{pmatrix} m_A \\ m_B \end{pmatrix} = 0.$$

To have non-trivial solutions, need
 $\det(\dots) = 0 \Rightarrow \left(\frac{\beta J}{2}\right)^2 = 1$, as before.

(c) Now, set $h_A = h$, $h_B = -h$
 (staggered fields)

Eq. (*) becomes:

$$\begin{cases} m_A = \tanh\left[\beta\left(\frac{J}{2}m_B + h\right)\right], \\ m_B = \tanh\left[\beta\left(\frac{J}{2}m_A - h\right)\right]. \end{cases}$$

Linearize and solve for $m_S = \frac{m_A - m_B}{2}$:

$$m_A - m_B = \beta \frac{J}{2} (m_B - m_A) + 2\beta h, \text{ or}$$

$$m_S = \frac{\beta h}{1 + \beta \frac{J}{2}} \quad \text{staggered magnetization}$$

(d) Staggered susceptibility:

$$\chi_S = \left(\frac{\partial m_S}{\partial h}\right)_T = \frac{\beta}{1 + \beta \frac{J}{2}}$$

At $T = T_c$, $\chi_c = \frac{\beta}{2}$ no divergence
 at T_c , stable
 anti-ferromagnetic
 mode