

Homework 3  
(Due February 11)

Problem 1

The table of all the isotopes of Carbon:

the atomic weight, A	the atomic mass (amu)	the abundance %	the half life	the decay mode(s)
8	8.037		$2 \times 10^{-23} s$	$2p$ ( ${}^8Be$ )
9	9.031		126.5 ms	+ 2 $\alpha$ ( ${}^9Li$ )
10	10.016		19.225 s	e ( ${}^{10B}$ )
11	11.011		20.39 min	e ( ${}^{11B}$ )
12	12.000	98.89		
13	13.003	1.11		
14	14.003		5730 years	$\beta$ ( ${}^{14N}$ )
15	15.010		2.449 s	$\beta$ ( ${}^{15N}$ )
16	16.015		0.447 s	$\beta$ ( ${}^{16N}$ )
17	17.023		193 ms	$\beta$ ( ${}^{17N}$ ) $\beta + n$ ( ${}^{16N}$ )
18	18.027		95 ms	$\beta$ ( ${}^{18N}$ )
19	19.035		49 ms	$\beta$ ( ${}^{19N}$ )
20	20.040		14 ms	$\beta$ ( ${}^{20N}$ )
21	21.050			$\beta$ ( ${}^{20N}$ )
22	22.057		> 200 ns	

The table of all isotopes of Iron:

45	45.014		> 350 nm	EC ( ${}^{46Mn}$ )
46	46		20 ms	EC ( ${}^{47Mn}$ )
47	46.992		27 ms	EC ( ${}^{48Mn}$ )
48	47.98		44 ms	EC ( ${}^{49Mn}$ )
49	48.973		70 ms	EC ( ${}^{50Mn}$ )
50	49.962		150 ms	EC ( ${}^{51Mn}$ )
51	50.956		305 ms	EC ( ${}^{52Mn}$ )
52	51.948		8.275 h, 45.9 s	EC ( ${}^{53Mn}$ )
53	52.945	5.85		
54	53.939		2.73 years	(EC ${}^{55Mn}$ )
55	54.9382			
56	55.934	91.45		
57	56.935	2.12		
58	57.933	0.28		
59	58.934		44.503 days	$\beta$ ( ${}^{59Co}$ )
60	59.934		$1.5 \times 10^6$ years	$\beta$ ( ${}^{60Co}$ )
61	60.936		5.98 min	$\beta$ ( ${}^{61Co}$ )
62	61.936		6.8 s	$\beta$ ( ${}^{62Co}$ )
63	62.94		6.1 s	$\beta$ ( ${}^{63Co}$ )
64	63.94		2 s	$\beta$ ( ${}^{64Co}$ )
65	64.944		0.4 s	$\beta$ ( ${}^{65Co}$ )
66	65.945		600 ms	$\beta$ ( ${}^{66Co}$ )
67	66.95		> 200 ns	$\beta$ ( ${}^{67Co}$ )
68	67.952		0.1 s	$\beta$ ( ${}^{68Co}$ )
69	68.957		> 150 s	$\beta$

### Problem 2

$T_p = 10.8 \text{ MeV}$ , Elastic scattering from  $^{197}\text{Au}$

(a) No, we don't relativistic kinematics due to  $M_p \approx 938 \text{ MeV} \gg T_p$

(b) The maxim. angle is  $\pi$ .

(c) At  $\theta = \pi$  the momentum of the proton can be found from conservation law of energy and momentum:

$$\left. \begin{aligned} T_p &= \frac{M_p v_p^2}{2} + \frac{M_{Au} v_{Au}^2}{2} \\ M_p v_0 &= M_{Au} v_{Au} - M_p v_p \end{aligned} \right\}$$

due to  $M_{Au} \gg M_p$

(d)  $p_{Au} = 0$ ;  $p_p \approx 2p_0 = 2M_p v_{p0}$

$$= \boxed{285 \text{ MeV}}$$

$T_e = 10.8 \text{ MeV}$ , electrons are elastically scattered from  $^{197}\text{Au}$ .

(a) Yes  $m_e = 0.511 \text{ MeV} \ll T_e$

(b) The maximum angle is  $\pi$ .

(c)  $T_e = \frac{m_e c^2}{\sqrt{1-v^2/c^2}} - m_e c^2 \Rightarrow (T_e + m_e c^2)^2 = m_e^2 c^4 [1 - v^2/c^2]^{-2} \Rightarrow$

$$\Rightarrow v = c \sqrt{1 - \left[ \frac{T_e}{m_e c^2} + 1 \right]^{-2}}$$

$$(T_e + m_e c^2)^2 = p^2 c^2 + m_e^2 c^4 \Rightarrow pc = \sqrt{(T_e + m_e c^2)^2 - m_e^2 c^4} = \sqrt{(10.8 + 0.511)^2 - 0.511^2}$$

$$= 11.3 \text{ MeV}$$

(d) Because of a huge difference in masses  $M_{Au} \gg m_e \Rightarrow p_{Au} = 0$

$$p_e = 2p_0 = \boxed{22.3 \text{ MeV}}$$

### Problem 3

(a) The form factor for a spherically symmetric nuclear charge density

$$F(q^2) = \int e^{i\vec{q}\cdot\vec{r}} f(\vec{r}) r^2 dr d\Omega = 2\pi \int_0^\pi \int_0^{2\pi} e^{iqr \cos\theta} f(\vec{r}) r^2 dr \sin\theta d\theta d\phi$$

$$= 2\pi \int_0^\pi \int_{-1}^1 e^{iqr x} f(r) r^2 dr dx = 0$$

$$= 2\pi \int_0^\infty f(r) r^2 dr \cdot \frac{e^{iqr/h} - e^{-iqr/h}}{2iq/h} = \frac{4\pi h}{q} \int_0^\infty r dr f(r) \sin\left(\frac{qr}{h}\right) dr$$

Due to the spherical symmetry  $f(\vec{r}) = f(r) = \frac{\rho(r)}{Ze}$ , hence

$$F(q^2) = \frac{4\pi h}{qZe} \int \rho(r) \cdot r \cdot \sin\left(\frac{qr}{h}\right) dr$$

Logged

(b) Using expansion  $\sin \alpha = 0 + f'(0) \alpha + \frac{f''(0)}{2!} \alpha^2 + \frac{f'''(0)}{3!} \alpha^3 + \dots$   
 and for  $\frac{qz}{\hbar} = \alpha$  the integral  $\frac{2^1}{2!}$  is zero,  
 the only function will be  $\sin\left(\frac{qz}{\hbar}\right) \approx \frac{qz}{\hbar} - \left(\frac{qz}{\hbar}\right)^3 \frac{1}{6} + \dots$   
 let's substitute that into the integral  $F(q^2)$ :

$$F(q^2) \approx \frac{4\pi\hbar}{q^2 e} \left[ \int_0^\infty \rho(z) \cdot z \frac{qz}{\hbar} dz - \int \rho(z) \cdot z \left(\frac{qz}{\hbar}\right)^3 \frac{1}{6} + \dots \right] =$$

$$= \int_0^\infty 4\pi f(z) \cdot z^2 dz - \frac{1}{6} \int_0^\infty f(z) \cdot q^2 \cdot \frac{z^4}{\hbar^4} 4\pi dz = \boxed{1 - \frac{1}{6} \frac{q^2}{\hbar^2} \langle z^2 \rangle + \dots}$$

### Problem 4

Notation  $nL_J = 15\frac{1}{2}$  The nuclei are in their ground state. 82 neutrons  
 + One extra neutron. New state?

Using the Shell Model, "82 neutrons" is a magic number, because neutrons in a nucleus are in filled shells. According to the plot from the lecture (or the picture below)

the next state above "82" is 1h<sub>7/2</sub>

