

Transform to Frequency

per solid angle, as a function of time, the total energy

$$= |\vec{A}(t)|^2 \quad \text{where } \vec{A}(t) :=$$

over time, the total energy

$$= \int_{-\infty}^{\infty} |A(t)|^2 dt = \int_{-\infty}^{\infty}$$

g the integral over t_e , we

$$\int_{-\infty}^{\infty} e^{i\omega(t_e + R(t_e)/c)} \left[\hat{n} \right]$$

there are not references to describe e .

$\hat{n} \cdot \vec{r}(t)$, where observer is in region where $\dot{\vec{\beta}} \neq 0$, which is R . Then

we

$$\int_{-\infty}^{\infty} e^{i\omega R/c} \int_{-\infty}^{\infty} e^{i\omega(t - \hat{n} \cdot \vec{r}(t)/c)}$$

useful to integrate by parts to discuss the low frequency limit, that this is useful as we integrate by parts, assuming that boundary terms can be discarded, and inserti

$$\frac{\omega^2}{\pi^2 c} \left| \int_{-\infty}^{\infty} e^{i\omega(t - \hat{n} \cdot \vec{r}(t)/c)} \hat{n} \right.$$

We will skip pages 6

We can define the energy per unit solid angle per unit frequency,

$$\frac{d^2 I}{d\omega d\Omega} = 2 |\vec{A}(\omega)|^2.$$

Our expression for the radiative part of the electric field,

$$R \vec{E}(t) = \frac{q}{c} \frac{\hat{n} \times ((\hat{n} - \vec{\beta}) \times \dot{\vec{\beta}})}{(1 - \hat{n} \cdot \vec{\beta})^3} \Bigg|_{t_e}.$$

In calculating $d^2 I / d\omega d\Omega$ the phase factor $e^{i\omega R/c}$ will be irrelevant.

We note that the piece in the integrand multiplying the exponential can be written as a total time derivative:

$$\begin{aligned} \frac{d}{dt} \left[\frac{\hat{n} \times (\hat{n} \times \vec{\beta})}{1 - \hat{n} \cdot \vec{\beta}} \right] &= \frac{\hat{n} \times (\hat{n} \times \dot{\vec{\beta}})}{1 - \hat{n} \cdot \vec{\beta}} + \frac{\hat{n} \times (\hat{n} \times \vec{\beta})(\hat{n} \cdot \dot{\vec{\beta}})}{(1 - \hat{n} \cdot \vec{\beta})^2} \\ &= \frac{[(\hat{n} \cdot \dot{\vec{\beta}})\hat{n} - \dot{\vec{\beta}}](1 - \hat{n} \cdot \vec{\beta}) + [(\hat{n} \cdot \vec{\beta})\hat{n} - \vec{\beta}](\hat{n} \cdot \dot{\vec{\beta}})}{(1 - \hat{n} \cdot \vec{\beta})^2} \end{aligned}$$

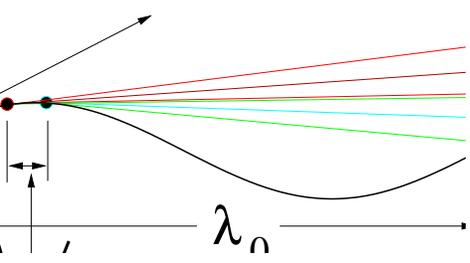
Wigglers and Undulators

The intense peaking of forward radiation from ultrarelativistic particles, and the blue-shifting thereof, is useful for condensed matter experimentalists and biologists who could make use of very intense short pulses of X-rays. Old high-energy accelerators needn't die, they become light-sources. Monochromatic sources would also be useful.

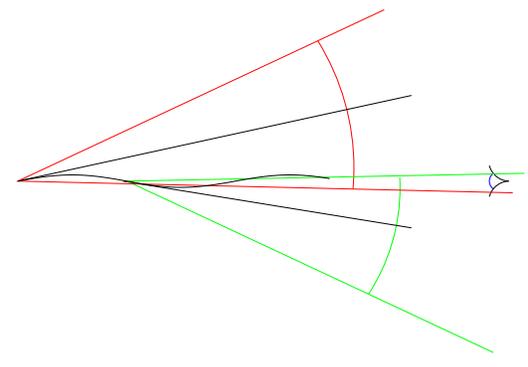
Wigglers and Undulators use a periodic sequence of alternately directed transverse magnets to produce transverse sinusoidal oscillations, $x = a \sin 2\pi z / \lambda_0$. The angle of the beam will vary by $\psi_0 = \Delta\theta = \frac{dx}{dz} = \frac{2\pi a}{\lambda_0}$.

The spread in angle of the forward radiation is $\theta_r \approx 1/\gamma$, centered on the momentary direction of the beam

the observer sees the source, that frequency a pulse to our eye only for one period, so $\Delta t_e \approx (\lambda_0 / c)$



intense region of the beam, and the beam is radiating coherently. In the particle's rest frame the disturbing fields have a Fitzgerald-contracted wavelength λ_0 / γ , going by at βc , so the particle sees itself oscillating at $\omega' = 2\pi c \gamma \beta / \lambda_0 \approx 2\pi c \gamma / \lambda_0$.



Thomson Scattering

the frequency observed

$$\omega = \frac{2\omega'}{\gamma(1 - \hat{n} \cdot \vec{\beta})} = \frac{4\pi}{\lambda_0(1 - \hat{n} \cdot \vec{\beta})}$$

coherent radiation, so the number of photons is N^2 and the frequency is $1/N$

electron has an electric field

$$\vec{E}(\vec{x}, t) = \vec{\epsilon}_0 E_0 e^{i\vec{k} \cdot \vec{x} - i\omega t}$$

it will have an acceleration

$$\dot{\vec{v}}(t) = \vec{\epsilon}_0 \frac{e}{m} E_0 e^{i\vec{k} \cdot \vec{x} - i\omega t}$$

is sufficiently limited to keep the particle non-relativistic ($\omega^2 \ll \lambda = 2\pi c / \omega$), the time-averaged cross section $\langle \dot{\vec{v}}^* \cdot \dot{\vec{v}} \rangle (\dot{\vec{v}}^* \cdot \vec{\epsilon})$ is

$$\frac{1}{2} \frac{e^2 |E_0|^2}{m^2} |\vec{\epsilon}^* \cdot \vec{\epsilon}_0|^2$$

We saw (14.18) that in the particle's rest frame the electric field is given by

$$\vec{E} = \frac{q}{c^2 R} \hat{n} \times (\hat{n} \times \dot{\vec{v}}),$$

so the amplitude corresponding to a particular polarization vector $\vec{\epsilon}$ is

$$\vec{\epsilon}^* \cdot \vec{E} = \frac{q}{c^2 R} \vec{\epsilon}^* \cdot (\hat{n} \times (\hat{n} \times \dot{\vec{v}})) = \frac{q}{c^2 R} \vec{\epsilon}^* \cdot \dot{\vec{v}},$$

Dividing this by the incident energy flux $c|E_0|^2 / 8\pi$ we get the cross section

$$\frac{d\sigma}{d\Omega} = \left(\frac{e^2}{mc^2} \right)^2 |\vec{\epsilon}^* \cdot \vec{\epsilon}_0|^2.$$

If the scattering angle is θ and the incident beam is unpolarized and the cross section summed over final polarizations, the factor of

$$\begin{aligned} & \frac{1}{2} \sum_i \sum_f |\vec{\epsilon}_f^* \cdot \vec{\epsilon}_i|^2 \\ &= \frac{1}{2\pi^2} \int_0^{2\pi} d\phi_i \int_0^{2\pi} d\phi_f \\ & \quad [(\cos \theta \cos \phi_f, \sin \phi_f, -\sin \theta \cos \phi_f) \cdot (\cos \phi_i, \sin \phi_i, 0)]^2 \\ &= \frac{1}{2\pi^2} \int_0^{2\pi} d\phi_i \int_0^{2\pi} d\phi_f [(\cos \theta \cos \phi_f \cos \phi_i + \sin \phi_f \sin \phi_i)]^2 \end{aligned}$$

polarized cross sections

$$\frac{d\sigma}{d\Omega} = \left(\frac{e^2}{mc^2} \right)^2 \frac{1 + \cos^2 \theta}{2}$$

and the **Thomson formula** for the cross section is

$$\sigma_T = \frac{8\pi}{3} \left(\frac{e^2}{mc^2} \right)^2$$

where r_e in parentheses is called the **classical electron radius**, roughly the radius of a sphere of charge e would have if $U = mc^2$. (The factor of $1/4\pi\epsilon_0$ for a charged sphere, is discarded

cross section could have been measured with an arbitrarily weak field, so recoil could be neglected, but quantum-mechanically the minimum energy hitting the electron is $\hbar\omega$, which gives a significant recoil if $\hbar\omega \approx mc^2$. In fact, if we take quantum mechanics into account we are considering Compton scattering, for which, we learned as a freshman, energy and momentum conservation insure that the outgoing photon has an increased wavelength,

$$\lambda' = \lambda + \frac{h}{mc}(1 - \cos \theta), \quad \text{or} \quad \frac{k'}{k} = \frac{1}{1 + \frac{\hbar\omega}{mc^2}(1 - \cos^2 \theta)}.$$

It turns out that the quantum mechanical calculation (for a scalar particle) is the classical result times $(k'/k)^2$:

$$\left. \frac{d\sigma}{d\Omega} \right|_{\text{QM, scalar}} = \left(\frac{e^2}{mc^2} \right)^2 \left(\frac{k'}{k} \right)^2 |\vec{\epsilon}^* \cdot \vec{\epsilon}_0|^2.$$

