

Physics 406
2012

HW#8
solutions

① Omar Q. 2

Eq. (10.12):

$$H_c(0) \approx \left(\frac{2nk_B^2}{\mu_0 E_F} \right)^{1/2} T_c$$

Recall that in 3D,

$$E_F \sim n^{2/3}$$

Thus $H_c(0) \sim n^{1/6} T_c$, the dependence on n is very weak

$H_c(0) \sim T_c$ for any superconductor

② Omar Q. 6

The binding energy of a Cooper pair (i.e. the energy required to break it up) is roughly $\Delta(T)$, the gap energy. It is easier to break up the Cooper pair as $T \uparrow$ due to thermal fluctuations. Thus we expect $\Delta(T)$ to decrease with $T \uparrow$ until $T = T_c$. At $T \geq T_c$ Cooper pair formation is no longer energetically favorable.

9. Omar Q.5

3

London equation:

$$\vec{B} = -\frac{m}{n_s e^2} \vec{\nabla} \times \vec{J}_s, \text{ where}$$

$$\vec{J}_s = -n_s e \vec{v}_s.$$

$$\begin{aligned} B_y &= -\frac{m}{n_s e^2} \left(\frac{\partial J_{s,x}}{\partial z} - \frac{\partial J_{s,z}}{\partial x} \right) = \\ &= \frac{m}{n_s e^2} \frac{\partial J_{s,z}}{\partial x} \end{aligned}$$

But $B_y(x) = B_y(0) e^{-x/\lambda}$, so that

$$J_{s,z}(x) = \left(\frac{n_s e^2}{m} \right) (-\lambda) B_y(x) =$$

$$= - \left(\frac{n_s^2 e^4}{m^2} \frac{m}{\mu_0 n_s e^2} \right)^{1/2} B_y(x) =$$

$$= - \left(\frac{n_s e^2}{\mu_0 m} \right)^{1/2} B_y(x) \equiv -J_{s,z}(0) e^{-x/\lambda}$$

④ Ch. 10, Pr. 1

Solenoid:

$$L \frac{dI}{dt} + IR = 0,$$

$$\frac{dI}{I} = - \underbrace{\frac{R}{L}}_{\frac{1}{\tau}} dt \Rightarrow I(t) = I(0) e^{-\tau t},$$

where τ is the damping timescale

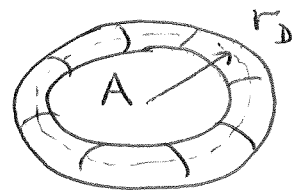
Current damped by 0.01%:

$$\frac{I}{I(0)} = 0.9999 = e^{-\tau \tau_c}, \text{ or}$$

$$\tau_c = - \frac{\log(0.9999)}{\tau}$$

$$\text{Now, } R = \rho \frac{l_w}{A} = \rho \frac{l_w}{\pi \left(\frac{d_w}{2}\right)^2}$$

$$L = \frac{\Phi}{I} = \frac{\mu_0 N^2 A}{25\pi r_D}$$



$$\mu_0 N I = B (25\pi r_D), \text{ or}$$

turns

$$B = \frac{\mu_0 N I}{25\pi r_D}$$

$$\text{Then } \Phi = NBA = \frac{\mu_0 N^2 A}{25\pi r_D} I$$

$$\text{So, } \frac{R}{L} = \frac{\rho l_w}{\pi \left(\frac{d_w}{2}\right)^2} \times \frac{25\pi r_D}{\mu_0 N^2 A}$$

Not all #'s are given \Rightarrow leave in symbolic form

⑤ 0 Ch. 10, Pr. 6

Perfect diamagnetism is considered the more fundamental property of a superconductor.

Indeed, $B=0$ inside a SC ($T < T_c$) does not follow from perfect conductivity:

$$\text{Ohm's law } \vec{E} = \rho \vec{j}$$

$\rho \rightarrow 0$ but j finite:
 $\vec{E} = 0$

Maxwell's equation

$$\vec{\nabla} \times \vec{E} = \frac{\partial \vec{B}}{\partial t} \Rightarrow \frac{\partial \vec{B}}{\partial t} = 0 \text{ since } \vec{E} = 0.$$

Thus a perfect conductor implies that $\vec{B} = \text{const}$, not $\vec{B} = 0$ as observed below T_c (Meissner effect). In other words, just a perfect conductor cannot have a change in magnetic flux after cooling through T_c .

⑥ Key points on the intermetallic superconductor MgB_2 .

a) Original idea: investigation of ternary compounds with 2 light atoms (Mg & B) and one intermediate one (Ti). Expected large DoS \Rightarrow large T_c .

b) Discovered that $T_c \sim 40K$ is actually due to binary compound MgB_2 .

c) MgB_2 wires can be made relatively easily \Rightarrow practical applications.

d) Observation of the isotope effect in MgB_2 \Rightarrow role of e -phonon interactions.

e) Combination of high critical field & low normal-state resistivity make MgB_2 attractive for making SC magnets.

f) Mechanism of SC unknown \Rightarrow two superconducting gaps?