Rutgers University Department of Physics & Astronomy

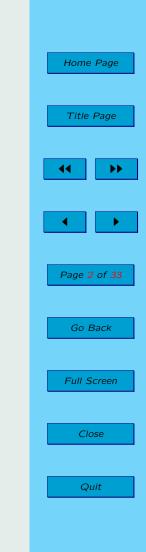
01:750:271 Honors Physics I Fall 2015

Lecture 21



Final Exam:

- Wednesday, Dec 21st, 8:00-11:00am in PHL.
- Final make up: Thursday, Dec 22nd, 10:00am 1:00pm in Serrin (physics building) E372.
 - Final: 20 questions = 25% final score



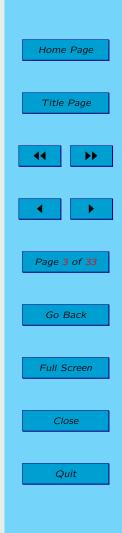
13. Gravitation

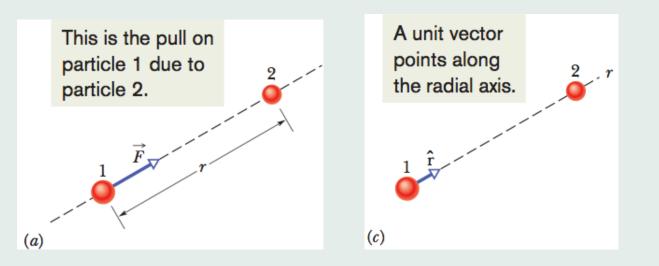
Newton's law of gravitation

Every point particle attracts every other particle with a gravitational force

$$F = G \frac{m_1 m_2}{r^2}$$

$$G = 6.67 \times 10^{-11} \,\mathrm{N \cdot m^2/kg^2} = 6.67 \times 10^{-11} \,\mathrm{m^3/kg \cdot s^2}$$

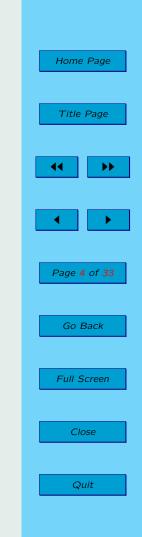




Gravitational force on particle 1 - vector form:

$$\vec{F}_{2 \text{ on } 1} = G \frac{m_1 m_2}{r^2} \hat{r}$$

 \hat{r} is the radial **unit** vector: $|\vec{r}| = 1$.



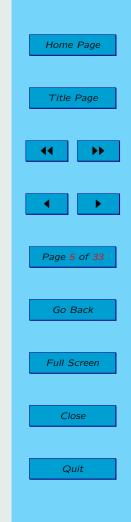
• Principle of Superposition

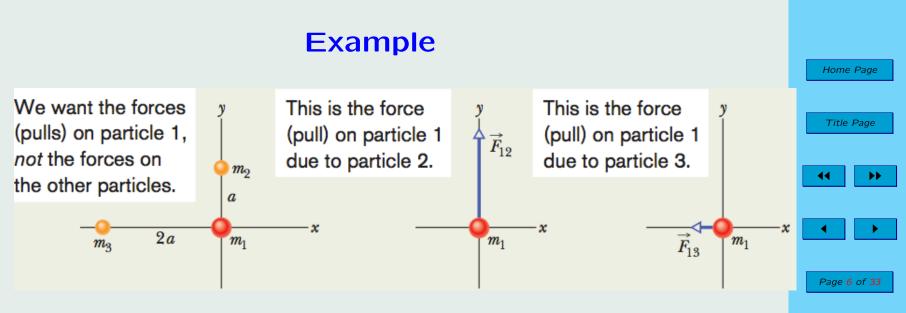
Given n interacting particles:

$$\vec{F}_{1,\text{net}} = \vec{F}_{12} + \vec{F}_{13} + \dots + \vec{F}_{1n}$$

 $\vec{F}_{1,\text{net}} =$ net gravitational force acting on particle 1

 $\vec{F}_{1,i} = \text{gravitational force on particle 1 from particle } i$





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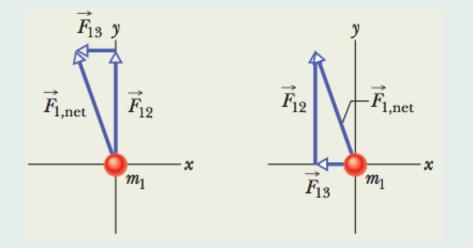
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• Isolated system of particles far from other massive objects.

• What is the magnitude of the gravitational force on particle 1 ?

$$ec{F}_{1,{\sf net}} = ec{F}_{12} + ec{F}_{13}$$



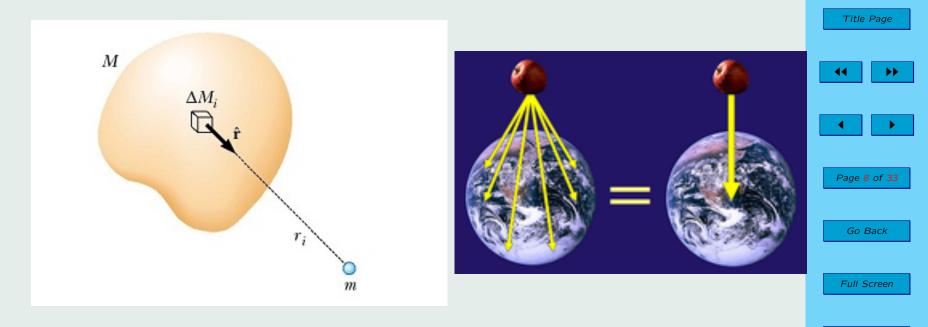
$$\vec{F}_{12} = G \frac{m_1 m_2}{a^2} \hat{j}$$

$$\vec{F}_{13} = -G\frac{m_1m_3}{4a^2}\hat{i}$$

$$\vec{F}_{1,\text{net}} = \frac{Gm_1}{a^2} \left(m_2 \hat{j} - \frac{m_3}{4} \hat{i} \right)$$
$$|\vec{F}_{1,\text{net}}| = \frac{Gm_1}{a^2} \sqrt{m_2^2 + \frac{m_3^2}{16}}$$

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• Gravitational force on a particle from an extended object:



$$\vec{F} = \int d\vec{F} = \int -G \frac{m dM}{r^2} \hat{r}$$

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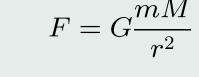
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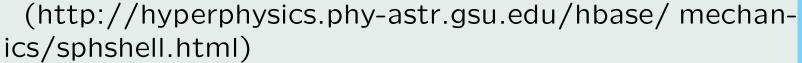
Shell theorem

A uniform spherical shell of matter attracts a particle that is outside the shell as if all the shell's mass were concentrated at its center.

F



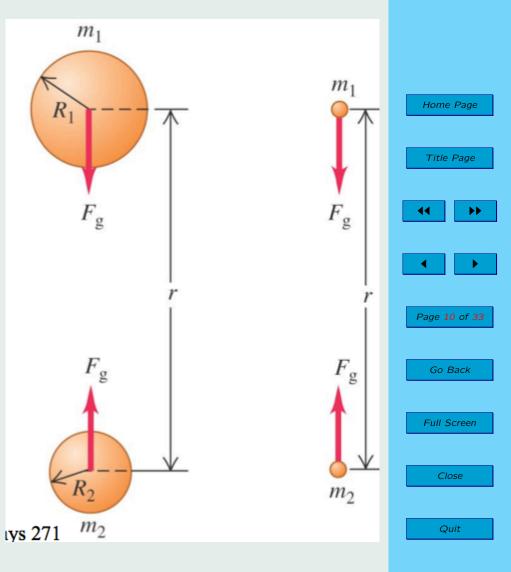
- \bullet *m* mass of particle
- M mass of shell
- r distance from particle to center of spherical shell



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Consequence:

The gravitational force between two **uniform** spherical distributions of mass is the same as if all the mass of each sphere were concentrated at its center



• Gravitation near Earth's surface

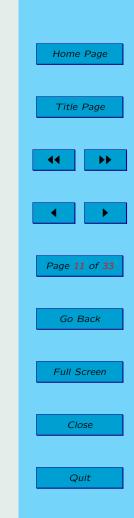


- Assume the Earth is exactly spherical.
- Assume uniform distribution of mass throughout the Earth
- Neglect rotation effects.

Gravitational force on a particle near Earth's surface:

$$\vec{F}_g = -G \frac{mM}{r^2} \hat{r} \Rightarrow a_g = \frac{GM}{r^2}$$

r = distance between particle and center of the Earth



Note: the free fall acceleration decreases with r

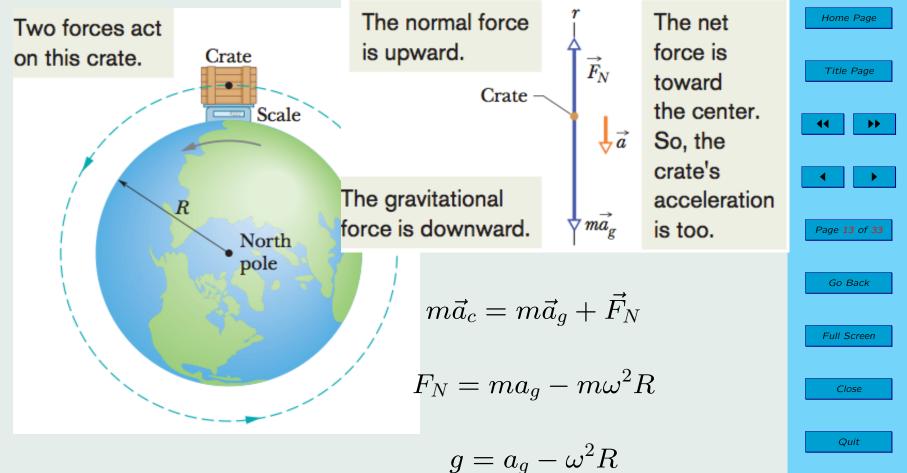
Altitude (km)	a_g (m/s ²)	Altitude Example
		Mean Earth
0	9.83	surface
8.8	9.80	Mt. Everest
36.6	9.71	Highest crewed balloon
400	8.70	Space shuttle orbit
35 700	0.225	Communications satellite

$$a_g = \frac{GM}{r^2} = \frac{GM}{(R+h)^2}$$

h = altitude = distance from particle to the surface of the Earth



• Rotation effects



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The gravitational acceleration on the surface of the Earth is g (neglecting rotation.) What will it be on the surface of a planet that has half the mass of the Earth and half its radius?

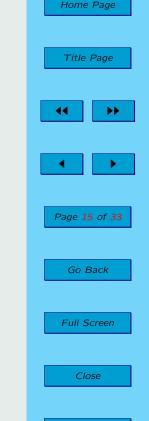
A) g/4
B) g/2
C) g
D) 2g



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The gravitational acceleration on the surface of the Earth is g (neglecting rotation.) What will it be on the surface of a planet that has half the mass of the Earth and half its radius?

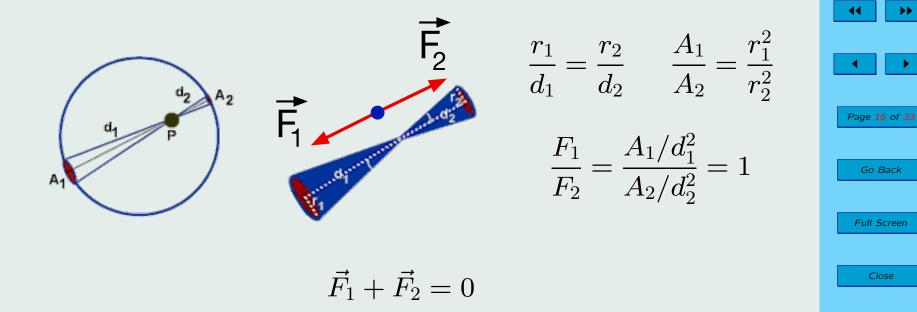
A) g/4B) g/2C) gD) 2g $g = G\frac{M}{R^2} \Rightarrow g'/g = M'R^2/M(R')^2 = 4/2 = 2$



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Gravitation inside the Earth

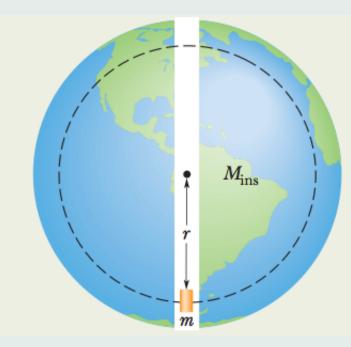
A uniform shell of matter exerts no net gravitational force on a particle located inside it.



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Suppose a very narrow tunnel is dug through the Earth along the North-South axis.

What is the gravitational force on a particle of mass m at distance r < R from the center?

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Any thin spherical shell of matter of radius $r_{\rm shell} > r$ does not yield any net force.

Only thin shells of radius $r_{\rm shell} < r$ give a nonzero force.

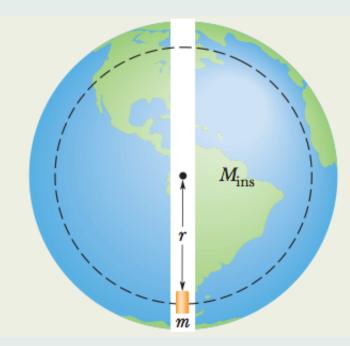
The gravitational force on the particle is the same as the force due to a sphere of radius r.

Assume spherical shape and uniform mass distribution. Mass density

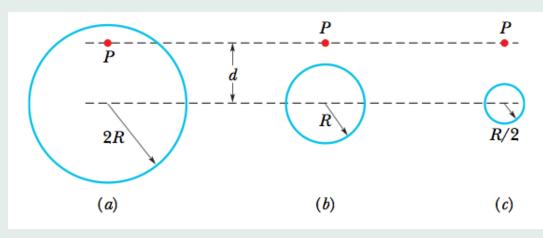
$$\rho = \frac{M}{4\pi R^3/3} = \frac{3M}{4\pi R^3}$$

 $F = \frac{Gm}{r^2} \times \text{Mass inside sphere of radius } r$ $= \frac{Gm}{r^2} \times \left(\frac{4\pi r^3 \rho}{3}\right) = \frac{4\pi Gm \rho}{3} r = \frac{GmM}{R^3} r$

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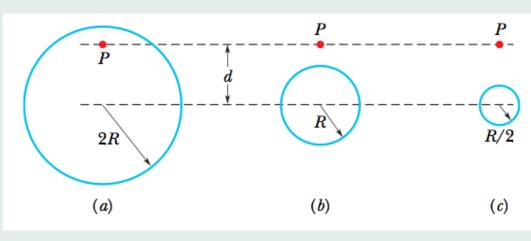


A) $F_a > F_b > F_c$ B) $F_a = F_b > F_c$ C) $F_a < F_b < F_c$ D) $F_a < F_b = F_c$ • All spherical shells are uniform and have the same mass *M*.

Rank the situations
 according to
 the magnitude of the
 gravitational
 force on the
 particle.

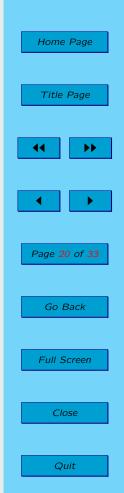
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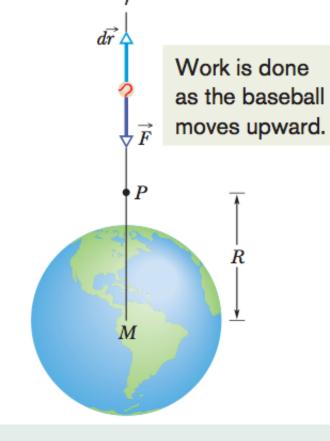


A) $F_a > F_b > F_c$ B) $F_a = F_b > F_c$ C) $F_a < F_b < F_c$ D) $F_a < F_b = F_c$ $F_a = 0, F_b = F_c = GmM/r^2$ • All spherical shells are uniform and have the same mass M.

Rank the situations
according to
the magnitude of the gravitational
force on the particle.

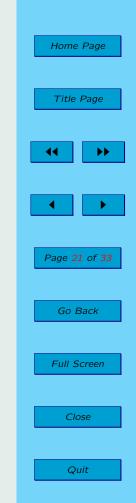


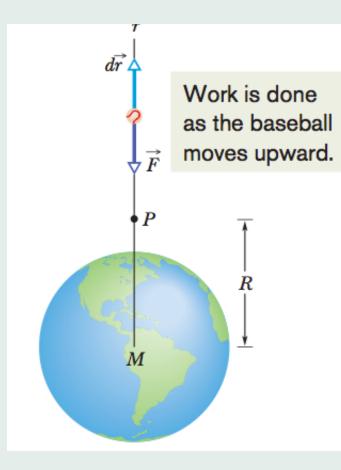
• Gravitational potential energy



Work done by gravitational force: suppose a baseball moves upward from a distance r_1 to a distance $r_2 >$ r_1 in the gravitational field of the Earth.

What is the work done by gravitational force?

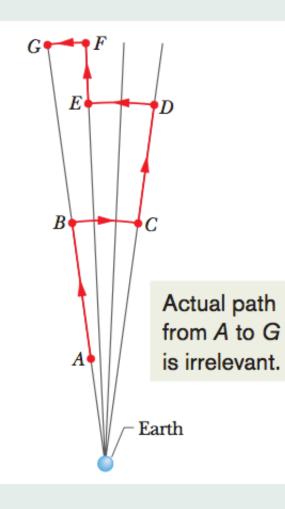




$$egin{aligned} W_{F_g} &= \int_{r_1}^{r_2} ec{F_g} \cdot dec{r} \ &= -GmM \int_{r_1}^{r_2} rac{dr}{r^2} \ &= GmM \left(rac{1}{r_2} - rac{1}{r_1}
ight) \end{aligned}$$

Does it depend on the path?

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 W_{F_g} does **not** depend on the path.

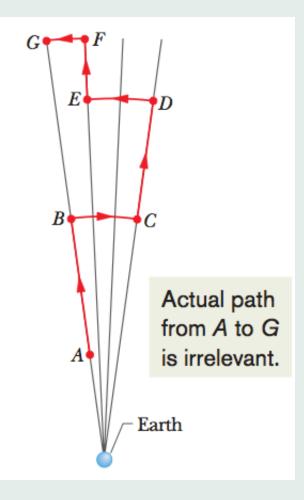
Contribution of **circular** arcs is 0:

$$\vec{F}_g \cdot d\vec{r} = 0$$

Contribution of radial arcs: $\vec{F_g} \cdot d\vec{r} = -\frac{GmM}{r^2}dr$

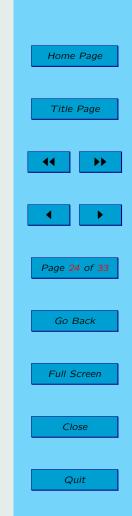
 $W_{F_g, ext{curved path}} = W_{F_g, ext{radial path}}$

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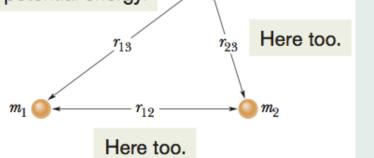


- The gravitational force is conservative
- Gravitational potential energy:

$$U(\infty) - U(r) = -W_{F_g,r o \infty} = rac{GmM}{r}$$



Gravitational potential energy: system of particles • For any pair of particles (i, j) $U_{ij} = -\frac{Gm_im_j}{U_{ij}}$

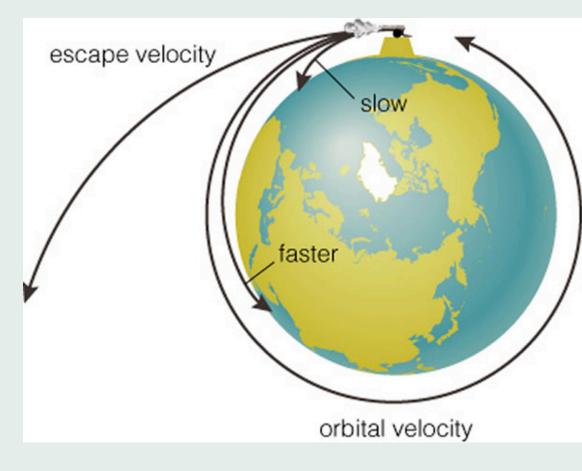


$$U_{\text{total}} = -\sum_{i < j} \frac{Gm_i m_j}{r_{ij}}$$

$$U_{\text{total}} = -\left(\frac{Gm_1m_2}{r_{12}} + \frac{Gm_2m_3}{r_{23}} + \frac{Gm_1m_3}{r_{13}}\right)$$

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How fast should а projectile be launched such that it escapes Earth's gravitational field? fast How should it be launched such that does not it fall back on Earth?

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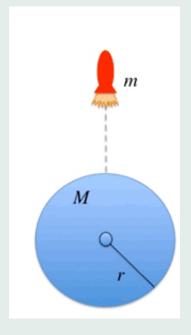
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i-Clicker



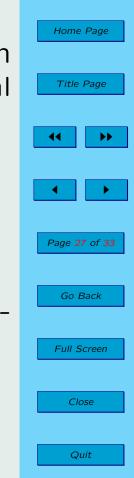
Find the initial speed of the rocket such that it escapes the Earth's gravitational field, moving infinitely far away.

A)
$$\sqrt{MG/R}$$

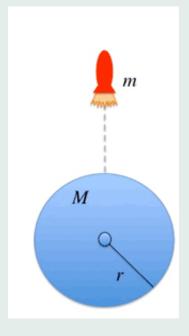
B)
$$\sqrt{2MG/R}$$

C)
$$\sqrt{MG/2R}$$

 $D)\ {\rm It}$ is impossible for the rocket to escape.



i-Clicker



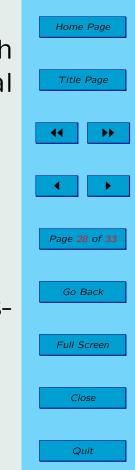
Find the initial speed of the rocket such that it escapes the Earth's gravitational field, moving infinitely far away.

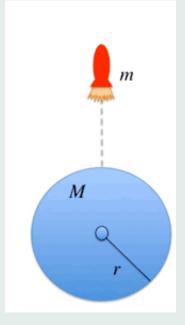
$$A) \sqrt{MG/R}$$

B)
$$\sqrt{2MG/R}$$

C)
$$\sqrt{MG/2R}$$

 $D)\ {\rm It}$ is impossible for the rocket to escape.





• Energy conservation

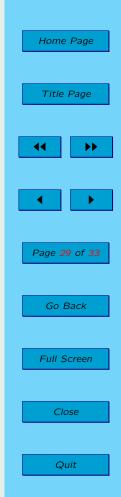
Initially: object on the surface of the Earth; kinetic and potential energy

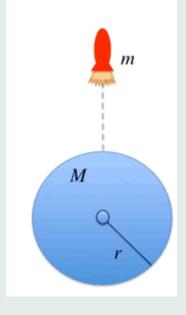
 $E_{mec} = K_0 + U_{\rm grav}$

Finally: object at ∞ ; kinetic energy

$$E_{mec} = K \ge 0$$

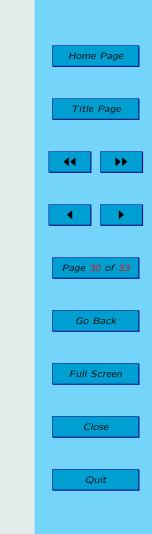
$$\frac{1}{2}mv_0^2 - \frac{GmM}{R} = \frac{1}{2}mv^2 \ge 0 \implies v_0 \ge \sqrt{\frac{2GM}{R}}$$



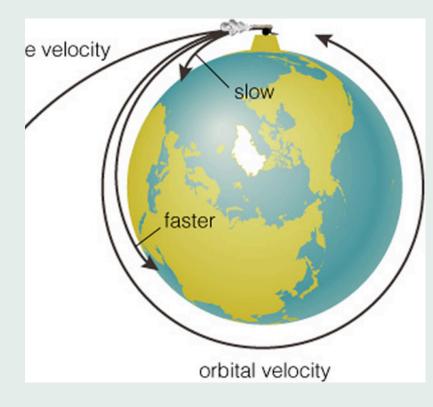


 $v_{\rm esc} = \sqrt{\frac{2GM}{R}}$

Escape speed



• i-Clicker



A projectile is launched horizontally from altitude *h* above the Earth's surface. How fast should a projectile be launched such that it moves on a circular trajectory around the Earth?

$$A) \sqrt{MG/(R+h)}$$

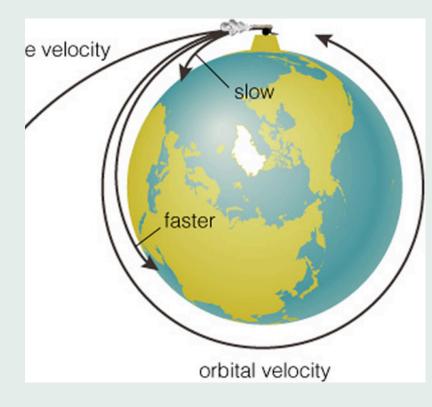
B)
$$\sqrt{2MG/(R+h)}$$

C) $\sqrt{MG/2(R+h)}$

D) It is impossible for the projectile to circle the Earth.

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• i-Clicker



A projectile is launched horizontally from altitude *h* above the Earth's surface. How fast should a projectile be launched such that it moves on a circular trajectory around the Earth?

A) $\sqrt{MG/(R+h)}$

$$B) \sqrt{2MG/(R+h)}$$

C) $\sqrt{MG/2(R+h)}$

D) It is impossible for the projectile to circle the Earth.

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$$m\vec{a}_c = \vec{F}_g$$

$$\frac{mv^2}{R+h} = \frac{GmM}{(R+h)^2}$$
$$v = \sqrt{\frac{GM}{R+h}}$$

Orbital velocity

