

(low wavelengths). For example, a black body at room temperature (300 K) with one square meter of surface area will emit a photon in the visible range (390-750 nm) at an average rate of one photon every 41 seconds, meaning that for most practical purposes, such a black body does not emit in the visible range [1]

### Blackbody simulators

Although a black body is a theoretical object (i.e. emissivity  $\epsilon = 1.0$ ), common applications define a source of infrared radiation as a black body when the object approaches an emissivity of 1.0, (typically  $\epsilon = 0.99$  or better). A source of infrared radiation less than 0.99 is referred to as a "grey body" [11]. Applications for black body simulators typically include the testing and calibration of infrared systems and infrared sensor equipment.

Super black is an example of such a material, made from a nickel-phosphorus alloy. More recently, a team of Japanese scientists created a material even closer to a black body, based on vertically aligned single-walled carbon nanotubes, which absorbs between 98% and 99% of the incoming light, in the spectral range from UV to far infrared [11].

### Equations governing black bodies

#### Planck's law of blackbody radiation

Main article: Planck's law

Planck's law states that

$$I(\nu, T) d\nu = \left( \frac{2h\nu^3}{c^2} \right) \frac{1}{e^{h\nu/kT} - 1} d\nu$$

where

$I(\nu, T) d\nu$  is the amount of energy per unit surface area per unit time per unit solid angle emitted in the frequency range between  $\nu$  and  $\nu + d\nu$  by a black body at temperature  $T$ .

$h$  is the Planck constant.

$c$  is the speed of light in a vacuum.

$k$  is the Boltzmann constant.

$\nu$  is frequency of electromagnetic radiation, and

$T$  is the temperature in kelvins.

#### Wien's displacement law

Main article: Wien's displacement law

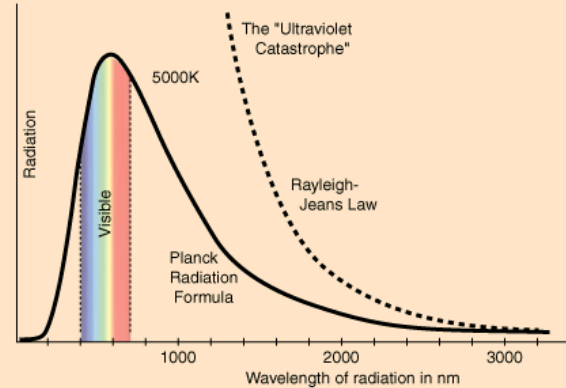
Wien's displacement law shows how the spectrum of black body radiation at any temperature is related to the spectrum at any other



A typical industrial "extended source plate" type black body.

## Rayleigh-Jeans vs Planck

Comparison of the classical Rayleigh-Jeans Law and the quantum [Planck radiation formula](#). Experiment confirms the Planck relationship.



### Comments on the development of the Rayleigh-Jeans Law

[Why does the Planck curve drop below the Rayleigh-Jeans?](#)

[Frequency plot](#)

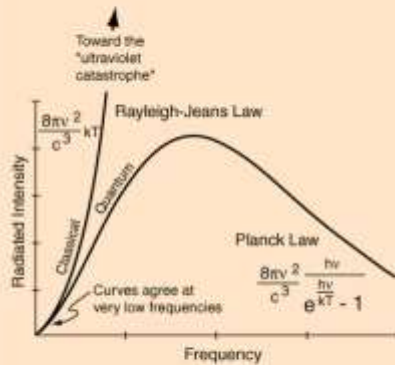
[Show that the two agree for long wavelengths](#)

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## Blackbody Intensity as a Function of Frequency



The Rayleigh-Jeans curve agrees with the [Planck radiation formula](#) for long wavelengths, low frequencies.

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Energy per unit volume per unit frequency

$$S_\nu = \frac{8\pi h}{c^3} \frac{\nu^3}{e^{h\nu/kT} - 1}$$

Energy per unit volume per unit wavelength

$$S_\lambda = \frac{8\pi hc}{\lambda^5} \frac{1}{e^{hc/\lambda kT} - 1}$$

[Example](#)