

Physics 343 Lecture # 6:
Bayesian methods & Fourier transforms

This week's schedule

Tonight 11:59pm: report for lab # 2 due by email (PDF please)

Monday – Thursday: hands-on sessions for lab # 3 (using ds9 and IDL to analyze archival millimeter array data)
+ attendance mandatory; active participation counts towards your course grade

Monday (Baker) & Thursday (Wu): regular office hours

Next week: “on call” office hours for lab # 3

NGC 1068

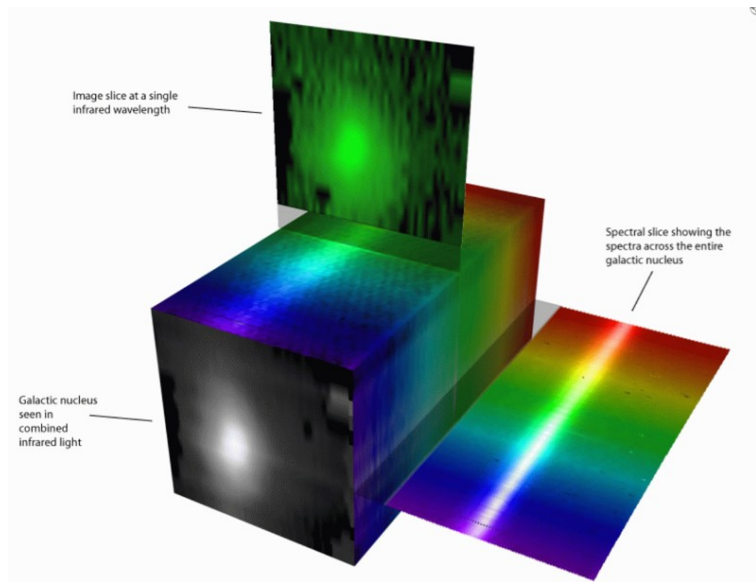
**Nearby spiral galaxy
with a powerful AGN
and lots of molecular
gas (traced by CO,
color-coded blue here).**

**Image credit: ALMA
(ESO/NAOJ/NRAO),
S. Takano et al.,
NASA/ESA *Hubble*
Space Telescope,
A. van der Hoeven**



Data for lab 3 come in the form of a *cube*

An observation with an array of radio telescopes will yield a **data “cube”** with two spatial dimensions and one frequency = redshift = velocity dimension.



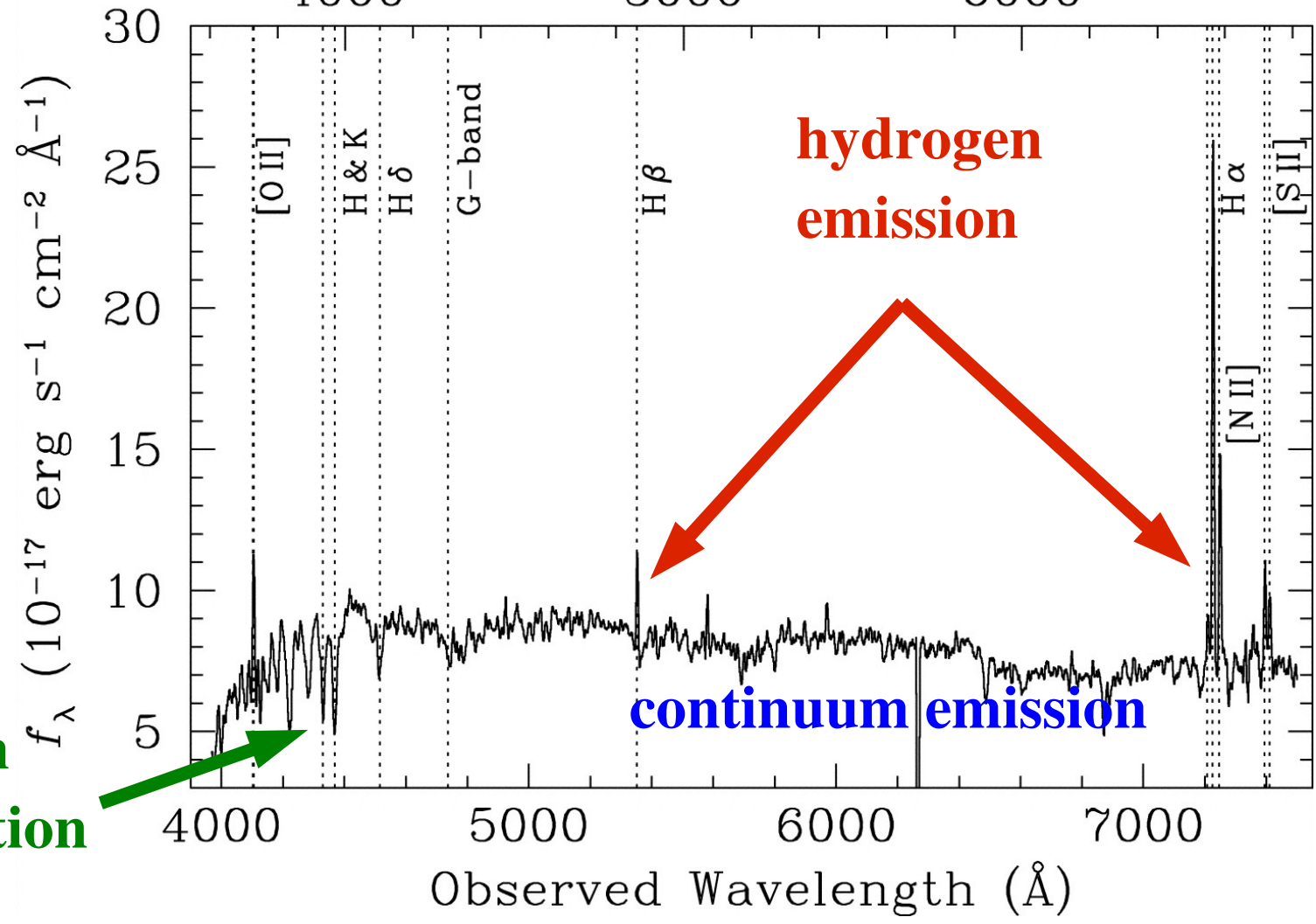
NGC1068 may have a **combination** of CO(1–0) rotational line emission (tracing molecular gas) and millimeter wavelength continuum emission (tracing free-free and/or synchrotron emission), in which case you will need to **separate** them within the data cube.

S. Todd & D. Pierce-Price

Each spatial pixel has own line+ctm spectrum

Rest-frame Wavelength (at $z_{\text{DLA}}=0.10104$)

4000 5000 6000



calcium absorption

hydrogen emission

continuum emission

Observed Wavelength (\AA)

Quiz