

**Physics 343 Lecture # 4:
gas dynamics (and more statistics)**

This week's schedule

Tonight 11:59pm: report for lab # 1 due by email (PDF please)

Monday – Thursday: hands-on sessions for lab # 2 (using archival SRT data in Excel format)

+ attendance mandatory; active participation (this week = getting a head start on your analysis) counts towards your course grade

Monday (Baker) & Thursday (Wu): regular office hours

Next week: “on call” office hours for lab # 2

Velocities in astronomy

Observed frequency and wavelength are related to **rest** (emitted) frequency and wavelength by a velocity (or redshift).

Exact relation = Doppler shift:

$$\nu_0/\nu = \lambda/\lambda_0 = 1 + z = \gamma (1 + v/c) \text{ for } \gamma = (1 - v^2/c^2)^{-1/2}$$

and for relative velocities, $\Delta\nu/\nu = \Delta\lambda/\lambda = \Delta z/(1 + z) = \Delta v/c$.

However, astronomers also make different approximations...

radio: $\nu \simeq \nu_0 (1 - v_{\text{rad}}/c)$

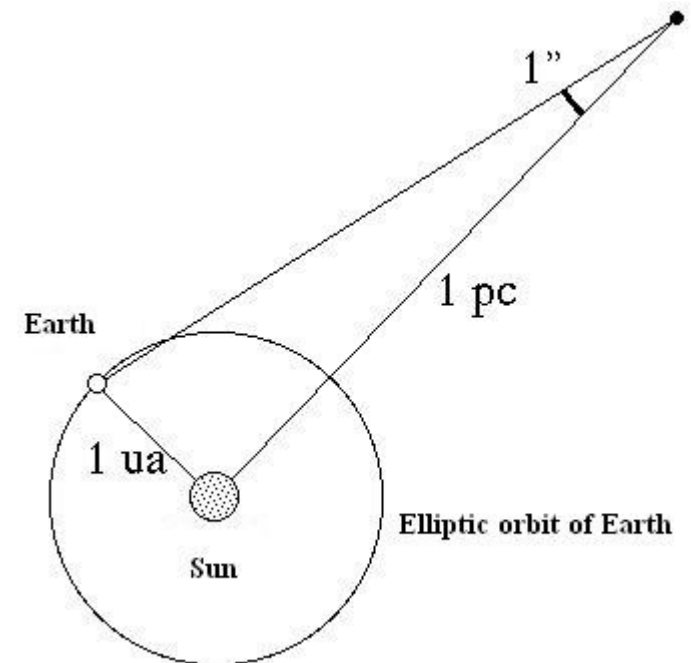
optical: $\lambda \simeq \lambda_0 (1 + v_{\text{opt}}/c)$

Distances in astronomy

Nearest stars can have distances measured by **parallax**:
the apparent shift in position relative to the background
pattern of more distant stars caused by the earth's
motion around the Sun.

1 parsec = 1 pc:
 3.089×10^{18} cm
~ 3.3 light years

Distances **inside** galaxies ~ kpc.
Distance **between** galaxies ~ Mpc.



Stellar components of spiral galaxies

Spiral galaxies have two principal components: **disk** and **bulge**.

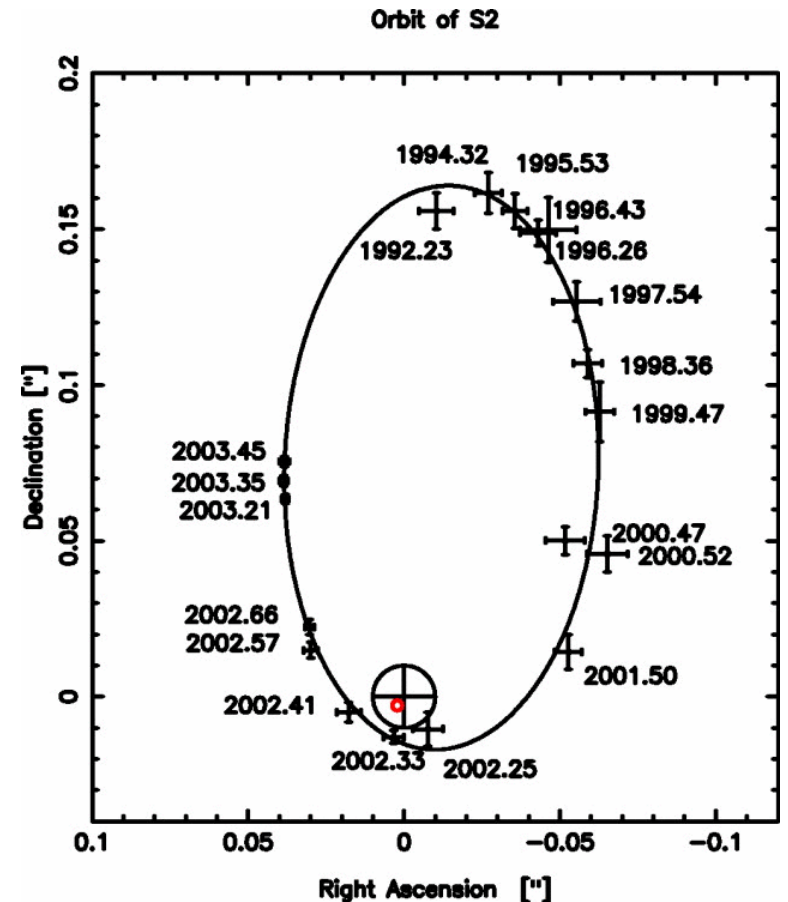


Where is the Sun within the Galaxy?

Note on terminology: Milky Way = “the Galaxy”; other “galaxies” are not capitalized.

Sun and solar system lie at a distance of **8 kpc** from the Galactic Center, where a supermassive black hole lies.

Eisenhauer et al. (2003)

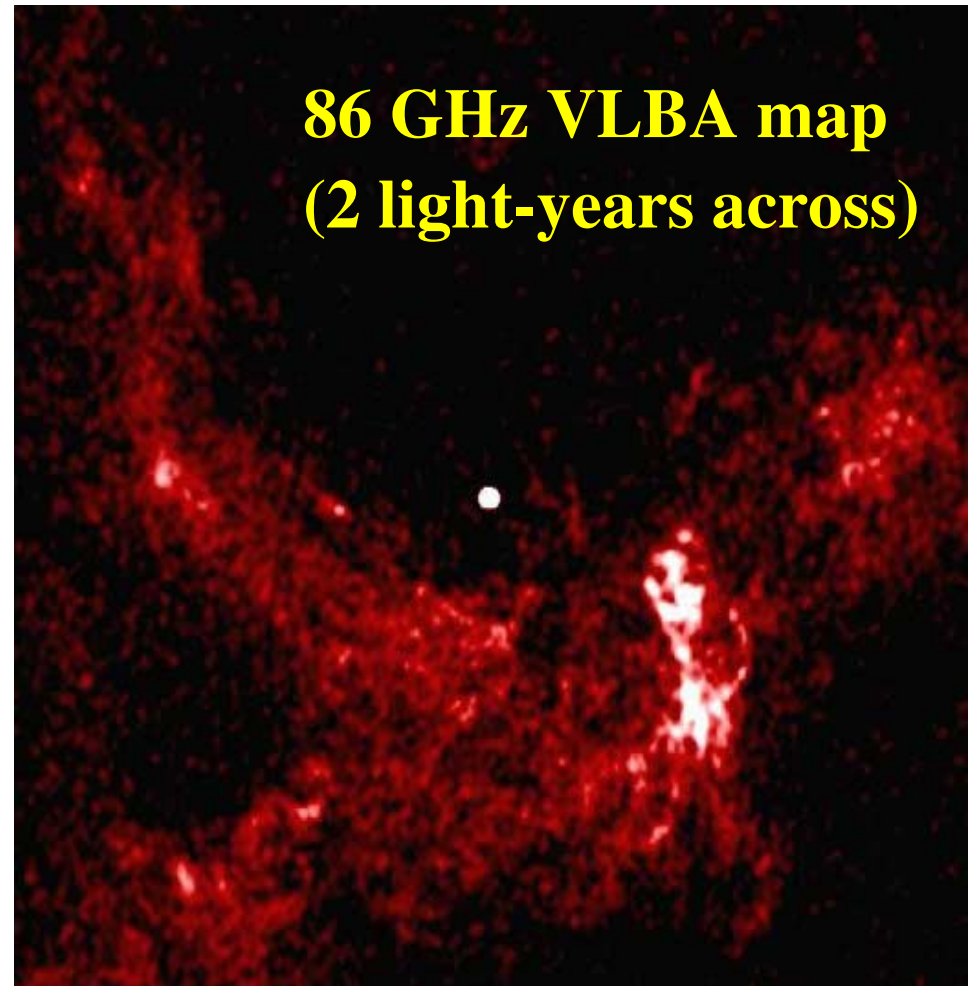


Central black hole is a radio source!

Motions of stars imply a large central mass – but from low proper motion of **Sgr A*** (“Sagittarius A–star”), we know that *it* must be massive.

Observations with the Very Long Baseline Array reveal a proper motion only due to the Sun's motion around the Galaxy.

Shen et al. (2005)



**86 GHz VLBA map
(2 light-years across)**

Galactic coordinates

The Sun is located in the disk.

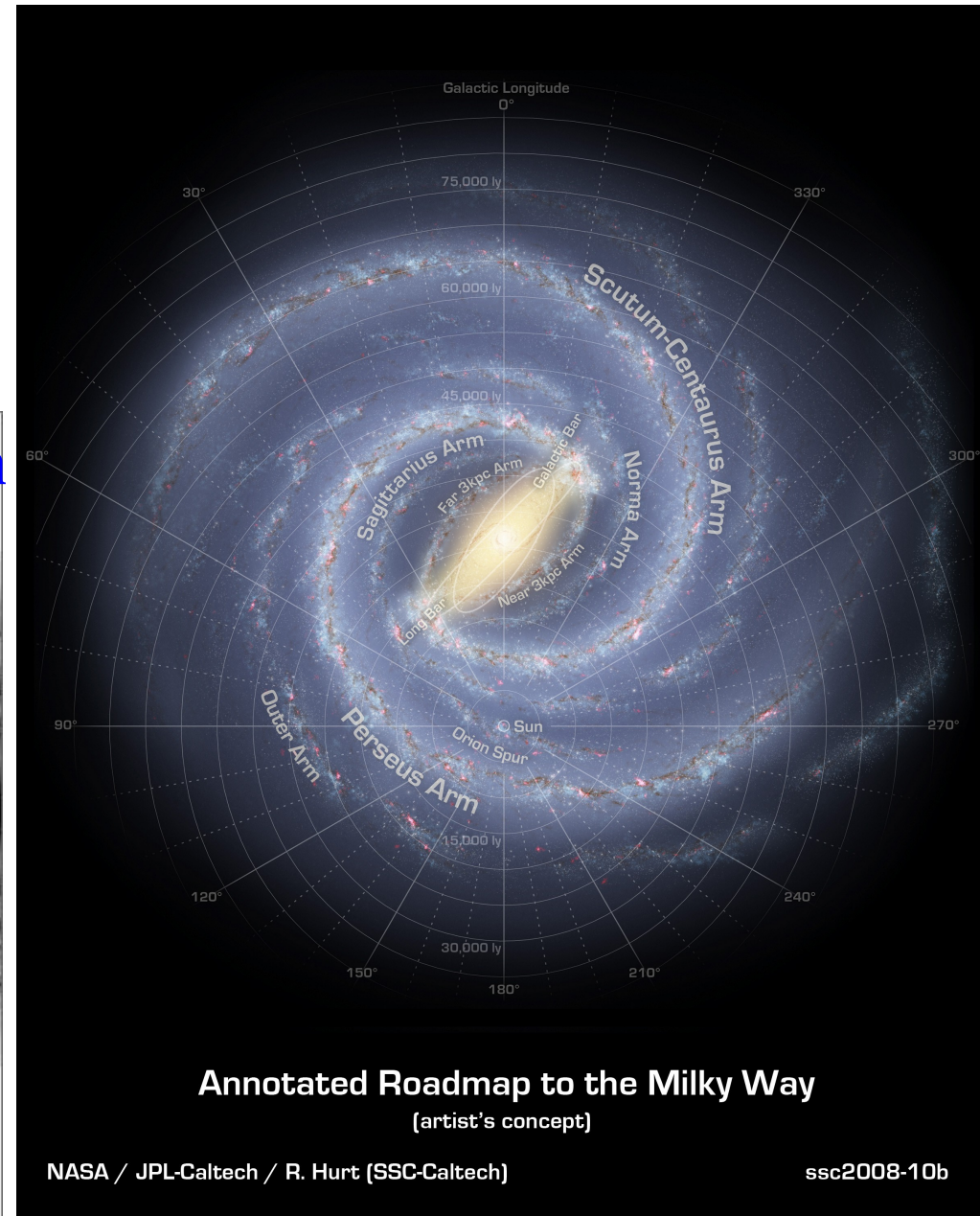
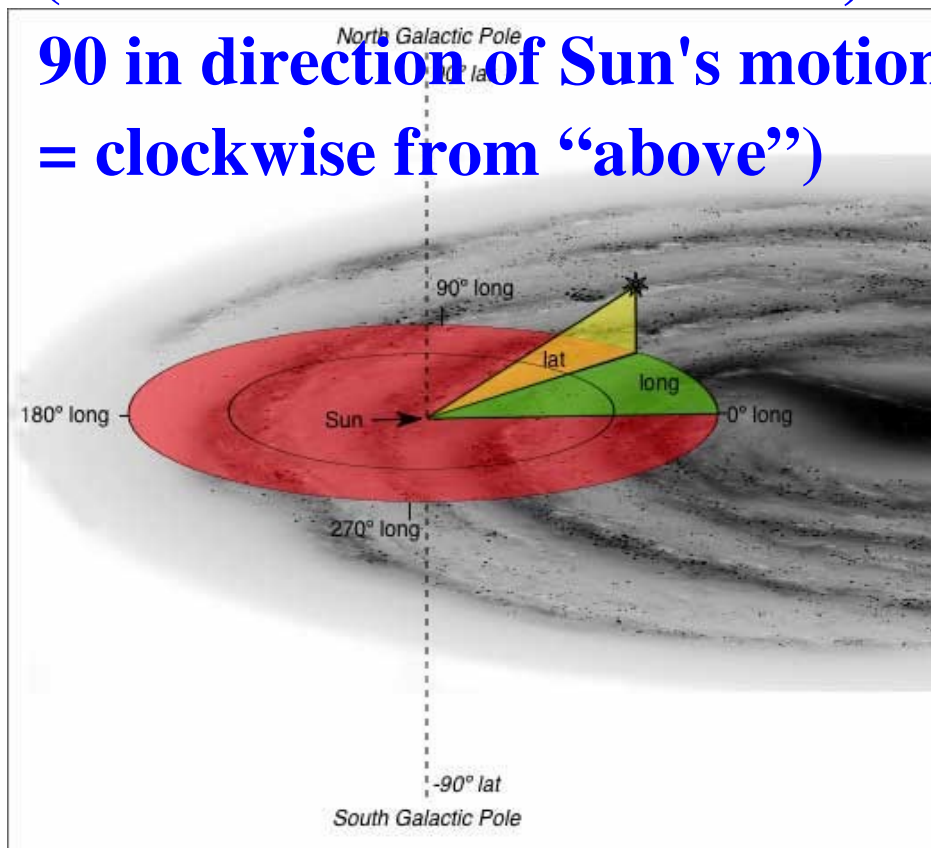
b = Galactic latitude

(above/below plane)

l = Galactic longitude

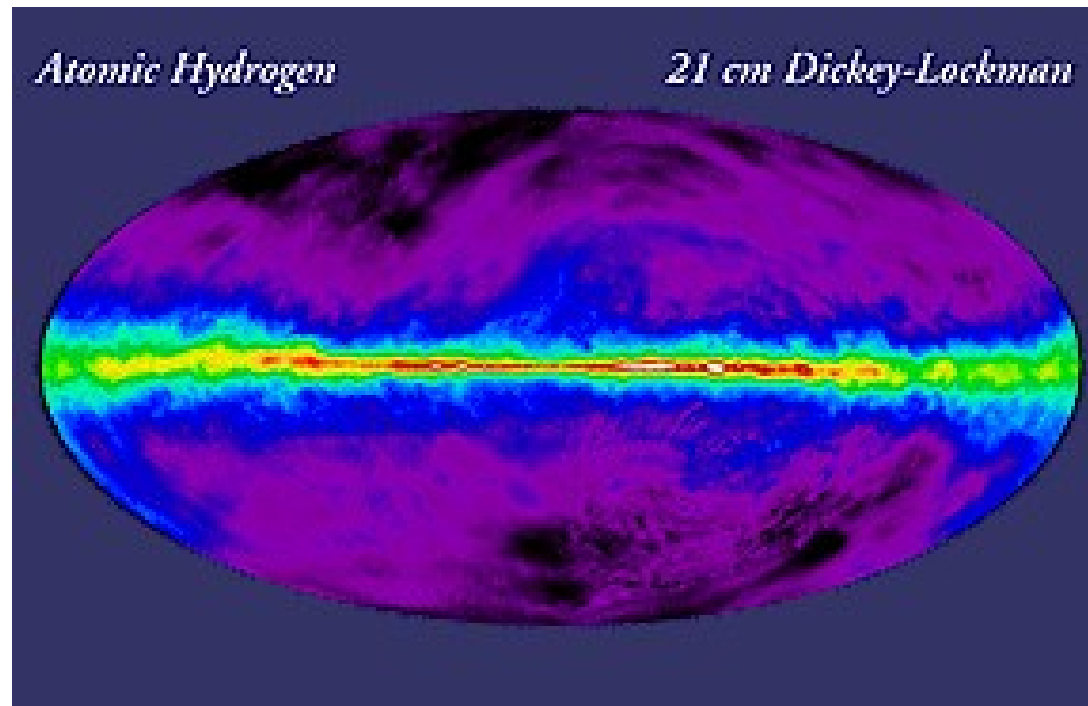
(0 towards Galactic Center;

90 in direction of Sun's motion
= clockwise from “above”)



HI in the Milky Way

**Nearly all the HI (neutral H) in the Galaxy is located in the disk.
Observed velocities governed by (a) rotation (b) random motions.**



(plotted in Galactic coordinates)

Gas dynamics: the Keplerian case

If an ensemble of gas clouds, each with mass m , is distributed in a disk orbiting a single massive object with $M \gg m$, then for each cloud we can write

$$F = mv^2/R = GMm/R^2$$

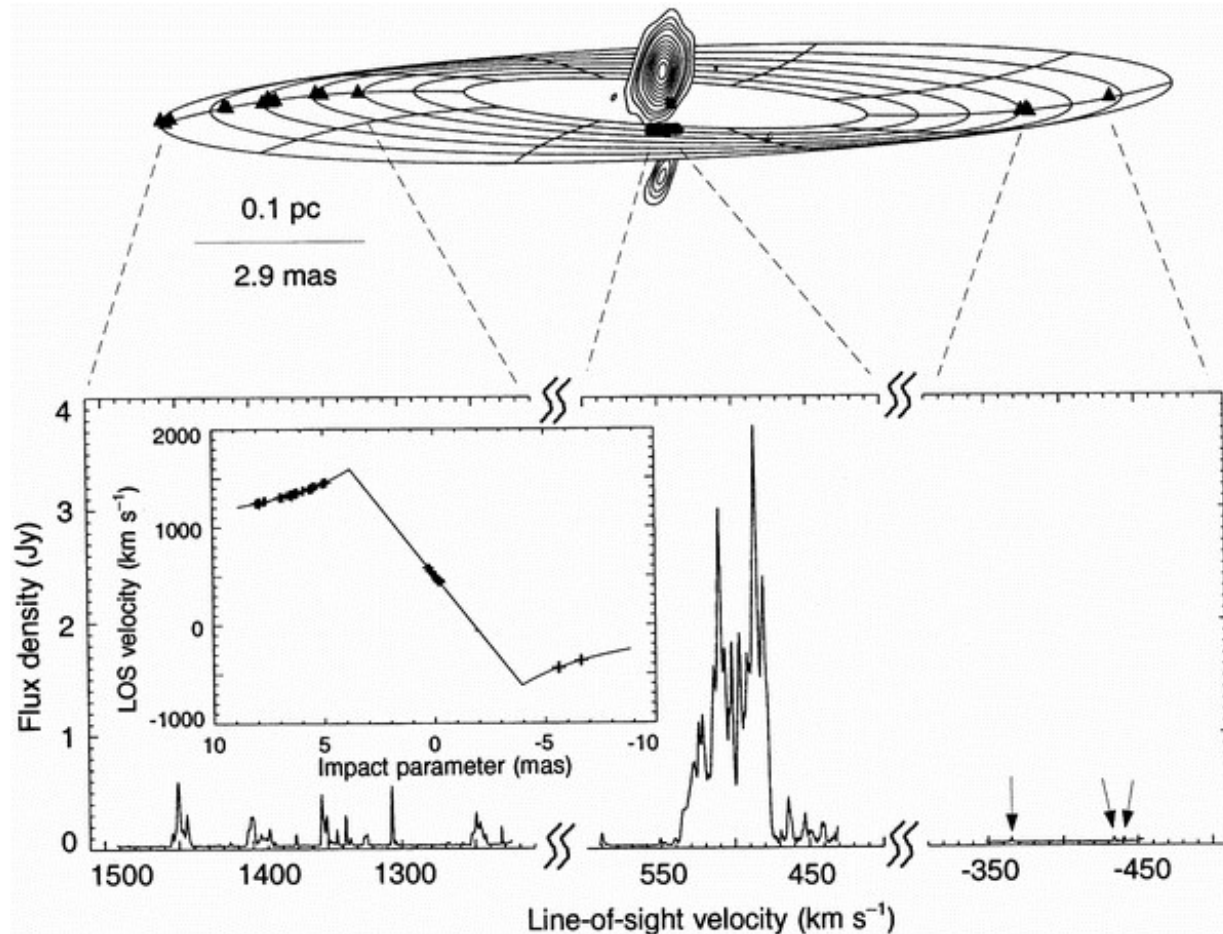
$$v^2R = GM$$

$$M = v^2R/G$$

which is equivalent to Kepler's third law for $v = 2\pi R/T$.

Gas dynamics: a Keplerian example

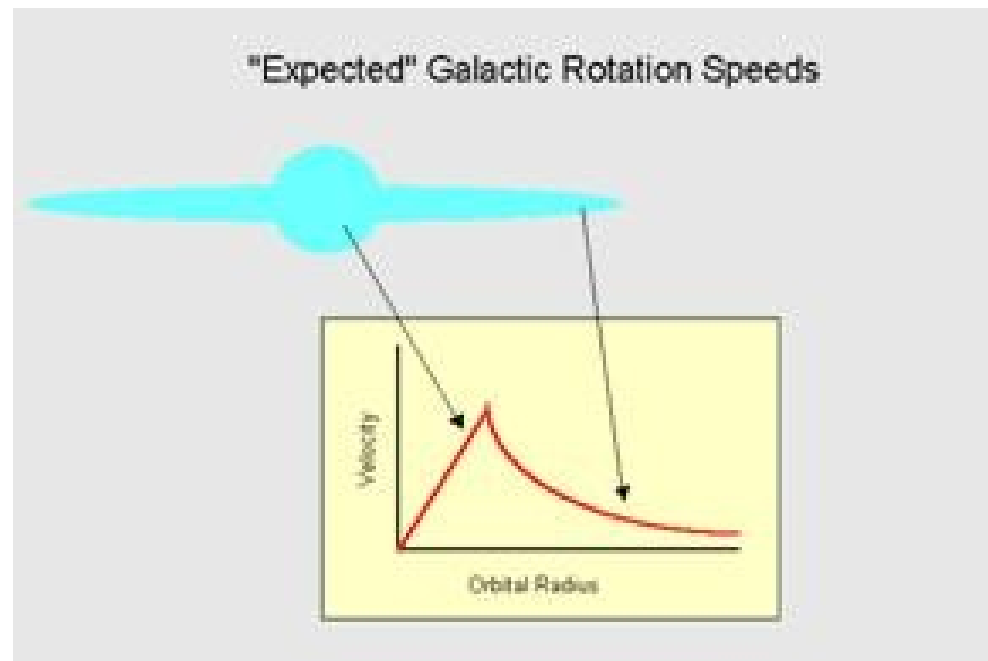
From Herrnstein et al. (1999): water masers tracing orbital motions around the central black hole in NGC4258.



Rotation curves in galaxies: expected

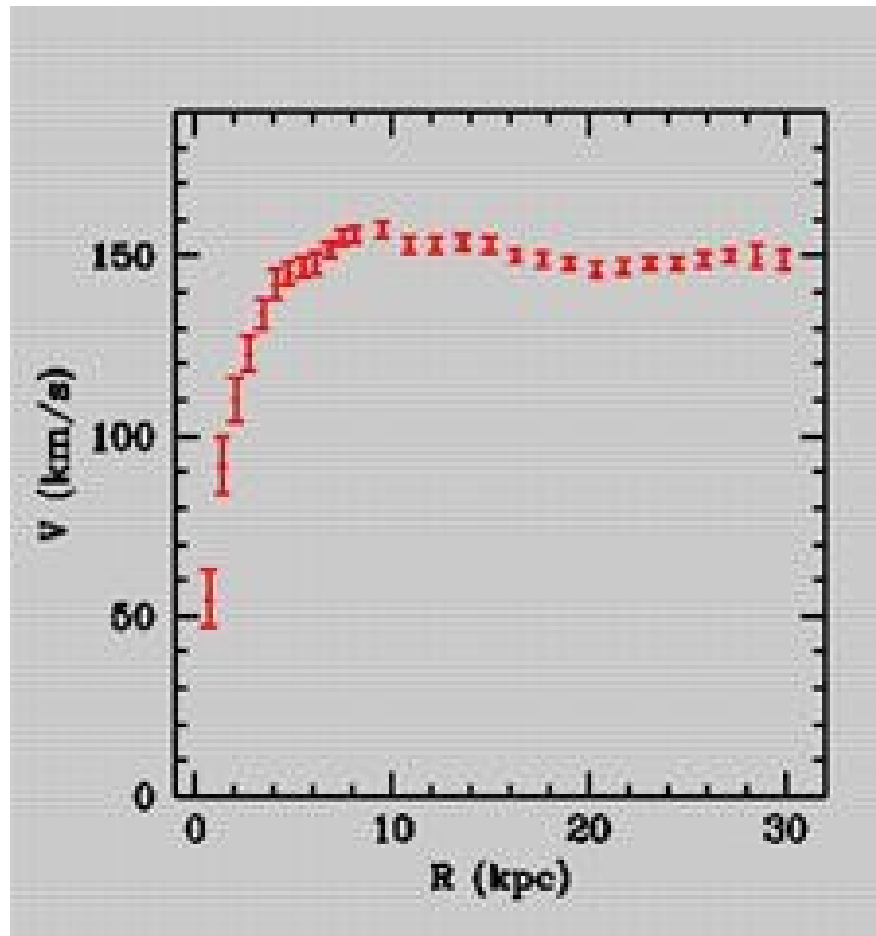
Rotation curves in galaxies are not Keplerian because we must replace M with the “interior mass” $M(<R)$ in the force equation... and $M(<R)$ is not constant as in the case of a central dominant mass.

What we expect, based on the central concentration of luminous matter (stars and gas):



Rotation curves in galaxies: observed

What we observe: **flat rotation curves**, implying the existence of additional non-luminous matter (i.e., **dark matter**).

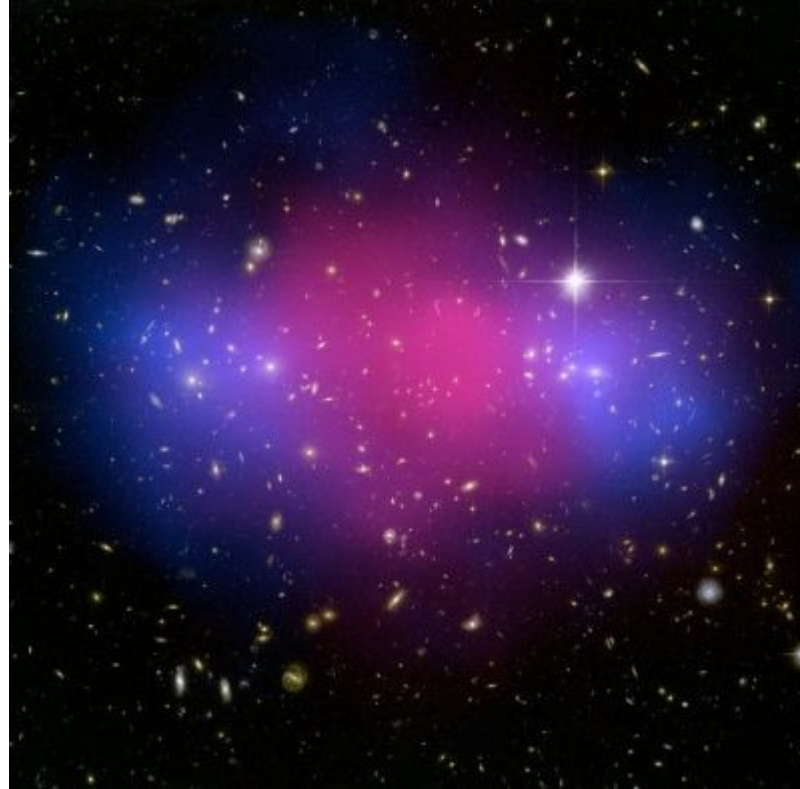


A heretical alternative?

A few bold souls have pointed out that once can just as easily relax the assumption of a universal law of gravitation as the assumption that all matter is luminous.

The idea that gravity might behave differently at low values of acceleration is known as **Modified Newtonian Dynamics (MOND). It works well in the context of spiral galaxy rotation curves, but not so well elsewhere.**

The “Bullet Cluster” vs. MOND



A collision between two clusters of galaxies: pink shows X-rays from hot gas, while blue shows mass based on weak lensing (apparently, mostly dark matter that is not “collisional”).

Inclination and rotation curves

If a galaxy is inclined relative to our line of sight, where

$i = 90^\circ$ means edge-on

$i = 0^\circ$ means face on

then the observed line of sight velocity is related to the intrinsic rotation velocity by $v_{\text{obs}} = v_{\text{rot}} \sin i$ if we make the assumption of intrinsic axisymmetry (i.e., roundness).

Quiz